Published for SISSA by 🖄 Springer

RECEIVED: September 16, 2016 REVISED: November 14, 2016 ACCEPTED: December 4, 2016 PUBLISHED: December 16, 2016

# One-loop effective actions and higher spins

# L. Bonora,<sup>*a,b*</sup> M. Cvitan,<sup>*c*</sup> P. Dominis Prester,<sup>*d*</sup> S. Giaccari,<sup>*c*</sup> B. Lima de Souza<sup>*a,b*</sup> and T. Štemberga<sup>*c*</sup>

<sup>b</sup>INFN, Sezione di Trieste,

Trieste, Italy

- <sup>c</sup> Theoretical Physics Division of Particles and Fields, Faculty of Science, University of Zagreb, Bijenička 32, HR-10000 Zagreb, Croatia
- <sup>d</sup>Department of Physics, University of Rijeka, Radmile Matejčić 2, 51000 Rijeka, Croatia

*E-mail:* bonora@sissa.it, mcvitan@phy.hr, pprester@phy.uniri.hr, sgiaccari@phy.hr, blima@sissa.it, tstember@phy.hr

ABSTRACT: The idea we advocate in this paper is that the one-loop effective action of a free (massive) field theory coupled to external sources (via conserved currents) contains complete information about the classical dynamics of such sources. We show many explicit examples of this fact for (scalar and fermion) free field theories in various dimensions d = 3, 4, 5, 6 coupled to (bosonic, completely symmetric) sources with a number of spins. In some cases we also provide compact formulas for any dimension. This paper is devoted to two-point correlators, so the one-loop effective action we construct contains only the quadratic terms and the relevant equations of motion for the sources we obtain are the linearized ones.

KEYWORDS: Higher Spin Gravity, Higher Spin Symmetry, Renormalization Group

ARXIV EPRINT: 1609.02088



<sup>&</sup>lt;sup>a</sup>International School for Advanced Studies (SISSA),

Via Bonomea 265, 34136 Trieste, Italy

## Contents

1	Inti	roduction. Higher spins are everywhere	2
2	Fre	e field theory models	3
	2.1	Generating functions and effective actions	6
3	An	7	
	3.1	Spin one and two — parity even sectors	7
	3.2	Linearized equations for spin 3 in parity even sector	8
4	Uni	iversal EOM and conserved structures for spin $s$	12
	4.1	Eom's from conserved structures	13
	4.2	Conserved odd parity structures	16
5	The	e general method	17
	5.1	Guidelines for the calculations	18
6	3 3d scalar effective field action tomography		
	6.1	3d msm: spin 1 current	20
	6.2	3d msm: e.m. tensor	21
	6.3	3d msm: spin 3 current	22
	6.4	3d msm: higher spin currents	24
7	<b>3</b> d :	fermion effective field theory action tomography	25
	7.1	3d mfm: spin 1 current	25
	7.2	3d mfm: e.m. tensor — even part	26
	7.3	3d mfm: e.m. tensor — odd part	27
	7.4	3d mfm: spin 3, even part	28
	7.5	3d mfm: spin 3, odd part	28
8	Ton	nography in 5d	29
	8.1	5d scalar model	29
		8.1.1 5d msm: spin 1 current	29
		8.1.2 5d msm: spin 2 current	30
		8.1.3 5d msm: spin 3 current	31
	8.2	5d fermion model	32
		8.2.1 5d mfm: spin 1 current	32
		8.2.2 5d mfm: e.m. tensor	32
		8.2.3 5d mfm: spin 3 current	33

9	Tom	nography in 4d	34
	9.1	4d scalar model	34
		9.1.1 4d msm: spin 1 current	34
		9.1.2 4d msm: spin 2 current	36
		9.1.3 4d msm: spin 3 current	37
	9.2	4d fermion model	38
		9.2.1 4d mfm: spin 1 current	38
		9.2.2 4d mfm: e.m. tensor	39
		9.2.3 4d mfm: spin 3 current	41
10	Tom	nography in 6d	<b>42</b>
	10.1	6d scalar model	42
		10.1.1 6d msm: spin 1 current	42
		10.1.2 6d msm: spin 2 current	43
		10.1.3 6d msm: spin 3 current	45
	10.2	6d fermion model	45
		10.2.1 6d mfm: spin 1 current	45
		10.2.2 6d mfm: spin 2 current	46
		10.2.3 6d mfm: spin 3 current	47
11	Spin	a $s$ current two-point correlators in any dimension	48
	11.1	Fermion model	48
		11.1.1 Fermion model — massive case	48
		11.1.2 d even	50
		11.1.3 d odd	54
		11.1.4 Spin 2	55
		11.1.5 Spin 3	55
	11.2	Scalar model	56
		11.2.1 Scalar model — massive case	56
		11.2.2 Concluding remark	57
12	Con	clusion	58
$\mathbf{A}$	Pro	of that a conserved structure can be written in terms of produc	ts of
	$\pi$ al	one	59
в	Mas	sless models	60
	B.1	Massless fermion model	60
	B.2	Massless scalar model	62
$\mathbf{C}$	4d f	ull amplitudes	64
	C.1	Spin 1	64
	C.2	Spin 2	65
	C.3	Spin 3	65

#### 1 Introduction. Higher spins are everywhere

The idea we wish to support in this paper is that the one-loop effective action of a free (massive) field theory coupled to external sources (via conserved currents) contains complete information about the possible classical dynamics of the sources. We exhibit several examples of this fact for (scalar and fermion) free field theories in various dimensions d = 3, 4, 5, 6 coupled to (bosonic) sources with a large number of spins. In some cases we also provide compact formulas for any dimension. In this paper we concentrate on two-point correlators, so the one-loop effective action we construct contains only the quadratic part. Consequently the equations of motion for the sources we obtain are the linearized ones. We postpone to a future work the analysis of one-point and three-point correlators. But our thesis is that the dynamics generated by the one-loop effective action (OLEA) contains all the information we need to reconstruct complete interacting equations of motion.

This paper is a follow-up of [1], which contains a few (mostly parity odd) 3d examples of what has just been said. The present paper is more general and systematic, not limited to 3d, and devoted especially to the parity even sector.

As we have just mentioned, the crucial issue here is the calculation of the two-point functions of free massive field theories coupled to external sources. We do it via Feynman diagrams. This is in principle a simple calculation, and to carry it out we resort to a method introduced by Davydychev and collaborators, [2-4]. However, as we will see, although we can derive from it very general formulas they are expressed in terms of hypergeometric functions and derivatives thereof and not easily 'readable'. For this reason it is often very useful to expand such results near their IR and UV fixed points. These expansions in powers of the mass m in odd dimensions, and m and  $\log m$  in even dimensions, allows us to single out the dynamics of the sources and will be referred to as *tomography*. In other words what we do is to describe RG trajectories of two-point current correlators that pass through those of free massive theories, but we focus in particular on their IR and UV expansions, where the physical content is more easily recognizable.

Such IR and UV expansions are necessary also for another reason: one has to check that the IR and UV limits of the one-loop effective action are well defined. We find in fact divergent and non-conserved term in the limit  $m \to \infty$ . These terms are local and can be subtracted. We subtract as well the IR finite terms, which are also local. This is to make the OLEA well defined and scheme independent. The results obtained in this way, in particular in the even parity sector, transferred to the OLEA, allow us to find the linearized Fronsdal eom's (see [5–9]) for all the source fields we have considered, in the nonlocal form introduced by Francia and Sagnotti, [10, 11]. In 3d we consider also the odd parity sector, and confirm the connection with Pope and Townsend's generalizations of Chern-Simons theory, already pointed out in [1].

In this paper, for the purpose of comparison, we also analyze massless free theories beside the corresponding massive ones. The difference between the two is that the latter allow us to control not only the UV but also the IR, while in the former only the UV is regularized. This explains the difference in the results. In general in the massless case we do not get all the information we can extract from the massive theory and many results are scheme dependent. Briefly stated, at least for the purpose of this paper, to make sure we get a complete information we must use massive models.

The subject of this paper is inspired by the idea of exploring theories with infinite many fields, [12, 13], in particular string theories and Vasiliev-type higher spin theories, [14–17]. As shown in the body of the paper higher spin fields appear naturally in the one-loop effective action of the simplest free theories in any dimension and it is possible to make contact with the literature on classical higher spin theories, [18–25]. Other sources of inspiration have been [26–32]. The idea of exploring the one-loop effective action is far from new: the list of works which may have some overlap with our paper includes [37–46], but is likely to be incomplete. From a technical point of view this paper continues the line of research started with [50–53] with more powerful techniques (a new Mathematica code).

The paper is organized as follows. In the next section we introduce the massive scalar and fermion model and define the relevant OLEA's. Section 3 is meant to explain the motivation for this research by means of simple concrete examples. We also introduce the issue of higher spin Fronsdal eom and their various forms. In section 4 we introduce a new representation of higher spin eom's in momentum space and their general form, which is independent on the dimension of space-time. Section 5 contains a short summary of Davydychev's method to compute one-loop Feynman diagrams. In section 6 to 10 we analyze the one-loop scalar and fermion model two-point functions and their IR and UV expansion (tomography) in 3, 5, 4 and 6 dimensions, respectively. In section 11 we produce the formulas for two-point correlators of spin 1, 2, 3 currents in any dimensions. Section 12 is devoted to the conclusion. Appendix A contains the demonstration of a result used in section 4 and appendix B the analysis of the massless scalar and fermion models.

#### 2 Free field theory models

In this paper we limit ourselves to two type of models, the free scalar and free fermion, although it is not hard to extend the analysis to other models. By the first we mean the complex scalar theory defined by the Lagrangian

$$L = \partial_{\mu}\phi^{\dagger}\partial^{\mu}\phi - m^{2}\phi^{\dagger}\phi \tag{2.1}$$

in any dimension. On shell the current

$$J_{\mu} = i \left( \phi^{\dagger} \partial_{\mu} \phi - \partial_{\mu} \phi^{\dagger} \phi \right)$$
(2.2)

is conserved. We can couple it to a gauge field via the action term  $\int d^d x A^{\mu}(x) J_{\mu}(x)$ . The scalar-scalar-gluon vertex with momenta p, p', k, respectively, (p incoming and p', k outgoing), and the propagator are, respectively,

$$i(p+p')_{\mu}\delta(p-p'-k), \qquad \frac{i}{p^2-m^2}$$
 (2.3)

But, of course we can define infinite many completely symmetric (on shell) conserved currents, of which (2.2) is only the simplest example:

$$J_{\mu_1\dots\mu_s} = i^s \phi^\dagger \stackrel{\leftrightarrow}{\partial_{\mu_1}} \dots \stackrel{\leftrightarrow}{\partial_{\mu_s}} \phi \tag{2.4}$$

They couple minimally to external spin s fields,  $a^{\mu_1...\mu_s}$ . The on-shell current conservation implies (to the lowest order) invariance under the gauge transformations

$$\delta a_{\mu_1\dots\mu_s} = \partial_{(\mu_1} \Lambda_{\mu_2\dots\mu_s)} \tag{2.5}$$

where round brackets stand for symmetrization.

In the case s = 2 the conserved current is the energy-momentum tensor and the external source is the metric fluctuation  $h_{\mu\nu}$ , where  $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$ . In this case the action is the integral of (2.1) multiplied by  $\sqrt{g}$ . The vertex for an incoming scalar with momentum p and outgoing scalar with momentum p' and an outgoing spin-s field with momentum k is

$$V_{sst}: \quad i(p+p')_{\mu_1}\dots(p+p')_{\mu_s}\delta^{(d)}(p-p'-k)$$
(2.6)

The free fermion model is represented by a Dirac fermion coupled to a gauge field. The action is

$$S[A] = \int d^3x \left[ i\bar{\psi}\gamma^{\mu}D_{\mu}\psi - m\bar{\psi}\psi \right], \quad D_{\mu} = \partial_{\mu} + A_{\mu}$$
(2.7)

where  $A_{\mu} = A^{a}_{\mu}(x)T^{a}$  and  $T^{a}$  are the generators of a gauge algebra in a given representation determined by  $\psi$ . We will use the antihermitean convention, so  $[T^{a}, T^{b}] = f^{abc}T^{c}$ , and the normalization  $\operatorname{tr}(T^{a}T^{b}) = -\delta^{ab}$ .

The current

$$J^a_\mu(x) = i\bar{\psi}\gamma_\mu T^a\psi \tag{2.8}$$

is (classically) covariantly conserved on shell as a consequence of the gauge invariance of (2.7)

$$(DJ)^a = (\partial^\mu \delta^{ac} + f^{abc} A^{b\mu}) J^c_\mu = 0$$
(2.9)

The next example involves the coupling to gravity

$$S[h] = \int d^3x \, e \left[ i\bar{\psi}E^{\mu}_a \gamma^a \nabla_{\mu}\psi - m\bar{\psi}\psi \right], \quad \nabla_{\mu} = \partial_{\mu} + \frac{1}{2}\omega_{\mu bc}\Sigma^{bc}, \quad \Sigma^{bc} = \frac{1}{4} \left[ \gamma^b, \gamma^c \right]. \tag{2.10}$$

The corresponding energy momentum tensor

$$T^{(g)}_{\mu\nu} = \frac{i}{4}\bar{\psi}\left(\gamma_{\mu}\overleftrightarrow{\nabla}_{\nu} + \gamma_{\nu}\overleftrightarrow{\nabla}_{\mu}\right)\psi.$$
(2.11)

is covariantly conserved on shell as a consequence of the diffeomorphism invariance of the action,

$$\nabla^{\mu} T_{\mu\nu}(x) = 0. \tag{2.12}$$

If we expand the metric around the flat spacetime,  $g_{\mu\nu}(x) = \eta_{\mu\nu} + h_{\mu\nu}(x)$ , then, contrary to spin-1 case, interaction is not linear in the gauge field, which is  $h_{\mu\nu}$ . However, for the

purposes of this paper, only linear term matters, and it is given by coupling the flat space energy-momentum tensor

$$T_{\mu\nu} = \frac{i}{4} \bar{\psi} \left( \gamma_{\mu} \overleftrightarrow{\partial}_{\nu} + \gamma_{\nu} \overleftrightarrow{\partial}_{\mu} \right) \psi.$$
(2.13)

to the metric fluctuation  $h_{\mu\nu}$ .

Again we can couple the fermions to more general fields. Consider the free action

$$S_0 = \int d^3x \, \left[ i\bar{\psi}\gamma^\mu \partial_\mu \psi - m\bar{\psi}\psi \right], \qquad (2.14)$$

and the spin three conserved current

$$J_{\mu_{1}\mu_{2}\mu_{3}} = -\frac{1}{2}\bar{\psi}\gamma_{(\mu_{1}}\partial_{\mu_{2}}\partial_{\mu_{3})}\psi - \frac{1}{2}\partial_{(\mu_{1}}\partial_{\mu_{2}}\bar{\psi}\gamma_{\mu_{3})}\psi + \frac{5}{3}\partial_{(\mu_{1}}\bar{\psi}\gamma_{\mu_{2}}\partial_{\mu_{3})}\psi - \frac{1}{3}\eta_{(\mu_{1}\mu_{2}}\partial^{\sigma}\bar{\psi}\gamma_{\mu_{3}})\partial_{\sigma}\psi + \frac{m^{2}}{3}\eta_{(\mu_{1}\mu_{2}}\bar{\psi}\gamma_{\mu_{3}})\psi$$
(2.15)

Using the equation of motion one can prove that

$$\partial^{\mu}J_{\mu\nu\lambda} = 0 \tag{2.16}$$

$$J_{\mu}{}^{\mu}{}_{\lambda} = -\frac{4}{9}m\left(-i\partial_{\lambda}\bar{\psi}\psi + i\bar{\psi}\partial_{\lambda}\psi + 2\bar{\psi}\gamma_{\lambda}\psi\right)$$
(2.17)

Therefore, the spin three current (2.15) is conserved on shell and its tracelessness is softly broken by the mass term. Similarly to the gauge field and the metric, we can couple the fermion  $\psi$  to a new external source  $b_{\mu\nu\lambda}$  by adding to (2.14) the term

$$\int d^3x J_{\mu\nu\lambda} b^{\mu\nu\lambda} \tag{2.18}$$

Due to the (on shell) current conservation this coupling is invariant (to lowest order) under the infinitesimal gauge transformations

$$\delta b_{\mu\nu\lambda} = \partial_{(\mu}\Lambda_{\nu\lambda)} \tag{2.19}$$

In the limit  $m \to 0$  we have also invariance under the local transformations

$$\delta b_{\mu\nu\lambda} = \Lambda_{(\mu}\eta_{\nu\lambda)} \tag{2.20}$$

which are usually referred to as (generalized) Weyl transformations and which induce the tracelessness of  $J_{\mu\nu\lambda}$  in any couple of indices.

We can construct on-shell conserved currents for any spin s, but their form is more complicated than in the scalar case. The explicit expressions can be found in [1].

We notice that to lowest order in the external sources the relevant action, in all cases above, takes the form of the free action + a linear interaction term such as (2.18). We make the identification  $a_{\mu} = A_{\mu}, a_{\mu\nu} \sim h_{\mu\nu}, a_{\mu\nu\lambda} \sim b_{\mu\nu\lambda}$ , with the obvious exception of the non-Abelian field in (2.7). The latter will be the only case in which we consider non-Abelian external sources.<sup>1</sup>

<sup>&</sup>lt;sup>1</sup>Also note that the nonlinearity present in spin-2 case, which is forced by the consistency requirements, is a signal that we should expect the same for higher-spin fields. However, this is not relevant in our two-point calculations.

#### 2.1 Generating functions and effective actions

In both scalar and fermion cases, the generating function for the external source  $a_{\mu_1...\mu_s}$  is

$$W[a,s] = W[0] + \sum_{n=1}^{\infty} \frac{i^{n-1}}{n!} \int \prod_{i=1}^{n} d^{3}x_{i} a^{\mu_{11}\dots\mu_{1s}}(x_{1})\dots a^{\mu_{n1}\dots\mu_{ns}}(x_{n}) \\ \times \langle 0|\mathcal{T}J^{(s)}_{\mu_{11}\dots\mu_{1s}}(x_{1})\dots J^{(s)}_{\mu_{n1}\dots\mu_{ns}}(x_{n})|0\rangle.$$
(2.21)

In particular  $a_{\mu} = A_{\mu}, a_{\mu\nu} = \frac{1}{4}h_{\mu\nu}$  and  $J^{(2)}_{\mu\nu} = 2T_{\mu\nu}$  with  $a_{\mu\nu\lambda} = b_{\mu\nu\lambda}$ . The full one-loop 1-pt correlator for  $J_{\mu_1...\mu_s}$  is

$$\langle\!\langle J_{\mu_1\dots\mu_s}^{(s)}(x)\rangle\!\rangle = \frac{\delta W[a,s]}{\delta a^{\mu_1\dots\mu_s}(x)} = \sum_{n=0}^{\infty} \frac{i^n}{n!} \int \prod_{i=1}^n d^3 x_i a^{\mu_{11}\dots\mu_{1s}}(x_1)\dots a^{\mu_{n1}\dots\mu_{ns}}(x_n) \\ \times \langle 0|\mathcal{T}J_{\mu_1\dots\mu_s}^{(s)}(x)J_{\mu_{11}\dots\mu_{1s}}^{(s)}(x_1)\dots J_{\mu_{n1}\dots\mu_{ns}}^{(s)}(x_n)|0\rangle.$$
(2.22)

The full one-loop conservation law for the energy-momentum tensor is

$$\nabla^{\mu} \langle\!\langle T_{\mu\nu}(x) \rangle\!\rangle = 0. \tag{2.23}$$

A similar covariant conservation should be written also for the other currents, but for s > 2we will content ourselves with the lowest nontrivial order in which the conservation law reduces to

$$\partial^{\mu_1} \langle\!\langle J^{(s)}_{\mu_1 \dots \mu_s}(x) \rangle\!\rangle = 0 \tag{2.24}$$

**Warning.** One must be careful when applying the previous formulas for generating functions. If the expression  $\langle 0|\mathcal{T}J_{\mu_{11}\ldots\mu_{1s}}^{(s)}(x_1)\cdots J_{\mu_{n1}\ldots\mu_{ns}}^{(s)}(x_n)|0\rangle$  in (2.21) is meant to denote the *n*-th point-function calculated by using Feynman diagrams, a factor  $i^n$  is already included in the diagram themselves and so it should be dropped in (2.21). When the current is the energy-momentum tensor an additional precaution is necessary: the factor  $\frac{i^{n-1}}{n!}$  must be replaced by  $\frac{i^{n-1}}{2^n n!}$ . The factor  $\frac{1}{2^n}$  is motivated by the fact that when we expand the action

$$S[\eta+h] = S[\eta] + \int d^d x \frac{\delta S}{\delta g^{\mu\nu}} \Big|_{g=\eta} h^{\mu\nu} + \cdots,$$

the factor  $\frac{\delta S}{\delta g^{\mu\nu}}\Big|_{g=\eta} = \frac{1}{2}T_{\mu\nu}$ . Another consequence of this fact will be that the presence of vertices with one graviton in Feynman diagrams will correspond to insertions of the operator  $\frac{1}{2}T_{\mu\nu}$  in correlation functions.

Our purpose in this paper is to compute the effective action for the external source fields at the quadratic order. As a consequence the first task is to compute the two-point functions

$$\langle 0|\mathcal{T}J^{(s)}_{\mu_1...\mu_s}(x) J^{(s)}_{\nu_1...\nu_s}(y)|0\rangle$$
 (2.25)

or their Fourier transforms

$$\tilde{J}_{\mu_1...\mu_s\nu_1...\nu_s}(k) = \langle 0|\mathcal{T}\tilde{J}^{(s)}_{\mu_1...\mu_s}(k)\,\tilde{J}^{(s)}_{\nu_1...\nu_s}(-k)|0\rangle$$
(2.26)

In the sequel we compute them by using the Feynman diagram technique. For all twopoint functions the only relevant diagram is the bubble diagram with one spin s line of ingoing momentum k and one with the same outgoing momentum and one scalar or fermion circulating in the internal loop. For instance the 2pt function for the current  $J^a$  in the fermion model is

$$\tilde{J}^{ab}_{\mu\nu}(k) = \langle \tilde{J}^{a}_{\mu}(k)\tilde{J}^{b}_{\nu}(-k)\rangle = \int \frac{d^{3}p}{(2\pi)^{3}} \operatorname{Tr}\left(\gamma_{\nu}T^{b}\frac{1}{\not{p}-m}\gamma_{\mu}T^{a}\frac{1}{\not{p}-\not{k}-m}\right)$$
(2.27)

while for the e.m. tensor it is

$$\tilde{T}_{\mu\nu\lambda\rho}(k) = \frac{1}{4} \langle \tilde{T}_{\mu\nu}(k)\tilde{T}_{\lambda\rho}(-k) \rangle$$

$$= -\frac{1}{64} \int \frac{d^3p}{(2\pi)^3} \operatorname{Tr}\left(\frac{1}{\not p - m}(2p - k)_{\mu}\gamma_{\nu}\frac{1}{\not p - \not k - m}(2p - k)_{\lambda}\gamma_{\rho}\right),$$

$$(2.28)$$

where symmetrization of indices  $(\mu, \nu)$  and  $(\lambda, \rho)$  and the factor  $\frac{1}{4}$  is introduced accordingly to the above warning.

#### 3 An appetizer in 3d

In [1] we calculated in particular the two-point function of the current  $J^a$  in the fermion model as well as its IR and UV limit. In the parity violating part we found a wellknown result: when Fourier antitransformed and inserted in the generating function of the OLEA (2.21) it gives rise to the linearized version of the gauge CS action in 3d (which is in fact conformal invariant). In [1] we did the same for the two-point correlator of the e.m. tensor for the fermion model, and proceeding the same way we found the linearized version of the gravity CS action. Something that was also known before, [38]. Repeating the same thing for the spin 3 current above we found instead a previously unknown result: the UV limit in particular leads to a linearized action that corresponds to a spin 3 CS generalization postulated long ago by Pope and Townsend, [33–36].

These were the results found in the parity odd part (in [1] we were mostly interested in the latter). But the even parity parts of the two-point correlators have perhaps even more interesting interpretations, so let us briefly analyze the parity even parts of the linearized effective actions obtained from 2-point current correlators in the free massive Dirac fermion quantum field theory in 3d in [1].

#### 3.1 Spin one and two — parity even sectors

The UV limit of the two-point function of the  $J^a$  currents are nonlocal conformal correlators, according to expectations, see [30]. The same is true for the e.m. tensor two-point function. But now let us focus on the IR limits. According to [1], for the  $J^a$  current two-point function, for large m we have

$$\tilde{J}^{ab(even)}_{\mu\nu}(k) = \frac{i}{4\pi} \frac{1}{3|m|} \delta^{ab}(k_{\mu}k_{\nu} - k^2\eta_{\mu\nu})$$
(3.1)

This term is local. Fourier anti-transforming it and inserting it into (2.21) it gives rise to the action

$$S \sim \frac{1}{m} \int d^3x \, \left( A^a_\mu \partial^\mu \partial^\nu A^a_\nu - A^a_\nu \Box A^{a\nu} \right) \tag{3.2}$$

which is the lowest term in the expansion of the YM action

$$S_{\rm YM} = -\frac{1}{g_{\rm YM}} \int d^3x \,\mathrm{Tr} \left(F_{\mu\nu}F^{\mu\nu}\right) \tag{3.3}$$

where  $g_{\rm YM} \sim |m|$ .

Now let us go to the IR limit of the even part of the 2pt e.m. tensor correlator. eq. (3.36) of [1] says

$$\langle T_{\mu\nu}(k)T_{\lambda\rho}(-k)\rangle_{\text{even}}^{\text{IR}} = \frac{i|m|}{96\pi} \left[ \frac{1}{2} \left( (k_{\mu}k_{\lambda}\eta_{\nu\rho} + \lambda \leftrightarrow \rho) + \mu \leftrightarrow \nu \right) - (k_{\mu}k_{\nu}\eta_{\lambda\rho} + k_{\lambda}k_{\rho}\eta_{\mu\nu}) - \frac{k^{2}}{2} \left( \eta_{\mu\lambda}\eta_{\nu\rho} + \eta_{\mu\rho}\eta_{\nu\lambda} \right) + k^{2}\eta_{\mu\nu}\eta_{\lambda\rho} \right].$$
(3.4)

This is a local expression multiplied by |m|. In fact Fourier anti-transforming it and inserting it into (2.19) it gives rise to the action

$$S \sim |m| \int d^3x \left( -2\partial_\mu h^{\mu\lambda} \partial_\nu h^\nu_\lambda - 2h \,\partial_\mu \partial_\nu h^{\mu\nu} - h^{\mu\nu} \Box h_{\mu\nu} + h \Box h \right)$$
(3.5)

which is the linearized Einstein-Hilbert action:

$$S_{EH} = \frac{1}{2\kappa} \int d^3x \sqrt{g} R \tag{3.6}$$

where  $\kappa \sim \frac{1}{|m|}$ .

These results for spin-1 and -2 are known have been known for a long time, see for instance [26]. Now, we ask the same question for the 2pt correlator of the 3-current (section 3.3). What action, if any, does it represent for the external source field?

#### 3.2 Linearized equations for spin 3 in parity even sector

Before presenting our results in 3d, let us briefly review the status of the linearized equations for the massless spin 3 field described by the completely symmetric field  $\varphi_{\mu\nu\lambda}$ . Historically the first formulation of equations for the *unconstrained* free massless spin 3 field was given by Fronsdal [5, 6]

$$\mathcal{F}_{\mu\nu\lambda} \equiv \Box \varphi_{\mu\nu\lambda} - \partial_{\underline{\mu}} \partial \cdot \varphi_{\underline{\nu\lambda}} + \partial_{\underline{\mu}} \partial_{\underline{\nu}} \varphi_{\underline{\lambda}}' = 0 \tag{3.7}$$

where underlined indices mean the sum over the minimum number of terms necessary to completely symmetrize the expression in  $\mu, \nu$  and  $\lambda$ , i.e. for instance

$$\partial_{\mu}\partial_{\cdot}\varphi_{\underline{\nu}\underline{\lambda}} = \partial_{\mu}\partial_{\alpha}\varphi^{\alpha}{}_{\nu\lambda} + \partial_{\nu}\partial_{\alpha}\varphi^{\alpha}{}_{\lambda\mu} + \partial_{\lambda}\partial_{\alpha}\varphi^{\alpha}{}_{\mu\nu} ,$$

and where a prime ' means that the tensor is traced over a pair of indices. In some formulas we shall use shorter notation in which all indexes are suppressed.

Under the gauge variation (2.19),  $\delta \varphi_{\mu\nu\lambda} = \partial_{\mu}\Lambda_{\nu\lambda} + \text{perm.}$ , the Fronsdal kinetic tensor transforms as  $\delta \mathcal{F}_{\mu\nu\lambda} = 3\partial_{\mu}\partial_{\nu}\partial_{\lambda}\Lambda'$ . It follows that the Fronsdal equation is invariant only on *restricted* gauge transformations satisfying  $\Lambda' = 0$  (this requirement holds for all higher spins). Also, the Fronsdal tensor is not divergence-free,  $\partial \cdot \mathcal{F} \neq 0$ , so one cannot directly couple the spin 3 field to a conserved (i.e., divergence-free) current using the Fronsdal equation. As we construct effective actions and corresponding equations for the higher spin fields by (minimally) coupling to conserved currents, it is obvious that Fronsdal's formalism is not suited for our purposes.

The formulation appropriate for our purposes was proposed in [10, 11], and analyzed in more detail in [47] (for a review, see [48]). It was shown that there is a one parameter class of equations for unconstrained spin 3 field, which are order 2 in derivatives, fully gauge invariant, and ready to be coupled to the external conserved current. These equations are most elegantly expressed by using gauge invariant linearized spin 3 Riemann tensor defined by

$$R_{\mu_1\nu_1\mu_2\nu_2\mu_3\nu_3} = \partial_{\mu_1}\partial_{\mu_2}\partial_{\mu_3}\varphi_{\nu_1\nu_2\nu_3} \quad (\text{antisymmetrised in all } (\mu_j,\nu_j)) \tag{3.8}$$

The spin 3 equations are parametrized by real number a and given by

$$\mathcal{G}(a)_{\mu\nu\lambda} \equiv \mathcal{A}(a)_{\mu\nu\lambda} - \eta_{\underline{\lambda}\underline{\nu}} \,\mathcal{A}(a)'_{\underline{\mu}} = 0 \tag{3.9}$$

$$\mathcal{A}(a)_{\mu\nu\lambda} \equiv \frac{1}{\Box} \partial \cdot R'_{\underline{\mu}\underline{\nu}\underline{\lambda}} + a \frac{\partial_{\underline{\nu}}\partial_{\underline{\lambda}}}{\Box^2} \partial \cdot R''_{\underline{\mu}}$$
(3.10)

where spin 3 Ricci tensors are defined by

$$R'_{\mu\nu\rho\sigma} \equiv \eta^{\alpha\beta} R_{\mu\nu\rho\alpha\sigma\beta} = 2\partial_{[\mu}\mathcal{F}_{\nu]\rho\sigma}$$
$$R''_{\mu\nu} \equiv \eta^{\rho\sigma} R'_{\mu\nu\rho\sigma} = 2\partial_{[\mu}\mathcal{F}'_{\nu]}$$
(3.11)

while their divergences are defined  $by^2$ 

$$\partial \cdot R'_{\mu\nu\lambda} = \partial_{\alpha} R^{\prime \alpha}{}_{\mu\nu\lambda}, \qquad \qquad \partial \cdot R^{\prime \prime}_{\mu} = \partial_{\alpha} R^{\prime \prime \alpha}{}_{\mu} \qquad (3.12)$$

What is the difference between equations with different a? First of all, it can be shown that regardless the value of a, the free field equation (3.9)–(3.10) is equivalent to Fronsdal equation (3.7). They start to differ when interactions are introduced. Note that equations (for any a) are non-local. From the purely mathematical side, the equation for a = 0 plays a special role because it is the least singular on-shell,<sup>3</sup> and because of this it was originally promoted in [10, 11]. However, it was later shown in [47] that equations with different parameters a propagate different set of excitations when coupled to a conserved external current  $J_{\mu\nu\lambda}$ ,

$$\mathcal{G}(a) = J, \qquad \partial \cdot J = 0 \qquad (3.13)$$

<sup>&</sup>lt;sup>2</sup>The Riemann tensor symmetries guarantee that the definitions for Ricci's and corresponding divergences (after symmetrization is taken into account) are essentially unique, in the sense that different choices for contracting indexes can differ only by a sign, or are vanishing [36].

<sup>&</sup>lt;sup>3</sup>In momentum space the on-shell condition is  $k^2 = 0$ .

In particular, it was shown that only equation with a = 1/2 propagates *spin 3 massless* excitations and nothing else, if one does not introduce additional constraints on  $\varphi$  or J. For a = 1/2 the tensor  $\mathcal{A}$  can be also written as

$$\mathcal{A}(1/2) = \mathcal{F} - \frac{\partial^3}{\Box^2} \partial \cdot \mathcal{F}'$$
(3.14)

Let us emphasize that this by itself does not mean that the equation with a = 1/2 is the "right one" to be used for the consistent coupling to the dynamical matter.

The non-locality of equations (3.9)-(3.9) can be 'cured' by multiplying with  $\Box^r$  with r large enough. It is obvious that the equation with a = 0 is special in that r = 1 already does the job, while for  $a \neq 0$  one needs r = 2. In this way one cures non-locality, but the price paid is that equations become higher-derivative (order 4 for a = 0 and order 6 for  $a \neq 0$ ). This opens up an additional question when one considers coupling to the conserved current J: should we do this as in (3.13), or should we couple the current in the local way,

$$\Box^r \mathcal{G}(a) = J, \qquad \partial \cdot J = 0 \qquad (3.15)$$

with r large enough?

The moral of the above analysis is that, due to several reasons, there is a large degeneracy in formulating equations of motion for the free massless spin 3 field, and it is not obvious that all formulations can be used as a basis for constructing consistent interacting quantized theories. It would be advantageous to know which formulation(s) are more promising, before embarking into such enterprise. We shall now argue that the induced action method may give us a hint.

In section 3.2.4 of [1] it was shown that the parity even part of the spin 3 two-point current correlator for a massive Dirac fermion in 3d is given by

$$\widetilde{\mathcal{J}}_{\mu_1\mu_2\mu_3\nu_1\nu_2\nu_3}^{(even)}(k) = \tau_b \left(\frac{k^2}{m^2}\right) |k|^5 \pi_{\mu_1\mu_2}^{(k)} \pi_{\mu_3\nu_1}^{(k)} \pi_{\nu_2\nu_3}^{(k)} + \tau_b' \left(\frac{k^2}{m^2}\right) |k|^5 \pi_{\mu_1\nu_1}^{(k)} \pi_{\mu_2\nu_2}^{(k)} \pi_{\mu_3\nu_3}^{(k)}$$
(3.16)

where  $\tau_b$  and  $\tau'_b$  are form factors presented in [1], and

$$\pi_{\mu\nu}^{(k)} = \eta_{\mu\nu} - \frac{k_{\mu}k_{\nu}}{k^2} \tag{3.17}$$

are projectors which guarantee conservation. From (2.22) it follows that the linearized effective equation in momentum space for the background spin 3 field minimally coupled to a conserved current in free QFT with massive Dirac field in 3d, is given by

$$\widetilde{\mathcal{J}}_{\mu_1\mu_2\mu_3\nu_1\nu_2\nu_3}(k)\,\widetilde{\varphi}^{\nu_1\nu_2\nu_3}(k) = \langle\!\langle \widetilde{J}^{(3)}_{\mu_1\mu_2\mu_3}(k)\rangle\!\rangle\,,\qquad k\cdot\widetilde{J}^{(3)}(k) = 0 \tag{3.18}$$

The form factors contain branch-cuts, which means that this equation is strongly nonlocal. The fact that there are two independent conserved structures present in (3.16), and so in (3.18), is directly connected with the one-parameter degeneracy introduced in (3.10).

In the IR region  $(|k^2|/m^2 < 4)$  the form factors are analytic, as expected, and the equation is weakly nonlocal (infinite sum of local terms) when expanded around |k|/m = 0.

Using the expansions of form factors from [1], we obtain that the leading term in the IR is given by

$$\widetilde{\mathcal{J}}_{\mu_1\mu_2\mu_3\nu_1\nu_2\nu_3}^{(even)}(k) \sim |m| \, k^4 \left( \pi_{\mu_1\mu_2}^{(k)} \pi_{\mu_3\nu_1}^{(k)} \pi_{\nu_2\nu_3}^{(k)} - \pi_{\mu_1\nu_1}^{(k)} \pi_{\mu_2\nu_2}^{(k)} \pi_{\mu_3\nu_3}^{(k)} \right) \tag{3.19}$$

Observe that this is the lowest derivative conserved local expression, which is unique. Now, plugging (3.19) into (3.18) and Fourier antitransforming, we obtain for the linearized induced equation in the coordinate space

$$|m| G_{\mu\nu\rho}(x) \sim \langle\!\langle J^{(3)}_{\mu_1\mu_2\mu_3}(x) \rangle\!\rangle, \qquad \partial \cdot J^{(3)} = 0$$
 (3.20)

where G is the conserved symmetric local tensor linear in  $\varphi$ , which is 4th-order in derivatives. As there is a unique such tensor, we can conclude (without doing any calculations) that it must be proportional to  $\Box \mathcal{G}(0)$ , with  $\mathcal{G}(0)$  defined in (3.9)–(3.10). Explicitly written,

$$G_{\mu\nu\lambda} = \partial_{\alpha} F^{\alpha}{}_{(\mu\nu\lambda)} \tag{3.21}$$

where

$$F_{\alpha\mu\nu\lambda} \equiv R'_{\alpha\mu\nu\lambda} - \frac{1}{2}R''_{\alpha\mu}\eta_{\nu\lambda} = 2\partial_{[\alpha}\left(\mathcal{F}_{\mu]\nu\lambda} - \frac{1}{2}\mathcal{F}'_{\mu]}\eta_{\nu\lambda}\right)$$
(3.22)

The result (3.20)–(3.22) is, in some sense, natural. First of all, it is the lowest derivative linear local parity invariant equation satisfying unrestricted gauge invariance and conservation condition. Also, the equation is of the same form as in spin 1 case, and we can identify the tensor F as spin 3 Maxwell tensor, while G appears to be spin 3 Riemann tensor (it is the lowest derivative local conserved gauge invariant parity even rank-3 tensor).<sup>4</sup>

Let us connect our result with the known constructions, reviewed above. It is obvious that our result (3.20)–(3.22) is the same as (3.15) with a = 0 and r = 1, i.e., we have obtained a local version of the equation proposed in [10, 11]. As we already mentioned, this equation does not propagate only spin 3 massless excitations, unless the conserved spin 3 current of the Dirac theory has some special properties which takes care of the redundant modes. This is the question we plan to investigate in the future.

Let us now briefly comment the UV limit  $(m/|k| \rightarrow 0)$ . After subtracting IR divergent terms (for a full explanation of this issue, see below) form factors in the UV limit tend to constants, which gives rise to a non-local correlator. However one of the subleading terms gives a combination of the two conserved quantities

$$A : k^{2} \pi_{\mu_{1}\nu_{1}} \pi_{\mu_{2}\nu_{2}} \pi_{\mu_{3}\nu_{3}}$$
  
$$B : k^{2} \pi_{\mu_{1}\mu_{2}} \pi_{\mu_{3}\nu_{1}} \pi_{\nu_{2}\nu_{3}}$$
(3.23)

which is not the same combination as the one present in IR limit (3.19). So, the corresponding induced linearized equation is also different.

A priori, one could freely linearly combine terms A and B and construct one parameter candidate equations for the free spin 3 field. For example, A by itself gives the following equation

$$\Box \varphi_{\mu\nu\lambda} - \partial_{\underline{\mu}} \partial \cdot \varphi_{\underline{\nu}\underline{\lambda}} + \frac{1}{\Box} \partial_{\underline{\mu}} \partial_{\underline{\nu}} \partial \cdot \partial \cdot \varphi_{\underline{\lambda}} - \frac{1}{\Box^2} \partial_{\mu} \partial_{\nu} \partial_{\lambda} \partial \cdot \partial \cdot \partial \cdot \varphi = 0$$
(3.24)

 $<sup>^{4}</sup>$ Conventions for naming objects in higher-spin metric-like formalism is notorious for its inconsistency. In the literature different objects are called Ricci tensor and Riemann tensor. We believe that our conventions are natural generalizations of spin 1 and 2 cases.

By combining with the traced equation, it can be shown that it is equivalent to the Fronsdal equation. The same can be shown for generic linear combination of A and B. There is though the special case, the combination 4B - 3A, which is traceless, for which the equation is

$$\Box\varphi_{\mu\nu\lambda} - 3\partial_{\mu}\partial_{\nu}\varphi_{\nu\lambda} + \frac{3}{4}\partial_{\mu}\partial_{\nu}\varphi_{\lambda}' - \frac{3}{4}\frac{1}{\Box}\partial_{\mu}\partial_{\nu}\partial_{\lambda}\partial_{\nu}\varphi' - \frac{1}{4}\frac{1}{\Box^{2}}\partial_{\mu}\partial_{\nu}\partial_{\lambda}\partial_{\nu}\partial_{\nu}\partial_{\nu}\varphi + \frac{9}{4}\frac{1}{\Box}\partial_{\mu}\partial_{\nu}\partial_{\nu}\partial_{\nu}\varphi_{\lambda} - \frac{3}{4}\eta_{\mu\nu}\Box\varphi_{\lambda}' + \frac{3}{4}\eta_{\mu\nu}\partial_{\lambda}\partial_{\nu}\varphi' + \frac{3}{4}\eta_{\mu\nu}\partial_{\nu}\partial_{\nu}\varphi_{\lambda} - \frac{3}{4}\eta_{\mu\nu}\frac{1}{\Box^{2}}\partial_{\lambda}\partial_{\nu}\partial_{\nu}\varphi = 0$$
(3.25)

In conclusion, we see that our simple analysis, based solely on the classification of possible conserved structures, recovers the Francia-Sagnotti analysis and gives an efficient method for analyzing higher spin actions. But, we emphasize that the induced action method, out of many possibilities, picks particular equations which are already coupled to particular external currents.

**Comment.** The previous results are limited to 3d and to the lowest spins. They are nevertheless enough to stir our interest and motivate a more in depth analysis. It is also clear enough that equations in the coordinate space are not always the best fit to generalizations to higher spins. Writing down the actions and equations of motion in the explicit form used so far becomes rapidly unwieldy with increasing spins and dimensions. Fortunately a language much sleeker than this and the formalism used so far in higher spin theories is at hand. We simply must go to momentum space and use the projector (3.17). Before plunging into the analysis of the results for 2pt correlators coming from Feynman diagrams, we'd better prepare the ground with a general analysis of their expected structure.

#### 4 Universal EOM and conserved structures for spin s

Our starting point is the 2-pt functions of symmetric conserved currents. We expect them to be conserved too, i.e. we expect to find 0 if we contract any index with the external momentum k. We exclude the presence of anomalies. In fact we will come across also some non-conservations, but they can be fixed by subtracting local counterterms. This aspect of our analysis is interesting in itself, but we will illustrate it later on in any detail. For the time being we ignore this fact and suppose that all 2-pt functions we deal with are conserved.

This said, the form of the conserved structures is universal, in the sense that is does not depend on the dimension d of spacetime. For spin s they can be easily constructed by means of the projector (3.17) and polarization vectors  $n_1, n_2$ :  $n_{1\mu}, n_{2\nu}$ .

For spin s let us write down the structures:

$$\tilde{A}_0^{(s)}(k \cdot n_1 \cdot n_2) = \frac{1}{(s!)^2} (n_1 \cdot \pi^{(k)} \cdot n_2)^s$$
(4.1)

$$\tilde{A}_{1}^{(s)}(k \cdot n_{1} \cdot n_{2}) = \frac{1}{(s!)^{2}} (n_{1} \cdot \pi^{(k)} \cdot n_{2})^{s-2} (n_{1} \cdot \pi^{(k)} \cdot n_{1}) (n_{2} \cdot \pi^{(k)} \cdot n_{2})$$

$$\dots \qquad (4.2)$$

$$\tilde{A}_{l}^{(s)}(k \cdot n_{1} \cdot n_{2}) = \frac{1}{(s!)^{2}} (n_{1} \cdot \pi^{(k)} \cdot n_{2})^{s-2l} (n_{1} \cdot \pi^{(k)} \cdot n_{1})^{l} (n_{2} \cdot \pi^{(k)} \cdot n_{2})^{l}$$
(4.3)

where  $n \cdot \pi^{(k)} \cdot m = n^{\mu} \pi_{\mu\nu} m^{\nu}$ . There are  $\lfloor s/2 \rfloor$  independent such terms.

Let us set

$$\tilde{E}^{(s)}(k \cdot n_1 \cdot n_2) = \sum_{l=0}^{\lfloor s/2 \rfloor} a_l \tilde{A}_l^{(s)}(k \cdot n_1 \cdot n_2)$$
(4.4)

where  $a_l$  are arbitrary constants. The explicit conserved structures are obtained by differentiating s times  $E^{(s)}$  with respect to  $n_1$  and s times with respect to  $n_2$ . One obtains in this way conserved tensors  $\tilde{E}^{(s)}_{\mu_1...\mu_s,\nu_1...\nu_s}(k)$ . Conservation is a consequence of the transversality property

$$k^{\mu}\pi_{\mu\nu} = 0 \tag{4.5}$$

and  $\tilde{E}^{(s)}$  writes

$$\tilde{E}_{\mu_1\dots\mu_s,\nu_1\dots\nu_s}(k) = \sum_{l=0}^{\lfloor s/2 \rfloor} a_l \tilde{A}_{l,\mu_1\dots\mu_s,\nu_1\dots\nu_s}(k)$$
(4.6)

This is the most general conserved structure for spin s (for a proof, see appendix A).

By Fourier anti-transforming and inserting into (2.21), one can construct the effective action corresponding to (4.6) multiplied by  $k^2$  for the spin s field  $B_{\mu_1...\mu_s,\nu_1...\nu_s}$  as follows

$$S_E \sim \int d^d x \, B^{\mu_1 \dots \mu_s} \Box E(\partial)_{\mu_1 \dots \mu_s, \nu_1 \dots \nu_s} B^{\nu_1 \dots \nu_s}$$

$$\tag{4.7}$$

where  $E(\partial)$  is the formal Fourier transform of  $\tilde{E}(k)$ , i.e. the same expression with  $k_{\mu}$  replaced by  $-i\partial_{\mu}$ . The eom is of course

$$\Box E(\partial)_{\mu_1\dots\mu_s,\nu_1\dots\nu_s} B^{\nu_1\dots\nu_s} = 0 \tag{4.8}$$

After canonically normalization, it depends on  $\lfloor s/2 \rfloor - 1$  arbitrary constants. This is the most general linearized eom for a completely symmetric spin s field.

From  $\tilde{E}^{(s)}(k)$  we can obtain the most general traceless combination, by taking the trace of (4.6) and imposing it to vanish. This can be done by differentiating the implicit expressions (4.1),..., (4.3),... with respect to  $\frac{\partial}{\partial n_{1\mu}} \frac{\partial}{\partial n_{1}^{\mu}}$ . The resulting equation is the recurrence relation

$$a_{l} = -\frac{(s-2l+2)(s-2l+1)}{2l(2(s-l-1)+d-1)}a_{l-1}$$
(4.9)

Setting  $a_0 = 1$  the solution is

$$a_{l} = \frac{(-1)^{l}}{2^{2l}l!} \frac{s!}{(s-2l)!} \frac{\Gamma\left(s + \frac{d-3}{2} - l\right)}{\Gamma\left(s + \frac{d-3}{2}\right)}$$
(4.10)

Replacing this in (4.6) we obtain a traceless conserved structure. In turn this gives rise to a traceless eom.

#### 4.1 Eom's from conserved structures

Any conserved structure (4.4) in coordinate space is in general a non-local differential operator. To each there corresponds a quadratic Lagrangian and a linearized eom. For the EOM it is enough to differentiate s times with respect to  $n_2^{\nu}$  and saturate the exposed indices with the spin s tensor field  $a^{\nu_1 \dots \nu_s}$ , multiply by  $k^2$ , set the result to zero and then

differentiate also s times w.r.t.  $n_1^{\mu}$ . For the Lagrangian one saturates the l.h.s. of the EOM with  $a^{\mu_1...\mu_s}$  and divide by 2.

Therefore we can represent the eom symbolically as

$$k^2 \sum_{l=0}^{\lfloor s/2 \rfloor} a_l \tilde{A}_l^{(s)}(k \cdot n_1 \cdot n_2) = 0$$

$$(4.11)$$

In the following instead of contracting the  $n_2$  indices with the field a, we will always leave  $n_2$  free and operate only on  $n_1$ . The operation will be essentially tracing two  $n_1$  indices. For instance tracing  $(n_1 \cdot \pi^{(k)} \cdot n_1)$  over  $n_1$  gives d-1.

Let us consider the spin 3 case. In this compact notation, the most general eom will be

$$k^{2} \left( a(n_{1} \cdot \pi^{(k)} \cdot n_{2})^{3} + b(n_{1} \cdot \pi^{(k)} \cdot n_{1})(n_{1} \cdot \pi^{(k)} \cdot n_{2})(n_{2} \cdot \pi^{(k)} \cdot n_{2}) \right) = 0$$
(4.12)

Taking the trace over  $n_1$  gives

$$(6a + (d+1)b)(n_1 \cdot \pi^{(k)} \cdot n_2)(n_2 \cdot \pi^{(k)} \cdot n_2) = 0$$
(4.13)

Thus, unless 6a + (d+1)b = 0, i.e. for generic coefficients a and b, the second piece of (4.12) vanishes on shell and we can simply drop it. Therefore the relevant eom for spin 3 is

$$k^2 (n_1 \cdot \pi^{(k)} \cdot n_2)^3 = 0 \tag{4.14}$$

i.e. (3.24).

Now we wish to prove that this is general, that is, for any spin s, for generic coefficients, the eom can be reduced to the form

$$k^2 (n_1 \cdot \pi^{(k)} \cdot n_2)^s = 0 \tag{4.15}$$

The strategy consists in taking the trace of (4.11) w.r.t. to  $n_1$  the maximum number of times and replacing the results in (4.11). For instance, for spin 4 we have to trace twice.

Tracing p times (4.11) we get

$$\sum_{l=p}^{\lfloor s/2 \rfloor} [c_{l-1}^{(p)}(s-2l+2)(s-2l+1) + 2c_{l}^{(p)}(l-p+1)(s+d-2p-1)] \cdot (n_{1} \cdot \pi^{(k)} \cdot n_{1})^{l-2} (n_{1} \cdot \pi^{(k)} \cdot n_{2})^{s-2l} (n_{2} \cdot \pi^{(k)} \cdot n_{2})^{l} = 0$$
(4.16)

where  $c_l^{(0)} = a_l, c_l^{(1)} = a_{l-1}(s-2l+2)(s-2l+1)+2la_l(s+d-3)$ , etc. The complete expression for  $c_l^{(p)}$  is not easy to compute, but these coefficients are generically non-vanishing. It is however possible to infer the important property that

$$c_l^{(p)} = 0, \qquad l (4.17)$$

For s = 2n after n tracings, i.e. p = n, we arrive at

$$k^2 (n_2 \cdot \pi^{(k)} \cdot n_2)^n = 0 \tag{4.18}$$

Now let us consider p = n - 1. Using (4.16) and (4.18) we arrive at

$$2k^{2}(d+1)c_{n-1}^{(n-1)}(n_{1}\cdot\pi^{(k)}\cdot n_{2})^{2}l(n_{2}\cdot\pi^{(k)}\cdot n_{2})^{n-1}=0$$

So, generically,

$$k^{2}(n_{1}\cdot\pi^{(k)}\cdot n_{2})^{2}l(n_{2}\cdot\pi^{(k)}\cdot n_{2})^{n-1} = 0$$
(4.19)

Now we proceed by induction. Suppose after q traces, i.e. p = n - q + 1, we have

$$k^{2}(n_{1}\cdot\pi^{(k)}\cdot n_{2})^{2i}(n_{2}\cdot\pi^{(k)}\cdot n_{2})^{n-i}=0, \qquad i=0,\ldots,q-1$$

Then, at level p = n - q, we remain with

$$2k^2 c_{n-q}^{(n-q)} (d+2q-1) (n_1 \cdot \pi^{(k)} \cdot n_2)^2 q (n_2 \cdot \pi^{(k)} \cdot n_2)^{n-q} = 0$$
(4.20)

from which the conclusion (4.15) follows.

For s = 2n + 1, we start from p = n

$$2k^2 c_n^{(n)} (d-1) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right) (n_2 \cdot \pi^{(k)} \cdot n_2)^n = 0$$
(4.21)

and we can repeat the induction procedure arriving at the same conclusion (4.15).

The next task is to recover the Fronsdal equation from (4.15).

To this end we take the trace of (4.15), i.e. apply to it  $\frac{\partial}{\partial n_1^{\mu}} \eta^{\mu\nu} \frac{\partial}{\partial n_1^{\nu}}$ . This is easily seen to give

$$\operatorname{tr}(n_1 \cdot \pi^{(k)} \cdot n_2)^2 = 2(n_2 \cdot n_2) - 2\frac{(n_2 \cdot k)^2}{k^2}$$
(4.22)

and, in general,

$$\operatorname{tr}(n_1 \cdot \pi^{(k)} \cdot n_2)^s = s(s-1)(n_1 \cdot \pi^{(k)} n_2)^{s-2} \operatorname{tr}(n_1 \cdot \pi^{(k)} \cdot n_2)^2 = 0$$
(4.23)

Using this we can easily calculate all the traces of (4.15). The end result is

$$\operatorname{tr}(n_1 \cdot \pi^{(k)} \cdot n_2)^s \sim \left( (n_2 \cdot n_2) - 2 \frac{(n_2 \cdot k)^2}{k^2} \right)^{\frac{s}{2}} = 0$$
(4.24)

for even s, and

$$\operatorname{tr}(n_1 \cdot \pi^{(k)} \cdot n_2)^s \sim \left( (n_2 \cdot n_2) - 2\frac{(n_2 \cdot k)^2}{k^2} \right)^{\frac{s-1}{2}} (n_1 \cdot \pi^{(k)} \cdot n_2) = 0$$
(4.25)

for odd s. These two equations have to be understood as follows: any solution that satisfies (4.15) also satisfies either (4.24) or (4.25). Therefore we can replace these two eqs. into (4.15). The viceversa is not true in general: i.e. if a solution satisfies (4.24) or (4.25), it may not satisfy (4.15). For the time being we assume that eq. (4.24) and (4.25) imply that in (4.15) we can make the replacement  $(n_2 \cdot k)^2 = k^2 (n_2 \cdot n_2)$  (see the comment below).

The result of this substitution is:

$$k^{2}(n_{1} \cdot n_{2})^{s} - s(n_{1} \cdot n_{2})^{s-1}(n_{1} \cdot k)(n_{2} \cdot k) + \binom{s}{2}(n_{1} \cdot n_{2})^{s-2}(n_{1} \cdot k)^{2}(n_{2} \cdot n_{2}) + \sum_{l=3}^{s}(-1)^{l}\binom{s}{l}(n_{1} \cdot n_{2})^{s-l}\frac{(n_{1} \cdot k)^{l}(n_{2} \cdot k)^{l-2}}{(k^{2})^{l-2}}(n_{2} \cdot n_{2}) = 0$$

$$(4.26)$$

The first line gives the spin s Fronsdal operator. Therefore (4.26) identifies the spin s nonlocal Fronsdal equation. The compensator takes the form

$$\alpha(n_1, n_2) = \sum_{l=3}^{s} (-1)^l \binom{s}{l} (n_1 \cdot n_2)^{s-l} \frac{(n_1 \cdot k)^{l-3} (n_2 \cdot k)^{l-2}}{(k^2)^{l-2}} (n_2 \cdot n_2)$$
(4.27)

#### 4.2 Conserved odd parity structures

It is easy to obtain also all the odd parity structures. The spin 1 odd parity conserved Lorentz structure (linear in  $n_1 \cdot n_2 \cdot k$ ) can only be

$$\tilde{C}_0^{(1)}(k \cdot n_1 \cdot n_2) = (n_1 \epsilon n_2 \cdot k), \qquad (n_1 \epsilon n_2 \cdot k) = \epsilon_{\mu\nu\lambda} n_1^{\mu} n_2^{\nu} k^{\lambda}$$
(4.28)

It is easy to realize that, for higher spin, the epsilon tensor can only appear in the form  $(n_1 \epsilon n_2, k)$  in every single term, thus it can be factored out. What remains is an even spin structure of one order less. So the most general odd conserved Lorentz structure will be a combination of

$$\tilde{C}_{0}^{(s)}(k \cdot n_{1} \cdot n_{2}) = (n_{1}\epsilon n_{2} \cdot k)\tilde{A}_{0}^{(s-1)}(k \cdot n_{1} \cdot n_{2}) 
\tilde{C}_{1}^{(s)}(k \cdot n_{1} \cdot n_{2}) = (n_{1}\epsilon n_{2} \cdot k)\tilde{A}_{1}^{(s-1)}(k \cdot n_{1} \cdot n_{2}) 
\dots \dots \\
\tilde{C}_{l}^{(s)}(k \cdot n_{1} \cdot n_{2}) = (n_{1}\epsilon n_{2} \cdot k)\tilde{A}_{l}^{(s-1)}(k \cdot n_{1} \cdot n_{2}) 
\dots \dots \dots$$
(4.29)

where  $A_0^{(0)} = 1$ , by definition. Let us define

$$\tilde{O}^{(s)}(k \cdot n_1 \cdot n_2) = \sum_{l=0}^{\lfloor s/2 \rfloor} c_l \tilde{C}_l^{(s)}(k \cdot n_1 \cdot n_2)$$
(4.30)

from which we can derive

$$O_{\mu_1\dots\mu_s,\nu_1\dots\nu_s}(\partial) = \sum_{l=0}^{\lfloor s/2 \rfloor} c_l C_{l,\mu_1\dots\mu_s,\nu_1\dots\nu_s}(\partial)$$
(4.31)

The odd parity action is supposed to be local (and higher derivative)

$$S_{O} = \int d^{d}x \, B^{\mu_{1}...\mu_{s}} \Box^{s-1} O(\partial)_{\mu_{1}...\mu_{s},\nu_{1}...\nu_{s}} B^{\nu_{1}...\nu_{s}}$$
(4.32)

Therefore the odd eom is

$$\Box^{s-1} O_{\mu_1 \dots \mu_s, \nu_1 \dots \nu_s}(\partial) B^{\nu_1 \dots \nu_s} = 0$$
(4.33)

The tracelessness condition (for spin s > 1) implies a recursion relation for the coefficients  $c_l$ :

$$c_{l} = -\frac{(s-2l+1)(s-2l)}{2l(2(s-l-2)+d+1)}c_{l-1}$$
(4.34)

Setting  $c_0 = 1$  the solution is:

$$c_l = \frac{(-1)^l}{2^{2l}l!} \frac{(s-1)!}{(s-2l-1)!} \frac{\Gamma\left(s + \frac{d-3}{2} - l\right)}{\Gamma\left(s + \frac{d-3}{2}\right)}$$
(4.35)

A comment on the non-local Fronsdal equation. In the previous derivations of eqs. (4.26) and (4.33), we have simplified a few steps by disregarding a number of alternatives. First, we have stated that several passages are *generic*, that is they do not hold in some very specific cases, leaving out in this way several (probably pathological) possibilities. Moreover, we have disregarded solutions that satisfy (4.24) or (4.25), but not (4.15). Therefore our conclusions concerning eqs. (4.26) and (4.33) are generic. They do not address more subtle questions, in particular the one pointed out in [47, 48]: the non-locality of the Fronsdal equation contains a large freedom, so an important issue is to select the form of the equation that gives rise to the correct propagator for the higher spin field, and not all non-local equations which give rise to the Fronsdal equation upon gauge fixing also give the correct propagator.<sup>5</sup> We cannot say, on the basis of our previous derivation, that our non-local Fronsdal equations have the property of generating the correct propagator, but we can verify this *a posteriori*, by analyzing the effective actions we obtain for the massive scalar and fermion models in various dimensions. We will return to this issue in the concluding section.

#### 5 The general method

In this section we illustrate the method to compute the 2-pt functions with Feynman diagrams. On first reading one can skip this section and go directly to the results in the next one. The integrals we have to compute in this paper are like the ones in (2.27) and (2.28), that is of the general form

$$\tilde{J}_{\mu_1\dots\mu_p}(d;\alpha,\beta;q_1,q_2,m) = \int \frac{d^d p}{(2\pi)^d} \frac{p_{\mu_1}\dots p_{\mu_p}}{\left((p+q_1)^2 - m^2\right)^{\alpha} \left((p+q_2)^2 - m^2\right)^{\beta}}$$
(5.1)

where, eventually,  $q_1 = 0, q_2 = -k$ . We will use the method invented by [2–4] to reduce the tensor integral to a sum of scalar ones

$$\tilde{J}_{\mu_{1}...\mu_{p}}\left(d;\alpha,\beta,\gamma;q_{1},q_{2},m\right) = \sum_{\substack{\lambda,\kappa_{1},\kappa_{2}\\2\lambda+\sum\kappa_{i}=p}} \left(-\frac{1}{2}\right)^{\lambda} (4\pi)^{p-\lambda} \left\{ [\eta]^{\lambda} [q_{1}]^{\kappa_{1}} [q_{2}]^{\kappa_{2}} \right\}_{\mu_{1}...\mu_{p}} \times (\alpha)_{\kappa_{1}} (\beta)_{\kappa_{2}} \tilde{I}^{(2)}(d+2(p-\lambda);\alpha+\kappa_{1},\beta+\kappa_{2};q_{1},q_{2},m), \quad (5.2)$$

where the symbol  $\left\{ [\eta]^{\lambda} [q_1]^{\kappa_1} \dots [q_N]^{\kappa_N} \right\}_{\mu_1 \dots \mu_M}$  stands for the complete symmetrization of the objects inside the curly brackets, for example

$$\{\eta q_1\}_{\mu_1\mu_2\mu_3} = \eta_{\mu_1\mu_2}q_{1\mu_3} + \eta_{\mu_1\mu_3}q_{1\mu_2} + \eta_{\mu_2\mu_3}q_{1\mu_1}.$$

The basic integral is now the scalar one

$$\tilde{I}^{(2)}(d;\alpha,\beta;q_1,q_2,m) = \int \frac{d^d p}{(2\pi)^d} \frac{1}{((p+q_1)^2 - m^2)^{\alpha} ((p+q_2)^2 - m^2)^{\beta}}$$
(5.3)

 $<sup>{}^{5}</sup>$ We thank the referee of this paper for stressing the importance of such an issue. On the other hand our intention in this section is to stress the universal features of the non-local Fronsdal equation.

For instance, the bubble integral for the s = 1 current in the scalar model

$$\tilde{J}_{\mu\nu}(k) = \int \frac{d^d p}{(2\pi)^d} \frac{(2p-k)_{\mu}(2p-k)_{\nu}}{(p^2-m^2)((p-k)^2-m^2)}$$
(5.4)

reduces to

$$\tilde{J}_{\mu\nu}(m,k) = -\frac{8\pi}{(2\pi)^{d+2}} \eta_{\mu\nu} \tilde{I}^{(2)}(d+2;1,1) + 8\frac{(4\pi)^2}{(2\pi)^{d+4}} k_{\mu}k_{\nu} \tilde{I}^{(2)}(d+4;1,3) \qquad (5.5)$$
$$+\frac{16\pi}{(2\pi)^{d+2}} k_{\mu}k_{\nu} \tilde{I}^{(2)}(d+2;1,2) + \frac{1}{(2\pi)^d} k_{\mu}k_{\nu} \tilde{I}^{(2)}(d;1,1)$$

The integral  $\tilde{I}^{(2)}(d; \alpha, \beta; k, m)$  can be cast into the form of a hypergeometric series

$$\tilde{I}_{IR}^{(2)}(d;\alpha,\beta;k,m) = 2^{-d}\pi^{-d/2}i^{1-d} \left(-m^2\right)^{-\alpha-\beta+\frac{d}{2}} \frac{\Gamma\left(-\frac{d}{2}+\alpha+\beta\right)}{\Gamma(\alpha+\beta)} \times {}_{3}F_2\left(\alpha,\beta,-\frac{d}{2}+\alpha+\beta \left|\frac{k^2}{4m^2}\right| \frac{k^2}{4m^2}\right)$$
(5.6)

This representation is valid for large m compared to k. When m is small compared to k another representation is available

$$\tilde{I}_{UV}^{(2)}(d;\alpha,\beta;k,m) = 2^{-d}\pi^{-d/2}i^{1-d} (k^2)^{-\alpha-\beta+\frac{d}{2}} \left\{ \frac{\left(\Gamma\left(\frac{d}{2}-\alpha\right)\Gamma\left(\frac{d}{2}-\beta\right)\Gamma\left(-\frac{d}{2}+\alpha+\beta\right)\right)}{\Gamma(\alpha)\Gamma(\beta)\Gamma(d-\alpha-\beta)} \times {}_{3}F_{2}\left(-\frac{d}{2}+\alpha+\beta,\frac{-d+\alpha+\beta+1}{2},\frac{-d+\alpha+\beta+2}{2}\left|\frac{4m^{2}}{k^{2}}\right) + \left(-\frac{m^{2}}{k^{2}}\right)^{\frac{d}{2}-\alpha}\frac{\Gamma\left(\alpha-\frac{d}{2}\right)}{\Gamma(\alpha)} {}_{3}F_{2}\left(\beta,\frac{-\alpha+\beta+1}{2},\frac{-\alpha+\beta+2}{2}\left|\frac{4m^{2}}{k^{2}}\right) + \left(-\frac{m^{2}}{k^{2}}\right)^{\frac{d}{2}-\beta}\frac{\Gamma\left(\beta-\frac{d}{2}\right)}{\Gamma(\beta)} {}_{3}F_{2}\left(\alpha,\frac{\alpha-\beta+1}{2},\frac{\alpha-\beta+2}{2}\left|\frac{4m^{2}}{k^{2}}\right)\right) + \left(-\frac{m^{2}}{k^{2}}\right)^{\frac{d}{2}-\beta}\frac{\Gamma\left(\beta-\frac{d}{2}\right)}{\Gamma(\beta)} {}_{3}F_{2}\left(\alpha,\frac{\alpha-\beta+1}{2},\frac{\alpha-\beta+2}{2}\left|\frac{4m^{2}}{k^{2}}\right)\right) \right\}$$

In the sequel we consider also massless models. The relevant results can be obtained from the massive models by taking the  $m \to 0$  limit. But they can also be obtained by setting m = 0 from the very beginning. In such a case the basic integral is

$$\tilde{I}^{(2)}(d;\alpha,\beta;q_1,q_2,0) = \int \frac{d^d p}{(2\pi)^d} \frac{1}{((p+q_1)^2)^{\alpha} ((p+q_2)^2)^{\beta}}$$
(5.8)  
=  $2^{-d} \pi^{-d/2} i^{1-d} (k^2)^{\frac{d}{2} - \alpha - \beta} \frac{\Gamma\left(\frac{d}{2} - \alpha\right) \Gamma\left(\frac{d}{2} - \beta\right) \Gamma\left(\alpha + \beta - \frac{d}{2}\right)}{\Gamma(\alpha) \Gamma(\beta) \Gamma(d - \alpha - \beta)}$ 

#### 5.1 Guidelines for the calculations

We will now set out to do explicit calculations and derive results for two-point functions in the scalar and fermion model in different dimensions. The method just outlined is the most convenient for our purposes, but it is nevertheless one out of many. In fact, even within it there are different possibilities or schemes. We expect that our results may depend on such schemes, but also to find a criterion to extract the scheme independent part. In most cases this is conservation and finiteness. In particular, by suitably choosing the scheme we will be able, for instance, to obtain both finiteness and conservation for spin 1 current in any dimension in the fermion model. The same is not as easy for higher spin currents. In generic spin current correlators and, therefore, in the corresponding one-loop effective actions, we will find, beside non-conserved terms, also terms that diverge in the IR limit  $m \to \infty$ . Fortunately these terms are finite in number and easy to identify by expanding the OLEA near the IR and the UV. Not only, all the nonconserved and all IR divergent terms are local. It is thus possible to subtract all the terms that diverge in the IR, which include, in particular, all the nonconserved ones and recover both conservation and finiteness in the IR.

In this process a particular attention has to be paid to the terms of order 0 in m, in even dimensions. In some cases they are local and conserved, and appear both in the IR and the UV. Even in this case we follow the attitude of subtracting the IR term from the corresponding UV one, on the assumption that physical information is contained in the difference between the UV and the IR, not in their absolute values.

Finally it should be added that the resulting IR and UV expansions are both convergent.

The calculations in the sequel are mainly carried out using a new Mathematica code [54].

To somewhat abbreviate the following formulas, at times we use the compact notation

$$\Pi_{\mathfrak{a}}^{(2)}(k, n_1, n_2) = \left(n_1 \cdot \pi^k \cdot n_2\right)^2 + \mathfrak{a}(n_1 \cdot \pi^k \cdot n_1)(n_2 \cdot \pi^k \cdot n_2),$$
(5.9)

$$\Pi_{\mathfrak{a}}^{(3)}(k, n_1, n_2) = \left(n_1 \cdot \pi^k \cdot n_2\right)^3 + \mathfrak{a}(n_1 \cdot \pi^k \cdot n_1)(n_1 \cdot \pi^k \cdot n_2)(n_2 \cdot \pi^k \cdot n_2), \quad (5.10)$$

where  $\mathfrak{a}$  is some constant.

The symbol k used in the above formula and in the sequel deserves an explanation: k saturated with  $n_1, n_2$  represents the vector  $k_{\mu}$ , while in the other cases it represents the modulus |k|. Finally, contrary to ([1]), the latter is  $k \equiv |k| = \sqrt{k^2}$ .

#### 6 3d scalar effective field action tomography

In this section we start the analysis of the two-point functions of spin higher than 1 currents.

Before reporting on the general spin s case we would like to analyze in detail a few low spin cases. It is in general possible to obtain compact expressions of the one-loop effective actions. However expanding it in powers of m near the IR and UV limits (an operation we call *tomography*) provides the most interesting information.

It is possible to use the parameter m to cut to slices the two-point function of currents of any spin. Let us consider the case of a massive scalar model (msm) in 3d. The basic formulas are (2.1), (2.2), (2.4), (2.6) and (5.4) together with the analogous ones for higher spins, with d = 3.

#### 6.1 3d msm: spin 1 current

This case is well known and simple, but it is excellent for pedagogical purposes. The exact 2-pt correlator for s = 1 is

$$(n_1 \cdot \tilde{J}(k) \cdot n_2) = \frac{i}{8\pi k^3} \left( -4m^2 \coth^{-1}\left(\frac{2m}{k}\right) + 2km + k^2 \coth^{-1}\left(\frac{2m}{k}\right) \right) (k \cdot n_1) (k \cdot n_2) + \frac{i}{8\pi k} (n_1 \cdot n_2) \left( 4m^2 \coth^{-1}\left(\frac{2m}{k}\right) + 2km - k^2 \coth^{-1}\left(\frac{2m}{k}\right) \right)$$
(6.1)

We can expand (6.1) in power of  $\frac{k}{m}$  (IR) or of  $\frac{m}{k}$  (UV). In the IR case we find

$$\mathcal{O}(m): \quad \frac{im\left(n_1 \cdot n_2\right)}{2\pi} \tag{6.2}$$

$$\mathcal{O}(m^{-1}): -\frac{ik^2}{24\pi m} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)$$
(6.3)

$$\mathcal{O}(m^{-3}): -\frac{ik^4}{480\pi m^3} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)$$
(6.4)

while the even powers of m vanish. The first is a (non-conserved and divergent in the IR limit) local term  $\sim \eta_{\mu\nu}$ , which must be subtracted away. The other terms are all conserved and proportional to the conserved structure

$$n_1 \cdot \pi^{(k)} \cdot n_2, \tag{6.5}$$

The UV expansion is instead

$$\mathcal{O}(m^0): -\frac{k}{16} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)$$
(6.6)

$$\mathcal{O}(m): \quad \frac{im}{2\pi k^2} \left(k \cdot n_1\right) \left(k \cdot n_2\right) \tag{6.7}$$

$$\mathcal{O}(m^2): \frac{m^2}{4k} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)$$
(6.8)

$$\mathcal{O}(m^3): \quad \frac{2im^3}{3\pi k^2} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right) \tag{6.9}$$

In fact we have  $\mathcal{O}(m^{2n}) = 0$  for  $n \ge 2$ . The only nonvanishing terms with even powers of m are  $\mathcal{O}(m^0), \mathcal{O}(m^2)$ . For these terms see the comment below.

Except (6.7) the other terms are conserved and proportional to (6.5). The terms proportional to (6.5) are all non-local in the UV, and local in the IR, in particular (6.3) is local and corresponds to the YM action in 3d, see (3.1).

The two nonconserved terms are (6.2) in the IR and (6.7) in the UV. The first is local and the second is nonlocal, but their divergence is the same and local:

$$-\frac{i}{2\pi}\left(k\!\cdot\!n_2\right)$$

This means that we can cancel it by subtracting a local term,  $\sim m \int d^3x \operatorname{tr}(A^2)$ . This amounts to subtracting the IR contribution (which is local) from the UV one. Indeed we get

$$\mathcal{O}_{\rm UV}(m) - \mathcal{O}_{\rm IR}(m) = -\frac{im}{2\pi} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right) \tag{6.10}$$

So the term of order m in the UV and IR conjure up to reform again the same conserved structure as all the other terms. Taking the UV and IR limits splits apart this conserved structure. The conclusion is that, up to a local term we can view the effective action as a sum of infinite many terms, all proportional to  $n_1 \cdot \pi^{(k)} \cdot n_2$  with coefficients proportional to various monomials of m and k. In compact form:

$$\frac{i}{8\pi k} \left(4m^2 \coth^{-1}\left(\frac{2m}{k}\right) - 2km - k^2 \coth^{-1}\left(\frac{2m}{k}\right)\right) n_1 \cdot \pi^{(k)} \cdot n_2 \tag{6.11}$$

#### 6.2 3d msm: e.m. tensor

We have to consider

$$(n_1^2 \cdot \tilde{T}(k) \cdot n_2^2) = n_1^\mu n_1^\nu \tilde{T}_{\mu\nu\lambda\rho}(k) n_2^\lambda n_2^\rho$$

Expanding in the IR we have

$$\mathcal{O}(m^3): + \frac{2im^3}{3\pi} \left( 2 \left( n_1 \cdot n_2 \right)^2 + \left( n_1 \cdot n_1 \right) \left( n_2 \cdot n_2 \right) \right)$$
(6.12)

$$\mathcal{O}(m): -\frac{im}{6\pi} \left( -(n_2 \cdot n_2) (k \cdot n_1)^2 - 4 (n_1 \cdot n_2) (k \cdot n_2) (k \cdot n_1) + 2k^2 (n_1 \cdot n_2)^2 + k^2 (n_1 \cdot n_1) (n_2 \cdot n_2) - (n_1 \cdot n_1) (k \cdot n_2)^2 \right)$$
(6.13)

$$\mathcal{O}(m^{-1}): \quad \frac{ik^4}{60\pi m} \Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2) \tag{6.14}$$

$$\mathcal{O}(m^{-3}): \frac{ik^6}{1680\pi m^3} \Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2)$$
(6.15)

while all the even powers vanish. The  $\mathcal{O}(m^3)$  and  $\mathcal{O}(m)$  terms are non-conserved, while the other terms are all conserved and proportional to the same structure.

In the UV we have

$$\mathcal{O}(m^0): \ \frac{k^3}{32} \Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2) \tag{6.16}$$

$$\mathcal{O}(m): \ \frac{im}{2\pi k^2} (k \cdot n_1)^2 (k \cdot n_2)^2 \tag{6.17}$$

$$\mathcal{O}(m^2): -\frac{m^2 k}{4} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(6.18)

$$\mathcal{O}(m^{3}): \frac{2im^{3}}{3\pi k^{4}} \left(-3 \left(k \cdot n_{1}\right)^{2} \left(k n_{2}\right)^{2}+k^{2} \left(n_{1} \cdot n_{1}\right) \left(k \cdot n_{2}\right)^{2}+4k^{2} \left(n_{1} \cdot n_{2}\right) \left(k \cdot n_{1}\right) \left(k \cdot n_{2}\right)+k^{2} \left(n_{2} \cdot n_{2}\right) \left(k \cdot n_{1}\right)^{2}\right)$$

$$(6.19)$$

$$\mathcal{O}(m^4): \quad \frac{m^4}{2k} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2) \tag{6.20}$$

$$\mathcal{O}(m^{2p}): 0, \quad \text{for} \quad p \ge 3$$

In fact we have  $\mathcal{O}(m^{2m}) = 0$  for  $m \ge 3$ . The only nonvanishing terms with even powers of m are  $\mathcal{O}(m^0), \mathcal{O}(m^2), \mathcal{O}(m^4)$  (again, about these terms, see the comment below)

All the terms are conserved except  $\mathcal{O}(m)$  and  $\mathcal{O}(m^3)$ . But putting together the analogous non-conserved terms in the UV and IR (that is subtracting the local IR terms from the (nonlocal) UV ones) we recover conservation.

$$\mathcal{O}_{\rm UV}(m) - \mathcal{O}_{\rm IR}(m) = \frac{imk^2}{3\pi} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(6.21)

$$\mathcal{O}_{\rm UV}(m^3) - \mathcal{O}_{\rm IR}(m^3) = -\frac{4im^3}{3\pi} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(6.22)

So we find a result analogous to the 1-current. Up to local terms the effective action is a sum of infinite many terms, all proportional to the same conserved structure (6.22) with coefficients proportional to various monomials of m and k. They form a convergent series both in the IR and in the UV. In compact form:

$$\frac{i}{48\pi k} \left( 48m^4 \coth^{-1}\left(\frac{2m}{k}\right) + 2km\left(5k^2 - 12m^2\right) - 24k^2m^2 \coth^{-1}\left(\frac{2m}{k}\right) + 3k^4 \coth^{-1}\left(\frac{2m}{k}\right) \right) \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(6.23)

It should be noticed that the massless model case gives the result:

$$(n_1^2 \cdot \tilde{T}(k) \cdot n_2^2) = \frac{k^3}{32} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(6.24)

This is conserved *but not traceless*, which is not surprising because a scalar massless model in  $d \ge 3$  is not conformal invariant.

Eq. (6.21) is conserved. It does not coincide with the linearized Einstein-Hilbert action (in particular it is nonlocal), but this is simply a nonlocal version of the same, in the same sense as we have already seen for spin 3 and higher in section 3.

#### 6.3 3d msm: spin 3 current

For the 3-spin current we have in the IR

$$\mathcal{O}(m^{5}): \frac{8im^{5}}{5\pi} \left( 2\left(n_{1} \cdot n_{2}\right)^{3} + 3\left(n_{1} \cdot n_{1}\right)\left(n_{2} \cdot n_{2}\right)\left(n_{1} \cdot n_{2}\right) \right)$$

$$\mathcal{O}(m^{3}): \frac{2im^{3}}{3\pi} \left( 3\left(n_{1} \cdot n_{1}\right)\left(n_{1} \cdot n_{2}\right)\left(k \cdot n_{2}\right)^{2} + 3\left(2\left(n_{1} \cdot n_{2}\right)^{2} + \left(n_{1} \cdot n_{1}\right)\left(n_{2} \cdot n_{2}\right)\right)\left(k \cdot n_{1}\right)\left(k \cdot n_{2}\right) \right)$$
(6.25)

$$+ (n_1 \cdot n_2) \left( 3 (n_2 \cdot n_2) (k \cdot n_1)^2 - k^2 \left( 2 (n_1 \cdot n_2)^2 + 3 (n_1 \cdot n_1) (n_2 \cdot n_2) \right) \right) \right)$$
 (6.26)

$$\mathcal{O}(m): \frac{im}{10\pi} \left( 3 (n_2 \cdot n_2) (k \cdot n_1)^3 + 3 (n_1 \cdot n_2) \left( 3 (k \cdot n_2)^2 - k^2 (n_2 \cdot n_2) \right) (k \cdot n_1)^2 + 3 (k \cdot n_2) \left( (n_1 \cdot n_1) (k \cdot n_2)^2 - k^2 \left( 2 (n_1 \cdot n_2)^2 + (n_1 \cdot n_1) (n_2 \cdot n_2) \right) \right) (k \cdot n_1) + k^2 (n_1 \cdot n_2) \left( k^2 \left( 2 (n_1 \cdot n_2)^2 + 3 (n_1 \cdot n_1) (n_2 \cdot n_2) \right) - 3 (n_1 \cdot n_1) (k \cdot n_2)^2 \right) \right)$$
(6.27)

$$\mathcal{O}(m^{-1}): -\frac{ik^{6}}{140\pi m}\Pi^{(3)}_{\frac{3}{2}}(k, n_{1}, n_{2})$$
(6.28)

The coefficients of even powers in m vanish, while the negative odd powers are all proportional to the conserved structure (6.28). The terms  $\mathcal{O}(m^5), \mathcal{O}(m^3), \mathcal{O}(m)$  are local and non-conserved.

In the UV we have

$$\mathcal{O}(m^0): -\frac{k^5}{64}\Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(6.29)

$$\mathcal{O}(m): \quad \frac{im}{2\pi k^2} (k \, n_1)^3 (k \, n_2)^3 \tag{6.30}$$

$$\mathcal{O}(m^2): \quad \frac{3m^2k^3}{16} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2) \tag{6.31}$$

$$\mathcal{O}(m^{3}): \frac{2im^{3}}{3\pi k^{4}}(k \cdot n_{1})(k \cdot n_{2}) \left(-5 (k \cdot n_{1})^{2} (k \cdot n_{2})^{2}+3k^{2} (n_{1} \cdot n_{1}) (k \cdot n_{2})^{2}+9k^{2} (n_{1} \cdot n_{2}) (k \cdot n_{1})(k \cdot n_{2})+3k^{2} (n_{2} \cdot n_{2}) (k \cdot n_{1})^{2}\right)$$

$$(6.32)$$

$$\mathcal{O}(m^4): -\frac{3m^4k}{4}\Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(6.33)

$$\mathcal{O}(m^{5}): \frac{8im^{3}}{5\pi k^{6}} \left( 6k^{4} (n_{1} \cdot n_{2})^{2} (k \cdot n_{1}) (k \cdot n_{2}) + 3k^{4} (n_{1} \cdot n_{1}) (n_{1} \cdot n_{2}) (k \cdot n_{2})^{2} + 3k^{4} (n_{1} \cdot n_{1}) (n_{2} \cdot n_{2}) (k \cdot n_{1}) (k \cdot n_{2}) + 3k^{4} (n_{1} \cdot n_{2}) (n_{2} \cdot n_{2}) (k \cdot n_{1})^{2} - 3k^{2} (n_{1} \cdot n_{1}) (k \cdot n_{2})^{3} - 9k^{2} (n_{1} \cdot n_{2}) (k \cdot n_{1})^{2} (k \cdot n_{2})^{2} - 3k^{2} (n_{2} \cdot n_{2}) (k \cdot n_{1})^{3} (k \cdot n_{2}) + 5 (k \cdot n_{1})^{3} (k \cdot n_{2})^{3} \right)$$

$$(6.34)$$

$$\mathcal{O}(m^6): \ \frac{m^6}{k} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$

The terms  $\mathcal{O}(m^{2n})$  with  $n \geq 4$  vanish. All terms are conserved, except  $\mathcal{O}(m), \mathcal{O}(m^3), \mathcal{O}(m^5)$ .

Proceeding as above we subtract from the non-conserved terms in the UV the homogeneous local non-conserved terms in the IR and obtain conserved terms:

$$\mathcal{O}_{\rm UV}(m) - \mathcal{O}_{\rm IR}(m) = -\frac{imk^4}{5\pi} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(6.35)

$$\mathcal{O}_{\rm UV}(m^3) - \mathcal{O}_{\rm IR}(m^3) = \frac{4im^3}{3\pi} k^2 \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2) \tag{6.36}$$

$$\mathcal{O}_{\rm UV}(m^5) - \mathcal{O}_{\rm IR}(m^5) = -\frac{16im^5}{5\pi} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(6.37)

Therefore up to local terms the effective action is a sum of infinite many terms, all proportional to the same conserved structure with coefficients proportional to various monomials of m and k. They form a convergent series both in the IR and in the UV. In compact form:

$$\frac{i}{480\pi k} \left(960m^{6} \coth^{-1}\left(\frac{2m}{k}\right) - 480km^{5} - 720k^{2}m^{4} \coth^{-1}\left(\frac{2m}{k}\right) + 320\left(k^{2}\right)^{3/2}m^{3} + 180k^{4}m^{2} \coth^{-1}\left(\frac{2m}{k}\right) - 66k^{4}km - 15k^{6} \coth^{-1}\left(\frac{2m}{k}\right)\right) \Pi_{\frac{3}{2}}^{(3)}(k, n_{1}, n_{2})$$
(6.38)

The term (6.36) is local and gives rise to an eom, which is the nonlocal version of the Fronsdal spin 3 equation of motion we have already met above.

#### 6.4 3d msm: higher spin currents

This scheme repeats itself for higher spin currents. For spin 4 there are 4 non-conserved terms in the IR and 4 in the UV, while the others are conserved or 0. Subtracting the IR non-conserved terms from the corresponding UV ones all the nonvanishing terms turn out to be proportional to the conserved structure:

$$\frac{1}{3} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^4 + \frac{1}{8} \left( n_1 \cdot \pi^{(k)} \cdot n_1 \right)^2 \left( n_2 \cdot \pi^{(k)} \cdot n_2 \right)^2 + \left( n_1 \cdot \pi^{(k)} \cdot n_1 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^2 \left( n_2 \cdot \pi^{(k)} \cdot n_2 \right)$$
(6.39)

All terms with even powers of m vanish, except  $m^0, m^2, m^4, m^6, m^8$ .

For spin 5 there are 5 non-conserved terms in the IR and 5 in the UV, while the others are conserved or 0. Subtracting the IR non-conserved terms from the corresponding UV ones all the nonvanishing terms turn out to be proportional to the conserved structure:

$$(n_1 \cdot \pi^{(k)} \cdot n_2)^5 + \frac{15}{8} (n_1 \cdot \pi^{(k)} \cdot n_1)^2 (n_1 \cdot \pi^{(k)} \cdot n_2) (n_2 \cdot \pi^{(k)} \cdot n_2)^2 + 5(n_1 \cdot \pi^{(k)} \cdot n_1) (n_1 \cdot \pi^{(k)} \cdot n_2)^3 (n_2 \cdot \pi^{(k)} \cdot n_2)$$
(6.40)

All terms with even powers of m vanish, except  $m^0, m^2, m^4, m^6, m^8, m^{10}$ .

**Comment 1.** As we have seen above any conserved structure is connected to a (nonlocal) higher spin field equation of motion. In particular eqs. (6.3) and (6.21) are conserved structures which represent the linearized YM and EH actions, respectively, the second one in a nonlocal version. Eq. (6.36) is non-local and gives rise to a variant of the nonlocal Fronsdal equation discussed in section 3. It is clear that any two-point correlator structure can be uniquely related to a given (linearized) equation of motion. The structure of the 2pt-functions conform to the general discussion in section 4. This will be confirmed by the forthcoming analysis.

It is remarkable that the conserved structures that appear in the above expansions are always the same for any fixed 2pt correlator. As we will see this is not the case for the effective field action originating from a fermion model.

**Comment 2.** The nonvanishing even m power terms are a finite number in all cases. They come from the fact that the UV expansion of  $\operatorname{coth}^{-1}$ 

$$\operatorname{coth}^{-1}\left(\frac{2m}{k}\right) = -\frac{i\pi}{2} + \frac{2m}{k} + \frac{8m^3}{3k^3} + \frac{32m^5}{5k^5} + \mathcal{O}\left(m^6\right)$$
(6.41)

contains the factor  $-\frac{i\pi}{2}$ . This is the reason why they are a finite number and do not contain the factor  $\frac{i}{\pi}$  like the others. The factor  $-\frac{i\pi}{2}$  comes from the logarithmic cut of coth<sup>-1</sup> and it is determined by the choice of the Riemann sheet. So it is scheme dependent. It is interesting to compare the  $\mathcal{O}(m^0)$  results with the massless model case, obtained via (5.8). In the massless case for spin 1 we get

$$-\frac{1}{16}k(n_1 \cdot \pi^{(k)} \cdot n_2) \tag{6.42}$$

for spin 2

$$\frac{k^3}{32}\Pi_{\frac{1}{2}}^{(2)}(k,n_1,n_2) \tag{6.43}$$

and for spin 3

$$-\frac{k^5}{64}\Pi^{(3)}_{\frac{3}{2}}(k,n_1,n_2) \tag{6.44}$$

These correlators are nonlocal and coincide with the  $\mathcal{O}_{\rm UV}(m^0)$  terms evaluated above.<sup>6</sup> To be precise there is an indeterminacy in their sign due to the branch point at k = 0originated from the choice of sign of the square root  $\sqrt{k^2}$ . This indeterminacy is present also in the  $m \to 0$  limit of the massive model and it is related to the choice of Riemann sheet mentioned above. As a consequence of it, in this paper we do not worry about the sign in front of the Maxwell and EH kinetic terms that appear in the effective actions. We postpone to a future work the task of finding a physically consistent prescription that eliminates this indeterminacy.

#### 7 3d fermion effective field theory action tomography

We consider now the same analysis for the massive fermion model (mfm). The starting point are eqs. (2.7), (2.10), (2.27), (2.28) and the like for higher spins (see also [1]).

#### 7.1 3d mfm: spin 1 current

This case is rather simple. It takes a very compact form

$$(n_1 \cdot \tilde{J}(k) \cdot n_2) = \frac{i}{8\pi k} \left( -\left(4m^2 \coth^{-1}\left(\frac{2m}{k}\right) - 2km + k^2 \coth^{-1}\left(\frac{2m}{k}\right)\right) (n_1 \cdot \pi^{(k)} \cdot n_2) + 4im \coth^{-1}\left(\frac{2m}{k}\right) \epsilon (k \cdot n_1 \cdot n_2) \right)$$

$$(7.1)$$

and is conserved without any subtraction.

Expanding, the term

$$\mathcal{O}(m^0): -\frac{\epsilon \left(k \cdot n_1 \cdot n_2\right)}{4\pi} \tag{7.2}$$

corresponds to the linearized CS action (here  $\epsilon (k \cdot n_1 \cdot n_2)$  means  $\epsilon_{\mu\nu\rho} k^{\mu} n_1^{\nu} n_2^{\rho}$ ), and the term

$$\mathcal{O}(m^{-1}): -\frac{i}{12\pi m} k^2 \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)$$
(7.3)

in the IR corresponds to the linearized YM action.

<sup>6</sup>Appendix B contains a complete analysis of two-point functions for massless scalar and fermion models.

#### 7.23d mfm: e.m. tensor — even part

For the e.m. tensor we have in the IR (all formulas below have to be multiplied by the factor  $\frac{1}{16}$ )

$$\mathcal{O}(m^3): -2\frac{im^3}{3\pi} \left( (n_1 \cdot n_2)^2 + (n_1 \cdot n_1) (n_2 \cdot n_2) \right)$$
(7.4)

$$\mathcal{O}(m): -\frac{imk^2}{6\pi}\Pi^{(2)}_{-1}(k, n_1, n_2)$$
(7.5)

$$\mathcal{O}(m^{-1}): \quad \frac{ik^4}{40\pi m} \Pi^{(2)}_{-\frac{1}{3}}(k, n_1, n_2) \tag{7.6}$$

$$\mathcal{O}(m^{-3}): \quad \frac{ik^6}{672m^3\pi} \Pi^{(2)}_{-\frac{1}{5}}(k, n_1, n_2) \tag{7.7}$$

The even powers vanish. The  $\mathcal{O}(m^3)$  term is not conserved, while the other terms are all conserved and proportional to different combinations of the two conserved structures.

In the UV we have

$$\mathcal{O}(m^0): \ \frac{k^3}{32} \Pi^{(2)}_{-\frac{1}{2}}(k, n_1, n_2) \tag{7.8}$$

$$\mathcal{O}(m): 0 \tag{7.9}$$

$$\mathcal{O}(m^2): \quad \frac{m^2}{8} k(n_1 \cdot \pi^{(k)} \cdot n_1)(n_2 \cdot \pi^{(k)} \cdot n_2) \tag{7.10}$$

$$\mathcal{O}(m^{3}): -2\frac{im^{3}}{3\pi k^{4}} \left( \left( k^{2} \left( n_{2} \cdot n_{2} \right) - 2 \left( k \cdot n_{2} \right)^{2} \right) \left( k \cdot n_{1} \right)^{2} + 2k^{2} \left( n_{1} \cdot n_{2} \right) \left( k \cdot n_{2} \right) \left( k \cdot n_{1} \right) + k^{2} \left( n_{1} \cdot n_{1} \right) \left( k \cdot n_{2} \right)^{2} \right)$$
(7.11)

$$\mathcal{O}(m^4): -\frac{m^4}{2k} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(7.12)

$$\mathcal{O}(m^5): -2\frac{4im^5}{5\pi k^2} \Pi^{(2)}_{\frac{1}{3}}(k, n_1, n_2)$$

$$\mathcal{O}(m^6): 0$$
(7.13)

$$O(m^0): 0$$

These are all conserved except  $\mathcal{O}(m^3)$ . But putting together the analogous non-conserved term in the UV and IR (that is subtracting the local IR term from the (nonlocal) UV one) we recover conservation:

$$\mathcal{O}_{\rm UV}(m^3) - \mathcal{O}_{\rm IR}(m^3) = 2\frac{im^3}{3\pi}\Pi_1^{(2)}(k, n_1, n_2)$$
(7.14)

eq. (7.5) is the linearized and *local* version of the EH equation of motion (see section 3). The other are non-local versions of the same (except (7.10)). Actually, according to our general philosophy the term  $\mathcal{O}_{IR}(m)$ , which is divergent in the IR limit, must be subtracted. It will therefore appear in the place of the vanishing term (7.9) with inverted sign.

Once again up to local terms the effective action is a sum of infinite many terms, which form a convergent series both in the IR and in the UV, all of them proportional to various combinations of the conserved structures with coefficients proportional to various monomials of m and k. In compact form:

$$-\frac{i}{96\pi k} \left(96m^{4} \coth^{-1}\left(\frac{2m}{k}\right) - 48km^{3} - 4k^{3}m - 6k^{4} \coth^{-1}\left(\frac{2m}{k}\right)\right) \left(n_{1} \cdot \pi^{(k)} \cdot n_{2}\right)^{2} -\frac{i}{96\pi k} \left(48m^{4} \coth^{-1}\left(\frac{2m}{k}\right) - 24km^{3} - 24k^{2}m^{2} \coth^{-1}\left(\frac{2m}{k}\right) + 10k^{3}m +3k^{4} \coth^{-1}\left(\frac{2m}{k}\right)\right) (n_{1} \cdot \pi^{(k)} \cdot n_{1})(n_{2} \cdot \pi^{(k)} \cdot n_{2}).$$
(7.15)

#### 7.3 3d mfm: e.m. tensor — odd part

In the IR (all formulas below have to be multiplied by the factor  $\frac{1}{16}$ )

$$\mathcal{O}(m^3): 0$$
  

$$\mathcal{O}(m^2): -\frac{m^2}{\pi} (n_1 \cdot n_2) \epsilon (k \cdot n_1 \cdot n_2)$$
(7.16)

$$\mathcal{O}(m^0): \quad \frac{k^2}{12\pi} \epsilon \left(k \cdot n_1 \cdot n_2\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right) \tag{7.17}$$

$$\mathcal{O}(m^{-2}): \frac{k^4}{240m^2\pi} \epsilon \left(k \cdot n_1 \cdot n_2\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)$$
(7.18)

the odd powers vanish. The  $\mathcal{O}(m^2)$  term is not conserved, while the other terms are all conserved and proportional to the unique odd conserved structure.

In the UV:

$$\mathcal{O}(m^0): 0$$
  
$$\mathcal{O}(m): -\frac{ikm}{8} \epsilon \left(k \cdot n_1 \cdot n_2\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)$$
(7.19)

$$\mathcal{O}(m^2): -\frac{m^2}{\pi k^2} \left(k \cdot n_1\right) \left(k \cdot n_2\right) \epsilon \left(k \cdot n_1 \cdot n_2\right)$$
(7.20)

$$\mathcal{O}(m^3): \quad i\frac{m^3}{2k}\epsilon\left(k\cdot n_1\cdot n_2\right)\left(n_1\cdot \pi^{(k)}\cdot n_2\right)$$
$$\mathcal{O}(m^4): \quad -\frac{4m^4}{3\pi k}\epsilon\left(k\cdot n_1\cdot n_2\right)\left(n_1\cdot \pi^{(k)}\cdot n_2\right)$$

 $\mathcal{O}(m^2)$  is not conserved, but

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{m^2}{\pi} \epsilon \left( k \cdot n_1 \cdot n_2 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right) \tag{7.21}$$

is. In summary, after subtracting  $\mathcal{O}_{\mathrm{IR}}(m^2)$  the odd 2-pt correlator is:

$$-\frac{m}{4\pi k}\left(4m^2 \operatorname{coth}^{-1}\left(\frac{2m}{k}\right) - 2km - k^2 \operatorname{coth}^{-1}\left(\frac{2m}{k}\right)\right) \epsilon\left(k \cdot n_1 \cdot n_2\right) \left(n_1 \cdot \pi^k, n_2\right) \quad (7.22)$$

The term (7.17) and, in a scaling limit, also (7.19), give rise to the linearized CS action as discussed in [1].

#### 7.4 3d mfm: spin 3, even part

This was already discussed in [1], so we report here only the final results. One must subtract the local terms  $\mathcal{O}(m^5), \mathcal{O}(m^3)$  in the IR, which are not conserved. After which the effective action becomes

$$-\frac{i}{216\pi k} \left( 192m^{6} \coth^{-1}\left(\frac{2m}{k}\right) - 96km^{5} - 48k^{2}m^{4} \coth^{-1}\left(\frac{2m}{k}\right)$$
(7.23)  
+16k<sup>3</sup>m<sup>3</sup> - 12k<sup>4</sup>m<sup>2</sup> \coth^{-1}\left(\frac{2m}{k}\right) - 6k^{5}m + 3k^{6} \coth^{-1}\left(\frac{2m}{k}\right) \right) \left(n\_{1} \cdot \pi^{(k)} \cdot n\_{2}\right)^{3}  
$$-\frac{i}{288\pi k} \left( 384m^{6} \coth^{-1}\left(\frac{2m}{k}\right) - 192km^{5} - 128k^{2}m^{4} \coth^{-1}\left(\frac{2m}{k}\right) + 48k^{3}m^{3} \right)$$
  
+28k<sup>4</sup>m<sup>2</sup>  $\coth^{-1}\left(\frac{2m}{k}\right) + 6k^{5}m - 3k^{6} \coth^{-1}\left(\frac{2m}{k}\right) \right) (n_{1} \cdot \pi^{(k)} \cdot n_{1})(n_{1} \cdot \pi^{(k)} \cdot n_{2})(n_{2} \cdot \pi^{(k)} \cdot n_{2})$ 

The  $\mathcal{O}_{\text{IR}}(m)$  term is conserved and has to be subtracted from it. The interpretation of these conserved structures in terms of massless Fronsdal eom has been discussed above. At each order they are different combinations of two conserved structures

$$(n_1 \cdot \pi^{(k)} \cdot n_2)^3$$
 and  $(n_1 \cdot \pi^{(k)} \cdot n_1)(n_1 \cdot \pi^{(k)} \cdot n_2)(n_2 \cdot \pi^{(k)} \cdot n_2)$  (7.24)

but it is actually easy to prove that all these combinations give rise to the same eom (after taking the trace of the resulting equation and re-inserting it). The only condition is that the coefficient of the first structure be nonvanishing.

#### 7.5 3d mfm: spin 3, odd part

One must subtract the local terms  $\mathcal{O}(m^4)$ ,  $\mathcal{O}(m^2)$  in the IR, which are not conserved. After which the effective action becomes:

$$-\frac{1}{216\pi k} \left(96m^{5} \coth^{-1}\left(\frac{2m}{k}\right) - 48km^{4} - 4k^{3}m^{2} - 3k^{4}m \coth^{-1}\left(\frac{2m}{k}\right)\right)$$
  

$$\cdot \epsilon \left(k \cdot n_{1} \cdot n_{2}\right) \left(n_{1} \cdot \pi^{(k)} \cdot n_{1}\right) \left(n_{2} \cdot \pi^{(k)} \cdot n_{2}\right)$$
(7.25)  

$$-\frac{1}{54\pi k} \left(48m^{5} \coth^{-1}\left(\frac{2m}{k}\right) - 24km^{4} - 24k^{2}m^{3} \coth^{-1}\left(\frac{2m}{k}\right) + 10k^{3}m^{2} + 3k^{4}m \coth^{-1}\left(\frac{2m}{k}\right)\right) \epsilon \left(k \cdot n_{1} \cdot n_{2}\right) \left(n_{1} \cdot \pi^{(k)} n_{2}\right)^{2}$$

The meaning of the term  $\mathcal{O}(m)$  in the UV (in the scaling limit)

$$-i\frac{m}{|k|}\epsilon \left(k\cdot n_{1}\cdot n_{2}\right)k^{4} \left(\frac{1}{36} \left(n_{1}\cdot \pi^{(k)}\cdot n_{2}\right)^{2} - \frac{1}{144} (n_{1}\cdot \pi^{(k)}\cdot n_{1})(n_{2}\cdot \pi^{(k)}\cdot n_{2})\right)$$
(7.26)

and  $\mathcal{O}(m^0)$  in the IR

$$\epsilon \left(k \cdot n_1 \cdot n_2\right) \left(\frac{k^4}{240\pi} (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) - \frac{2k^4}{135\pi} \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^2\right)$$
(7.27)

have already been discussed in [1].

#### 8 Tomography in 5d

There is no substantial difference between 3d and 5d. We start from the same formulas as in 3d and change only the dimension. For obvious reasons of readability we limit ourselves to the even parity part and the lowest spins, although the generalization is at hand.

#### 8.1 5d scalar model

#### 8.1.1 5d msm: spin 1 current

The analog of eq. (6.1) is

$$(n_1 \cdot \tilde{J}(k) \cdot n_2) = -\frac{i}{768\pi^2 k^3} \left( k^2 (n_1 \cdot n_2) \left( 40km^3 + 3\left(k^2 - 4m^2\right)^2 \coth^{-1}\left(\frac{2m}{k}\right) - 6k^3m \right) + 3\left(8km^3 - \left(k^2 - 4m^2\right)^2 \coth^{-1}\left(\frac{2m}{k}\right) + 2k^3m\right) (k \cdot n_1) (k \cdot n_2) \right)$$
(8.1)

This is not conserved, but the divergence is local. Expanding in powers of m like in 3d, we get in the IR

$$\mathcal{O}(m^3): -\frac{im^3}{12\pi^2} (n_1 \cdot n_2) \tag{8.2}$$

$$\mathcal{O}(m) : \frac{im}{48\pi^2} k^2 (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(8.3)

$$\mathcal{O}(m^{-1}) : -\frac{\imath}{960\pi^2 m} k^4 (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(8.4)

All terms corresponding to even powers of m vanish. In the UV we have instead

$$\mathcal{O}(m^0): -\frac{1}{512\pi}k^3(n_1 \cdot \pi^{(k)} \cdot n_2)$$
(8.5)

$$\mathcal{O}(m): 0$$
  
 $\mathcal{O}(m^2): m^2 \frac{1}{64\pi} k(n_1 \cdot \pi^{(k)} \cdot n_2)$  (8.6)

$$\mathcal{O}(m^3): -m^3 \frac{i}{12\pi^2} \frac{(k \cdot n_1) (k \cdot n_2)}{k^2}$$
 (8.7)

$$\mathcal{O}(m^4): -m^4 \frac{1}{32\pi k} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
 (8.8)

All even powers of  $m \ge 6$  vanish. All these terms are conserved except  $\mathcal{O}_{IR}(m^3)$  and  $\mathcal{O}_{UV}(m^3)$ . But once again  $\mathcal{O}_{IR}(m^3)$  is local and can be subtracted, and

$$\mathcal{O}_{\rm UV}(m^3) - \mathcal{O}_{\rm IR}(m^3) = \frac{im^3}{12\pi^2} (n_1 \cdot \pi^{(k)} \cdot n_2) \tag{8.9}$$

The term  $\mathcal{O}(m)$  is conserved but divergent in the IR limit. Therefore, according to our recipe, it must be subtracted and will appear with opposite sign in the UV list, where the corresponding term is missing. This term yields the Maxwell (or linearized YM) action and EOM, with a coupling  $\sim m$ .

#### 8.1.2 5d msm: spin 2 current

For the full 2-pt function of the e.m. tensor is much too cumbersome see section 11. As expected it is not conserved, as will be clear from the expansion in powers of m, but the terms responsible for the non-conservation are local. In the IR we have

$$\mathcal{O}(m^5) : -\frac{im^5}{15\pi^2} \left( 2 \left( n_1 \cdot n_2 \right)^2 + \left( n_1 \cdot n_1 \right) \left( n_2 \cdot n_2 \right) \right)$$
(8.10)

$$\mathcal{O}(m^3) : -\frac{im^3}{36\pi^2} \left( (n_2 \cdot n_2) (k \cdot n_1)^2 + 4 (n_1 \cdot n_2) (k \cdot n_2) (k \cdot n_1) -2k^2 (n_1 \cdot n_2)^2 - k^2 (n_1 \cdot n_1) (n_2 \cdot n_2) + (n_1 \cdot n_1) (k \cdot n_2)^2 \right)$$
(8.11)

$$\mathcal{O}(m): -\frac{imk^4}{120\pi^2} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(8.12)

$$\mathcal{O}(m^{-1}): \frac{ik^6}{3360\pi^2 m} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(8.13)

The terms corresponding to odd powers of m vanish. In the UV we have

$$\mathcal{O}(m^0): \ \frac{k^5}{1536\pi} \Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2) \tag{8.14}$$

$$\mathcal{O}(m^2): -\frac{m^2 k^3}{128\pi} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(8.15)

$$\mathcal{O}(m^3): -\frac{im^3}{12\pi^2 k^2} (kn_1)^2 (k \cdot n_2)^2$$
(8.16)

$$\mathcal{O}(m^4): \quad \frac{m^4 k}{32\pi} \prod_{\frac{1}{2}}^{(2)} (k, n_1, n_2) \tag{8.17}$$

$$\mathcal{O}(m^{5}): -\frac{im^{5}}{15\pi^{2}k^{4}} \left(k^{2} \left(n_{1} \cdot n_{1}\right) \left(k \cdot n_{2}\right)^{2} + 4k^{2} \left(n_{1} \cdot n_{2}\right) \left(k \cdot n_{1}\right) \left(k \cdot n_{2}\right) + k^{2} \left(n_{2} \cdot n_{2}\right) \left(k \cdot n_{1}\right)^{2} - 3k^{4} \left(k \cdot n_{1}\right)^{2} \left(k \cdot n_{2}\right)^{2}\right)$$

$$(8.18)$$

$$\mathcal{O}(m^6): -\frac{m^6}{24\pi k} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(8.19)

The terms  $\mathcal{O}(m)$  and  $\mathcal{O}(m^{2k})$  with  $k \ge 4$  vanish. All the nonvanishing terms are conserved except those of order 3 and 5. But the non-conserved terms in the IR are local and

$$\mathcal{O}_{\rm UV}(m^5) - \mathcal{O}_{\rm IR}(m^5) = \frac{2im^5}{15\pi^2} \Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2)$$
(8.20)

$$\mathcal{O}_{\rm UV}(m^3) - \mathcal{O}_{\rm IR}(m^3) = -\frac{im^3k^2}{18\pi^2} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(8.21)

These are conserved. Eq. (8.21) gives rise to a (nonlocal) version of the linearized EH action. Also in this case the term  $\mathcal{O}_{\text{IR}}(m)$  must be subtracted, although conserved, because it is divergent in the IR; as a consequence it will appear with opposite sign in the UV list, where the corresponding term is missing.

#### 8.1.3 5d msm: spin 3 current

Once again the 2pt correlators of spin 3 currents can be calculated exactly, see section 11, but we will skip it here and go to the IR expansion. The terms  $\mathcal{O}(m^7), \mathcal{O}(m^5), \mathcal{O}(m^3)$  are not conserved, but local, while

$$\mathcal{O}(m): \quad \frac{imk^{6}}{280\pi^{2}} \Pi_{\frac{3}{2}}^{(3)}(k, n_{1}, n_{2}) \tag{8.22}$$

$$\mathcal{O}(m^{-1}): -\frac{ik^8}{10080\pi^2 m} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(8.23)

Moreover  $\mathcal{O}_{\mathrm{IR}}(m^n) = 0$  for n even.

Near the UV the nonvanishing terms are:

$$\mathcal{O}(m^0): -\frac{k^7}{4096\pi} \prod_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(8.24)

$$\mathcal{O}(m^2): \frac{m^2 k^5}{256\pi} \prod_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(8.25)

$$\mathcal{O}(m^4): -\frac{3m^4k^3}{128\pi} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(8.26)

$$\mathcal{O}(m^6): \frac{km^6}{16\pi} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(8.27)

$$\mathcal{O}(m^8): -\frac{m^8}{16\pi k} \prod_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(8.28)

while  $\mathcal{O}(m^{2n}) = 0$  for  $n \ge 5$ . As for the odd *m* power terms they are conserved for  $n \ge 9$ :

$$\mathcal{O}(m^9): -\frac{32im^9}{315\pi^2 k^2} \prod_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(8.29)

while  $\mathcal{O}(m) = 0$  and  $\mathcal{O}(m^3), \mathcal{O}(m^5), \mathcal{O}(m^7)$  are non-local and non-conserved. But once again

$$\mathcal{O}_{\rm UV}(m^7) - \mathcal{O}_{\rm IR}(m^7) = \frac{8im^7}{35\pi^2} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(8.30)

$$\mathcal{O}_{\rm UV}(m^5) - \mathcal{O}_{\rm IR}(m^5) = -\frac{2im^5k^2}{15\pi^2} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(8.31)

$$\mathcal{O}_{\rm UV}(m^3) - \mathcal{O}_{\rm IR}(m^3) = \frac{im^3k^4}{30\pi^2} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)) \tag{8.32}$$

The term (8.31) corresponds to the spin 3 Fronsdal EOM. As we see from these examples the scheme for 5d is similar to 3d. Once again the term  $\mathcal{O}_{\text{IR}}(m)$  must be subtracted, although conserved, because it is divergent in the IR; as a consequence it will appear with opposite sign in the UV list, where the corresponding term is missing.

#### 8.2 5d fermion model

#### 8.2.1 5d mfm: spin 1 current

The analog of eq. (7.1) (for the even part) is

$$(n_1 \cdot \tilde{J}(k) \cdot n_2) = -\frac{i}{128\pi^2 k} \left( -16m^4 \coth^{-1}\left(\frac{2m}{k}\right) + 8km^3 - 8k^2m^2 \coth^{-1}\left(\frac{2m}{k}\right) - 6k^3m + 3k^4 \coth^{-1}\left(\frac{2m}{k}\right) \right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(8.33)

which is conserved.

Expanding in powers of m like in 3d, all coefficients have of course the same conserved structure. In the IR all even m-power coefficient vanish and, for instance,

$$\mathcal{O}(m): \frac{im}{12\pi^2} k^2 (n_1 \cdot \pi^{(k)} \cdot n_2) \tag{8.34}$$

which (with reversed sign) corresponds to the Maxwell action. In the UV we have instead,

$$\mathcal{O}(m^0): -\frac{3}{256\pi}k^3(n_1 \cdot \pi^{(k)} \cdot n_2)$$
(8.35)

$$\mathcal{O}(m^2): m^2 \frac{1}{32\pi} k(n_1 \cdot \pi^{(k)} \cdot n_2)$$
 (8.36)

$$\mathcal{O}(m^4): m^4 \frac{1}{16\pi k} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
 (8.37)

$$\mathcal{O}(m^5): m^5 \frac{4i}{15\pi^2 k^2} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
 (8.38)

$$\mathcal{O}(m^7): \ m^7 \frac{32i}{105\pi^2 k^4} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(8.39)

... ...

while  $\mathcal{O}(m) = \mathcal{O}(m^3) = \mathcal{O}(m^n) = 0$  for even  $n \ge 6$ . According to our recipe the term  $\mathcal{O}(m)$  must be subtracted and will appear in the UV list with opposite sign

### 8.2.2 5d mfm: e.m. tensor

In this subsection every result must be multiplied by a factor of  $\frac{1}{16}$ .

In the IR the even m-power coefficients vanish. The nonvanishing ones are

$$\mathcal{O}(m^5): \frac{2im^5}{15\pi^2} \left( (n_1 \cdot n_2)^2 + (n_1 \cdot n_1)(n_2 \cdot n_2) \right)$$
(8.40)

$$\mathcal{O}(m^3): \quad \frac{ik^2m^3}{18\pi^2}\Pi^{(2)}_{-1}(k, n_1, n_2) \tag{8.41}$$

$$\mathcal{O}(m): -\frac{ik^4m}{40\pi^2}\Pi^{(2)}_{-\frac{1}{3}}(k, n_1, n_2)$$
(8.42)

$$\mathcal{O}(m^{-1}): \quad \frac{ik^6}{672\pi^2 m} \Pi^{(2)}_{-\frac{1}{5}}(k, n_1, n_2)$$
(8.43)

Except  $\mathcal{O}(m^5)$  they are all conserved. In the UV we find  $\mathcal{O}(m) = \mathcal{O}(m^3) = \mathcal{O}(m^{2n}) = 0$ , for even  $n \ge 4$ , and

$$\mathcal{O}(m^0): \ \frac{k^5}{384\pi} \prod_{k=1}^{(2)} (k, n_1, n_2)$$
(8.44)

$$\mathcal{O}(m^2): -m^2 \frac{k^3}{64\pi} \prod_{-\frac{1}{2}}^{(2)} (k, n_1, n_2)$$
(8.45)

$$\mathcal{O}(m^4): -m^4 \frac{k}{32\pi} (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2)$$
(8.46)

$$\mathcal{O}(m^6): \ \frac{m^6}{12\pi k} \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2) \tag{8.47}$$

$$\mathcal{O}(m^7): \quad \frac{8im^7}{35\pi k^2} \Pi^{(2)}_{\frac{1}{3}}(k, n_1, n_2) \tag{8.48}$$

 $\mathcal{O}(m^5)$  is nonlocal and non-conserved, but

$$\mathcal{O}_{\rm UV}(m^5) - \mathcal{O}_{\rm IR}(m^5) = -\frac{2im^5}{15\pi^2} \Pi_1^{(2)}(k, n_1, n_2)$$
(8.49)

The remaining terms are conserved. In particular  $\mathcal{O}_{\rm UV}(m^2)$  corresponds to the linearized EH action. The terms  $\mathcal{O}_{\rm IR}(m)$ ,  $\mathcal{O}_{\rm IR}(m^3)$  are conserved but divergent in the IR limit. So they must be subtracted and will appear in the UV list with opposite sign.

#### 8.2.3 5d mfm: spin 3 current

We give a brief account because this case varies with respect to the scalar model only in one respect: the various conserved terms in the m expansion do not have always the same conserved structure like in the latter case. In the IR the even power terms vanish, while the odd power terms  $\mathcal{O}(m^n)$  are nonvanishing for  $n \leq 7$ . Moreover  $\mathcal{O}(m^7), \mathcal{O}(m^5)$  are not conserved, while all the others are. For instance

$$\mathcal{O}(m^3): -\frac{8im^3k^4}{405\pi^2} \Pi^{(3)}_{-1}(k, n_1, n_2)$$
(8.50)

$$\mathcal{O}(m): \quad \frac{4imk^6}{945\pi^2} \Pi^{(3)}_{-\frac{37}{64}}(k, n_1, n_2) \tag{8.51}$$

In the UV  $\mathcal{O}(m) = \mathcal{O}(m^3) = \mathcal{O}(m^n) = 0$  for even  $n \ge 10$ , while  $\mathcal{O}(m^0), \mathcal{O}(m^2), \mathcal{O}(m^4), \mathcal{O}(m^6), \mathcal{O}(m^8)$  are nonvanishing and conserved. For instance

$$\mathcal{O}(m^8): \frac{m^8}{18\pi k} \prod_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(8.52)

The odd powers  $\mathcal{O}(m^n)$  are nonvanishing for n > 0 and conserved except for n = 5, 7. But again

$$\mathcal{O}_{\rm UV}(m^5) - \mathcal{O}_{\rm IR}(m^5) = \frac{im^5k^2}{15\pi^2} \left( (n_1 \cdot \pi^{(k)} \cdot n_1)(n_1 \cdot \pi^{(k)} \cdot n_2)(n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$
(8.53)

$$\mathcal{O}_{\rm UV}(m^7) - \mathcal{O}_{\rm IR}(m^7) = -\frac{128im^7}{945\pi^2} \Pi^{(3)}_{\frac{23}{16}}(k, n_1, n_2)$$
(8.54)

It is curious that the fermionic model in 5d does not reproduce exactly the spin 3 Fronsdal operator. In fact the term (8.53) has the right form but lacks the essential  $k^2 (n_1 \cdot \pi^{(k)} \cdot n_2)^3$  part. This has to be considered a combinatorial coincidence. The terms  $\mathcal{O}_{\text{IR}}(m)$ ,  $\mathcal{O}_{\text{IR}}(m^3)$  are conserved but divergent in the IR limit. So they must be subtracted and will appear in the UV list with opposite sign.

**Comment.** The structure of the 2pt functions in 5d essentially repeats the scheme of 3d. The *m*-power expansions both in the IR and in the UV are similar: in the IR there are non conserved *local* terms, while in the UV there are non-conserved *nonlocal* terms. Subtracting the former from the latter one obtains conserved structures (and a finite IR limit). All the other terms are conserved and have analogous types of structures in both the fermionic and the scalar model.

#### 9 Tomography in 4d

Even dimensional models present an additional problem concerning their regularization. For odd d works by itself as a complete regulator in carrying out the integrals generated by the Feynman diagrams. This is not anymore true for even d. The way out is well-known, we will set  $d = 4 + \varepsilon$ . Another difference we will come across with, which is related to this, is the appearance of *log* terms in the form factors. We will again expand the two-point functions in powers of m near the IR and UV limits.

In almost all the two-point correlators and, therefore, in all the one-loop effective actions, we will find non-conserved terms and terms that diverge in the IR  $m \to \infty$ , like in the odd dimensional case, but we will find also  $\varepsilon$ -divergent terms. Our general attitude is to recover both conservation and finiteness in the IR. This is possible because all the nonconserved and all divergent terms in the IR, as well as all  $\varepsilon$ -divergent terms, are local. We will therefore subtract all the terms that diverge in the IR and in  $\varepsilon$ . They include, in particular, all the nonconserved ones.

There remains however an ambiguity. Beside divergent and/or nonconserved terms, in the case of  $m^0$  we meet also finite contributions, both in the IR and in the UV. Also for these terms we subtract the IR from the UV contribution, on the assumption that it is this difference that contains the physical information.

#### 9.1 4d scalar model

The basic formulas are again (2.1), (2.2), (2.4), (2.6) and (5.4) together with the analogous ones for higher spins, with  $d = 4 + \varepsilon$ .

#### 9.1.1 4d msm: spin 1 current

The full formula for the 2pt correlator is expressed in terms of hypergeometric functions and parameter derivatives thereof, and we dispense with writing it down explicitly here, see however section 11. We will focus on the power of m expansions. As just mentioned, we have to consider also  $\log(m)$  and  $\frac{1}{\varepsilon}$  factors. In the IR we find

$$\mathcal{O}(m^2): -\frac{im^2}{8\pi^2} \left(\gamma - 1 - \log(4\pi) + 2\log(m) + \frac{2}{\varepsilon}\right) (n_1 \cdot n_2)$$

$$\mathcal{O}(m): 0$$

$$(9.1)$$

$$\mathcal{O}(\log(m)): \quad \frac{i\log(m)}{24\pi^2} k^2 (n_1 \cdot \pi^k \cdot n_2) \tag{9.2}$$

$$\mathcal{O}(m^0): \quad \frac{ik^2}{48\pi^2} \left(\gamma - \log(4\pi) + \frac{2}{\varepsilon}\right) (n_1 \cdot \pi^{(k)} \cdot n_2) \tag{9.3}$$
$$\mathcal{O}(m^{-1}): \quad 0$$

$$\mathcal{O}(m^{-2}): -\frac{ik^4}{480\pi^2 m^2} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(9.4)

These coefficients are conserved except  $\mathcal{O}(m^2)$ . All the odd powers of m vanish.

In the UV we find:

$$\mathcal{O}(m^{0}): -i\frac{k^{2}}{144\pi^{2}} \left(8 - 3\gamma - \log\left(\frac{1}{64\pi^{3}}\right) - 3\log\left(-k^{2}\right) - \frac{6}{\varepsilon}\right) (n_{1} \cdot \pi^{(k)} \cdot n_{2})$$
(9.5)

$$\mathcal{O}(m): 0 \tag{9.6}$$

$$\mathcal{O}(m^{2}): -\frac{im^{2}}{24\pi^{2}k^{2}} \left( \left( -3\log\left(-\frac{k^{2}}{m^{2}}\right) + 3\right) (k \cdot n_{1})(k \cdot n_{2}) + k^{2}(n_{1} \cdot n_{2}) \left( 3(-2 + \gamma - \log(4\pi)) + 3\log\left(-k^{2}\right) + \frac{6}{\varepsilon} \right) \right)$$
(9.7)

$$\mathcal{O}(m^3): 0 \tag{9.8}$$

$$\mathcal{O}(m^4): -i\frac{m^4}{16\pi^2 k^2} \left(-2\log\left(-\frac{k^2}{m^2}\right) - 3\right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
... (9.9)

All odd powers of m vanish. The even powers are conserved except (9.7). Subtracting from the latter the analogous (local) non-conserved term in the IR we find a conserved term

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = -\frac{im^2}{8\pi^2} \left(2\log\left(-\frac{k^2}{m^2}\right) - 1\right) (n_1 \cdot \pi^{(k)} \cdot n_2) \tag{9.10}$$

The  $\mathcal{O}(\log(m))$  term is divergent in the IR, and the  $\mathcal{O}(m^0)$  is divergent in the  $\varepsilon \to 0$  limit. Luckily they are local and can be subtracted with the following result:

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) = -\frac{ik^2}{144\pi^2} \left(-3\log\left(-\frac{k^2}{m^2}\right) + 8\right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(9.11)

This term corresponds to the linearized Maxwell action with an energy dependent coupling.

# 9.1.2 4d msm: spin 2 current

In the IR the odd powers of m vanish. The nonvanishing even powers are

$$\mathcal{O}(m^4): -\frac{im^4}{16\pi^2} \left( 2\left(n_1 \cdot n_2\right)^2 + (n_1 \cdot n_1)(n_2 \cdot n_2) \right) \left( 2\gamma - 3 - 2\log(4\pi) + 4\log(m) + \frac{4}{\varepsilon} \right)$$
(9.12)

$$\mathcal{O}(m^{2}): -\frac{im^{2}}{24\pi^{2}} \left( (n_{2} \cdot n_{2}) (k \cdot n_{1})^{2} + 4(n_{1} \cdot n_{2})(k \cdot n_{2})(k \cdot n_{1}) -k^{2} \left( 2 (n_{1} \cdot n_{2})^{2} + (n_{1} \cdot n_{1})(n_{2} \cdot n_{2}) \right) + (n_{1} \cdot n_{1}) (k \cdot n_{2})^{2} \right) \\ \cdot \left( \gamma - 1 - \log(4\pi) + 2\log(m) + \frac{2}{\varepsilon} \right)$$
ilor(m)

$$\mathcal{O}(\log(m)): -\frac{i\log(m)}{60\pi^2}k^4\Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(9.14)

$$\mathcal{O}(m^{0}): -\frac{ik^{4}}{120\pi^{2}}\Pi_{\frac{1}{2}}^{(2)}(k, n_{1}, n_{2})\left(\gamma - \log(4\pi) + \frac{2}{\varepsilon}\right)$$
(9.15)

The first two terms are not conserved, the logarithmic term is conserved but divergent in the IR, the  $m^0$  term is divergent in the limit  $\varepsilon \to 0$ . They all must be subtracted. The remaining terms are conserved.

In the UV all the odd powers of m vanish. The nonvanishing even powers are

$$\mathcal{O}(m^{0}): \frac{ik^{4}}{1800\pi^{2}} \left( 46 - 15\gamma + 15\log(4\pi) - 15\log(-k^{2}) - \frac{30}{\varepsilon} \right)$$

$$\cdot \Pi_{\frac{1}{2}}^{(2)}(k, n_{1}, n_{2})$$

$$(9.17)$$

$$\mathcal{O}(m^{2}): -\frac{im^{2}}{72\pi^{2}} \left( \left( \frac{6}{\varepsilon} - 8 + 3\gamma + \log\left(\frac{1}{64\pi^{3}}\right) + 3\log\left(-k^{2}\right) \right) \right)$$

$$\times \left( (n_{2} \cdot n_{2}) \left(k \cdot n_{1}\right)^{2} + 4(n_{1} \cdot n_{2})(k \cdot n_{2})(k \cdot n_{1}) + (n_{1} \cdot n_{1}) \left(k \cdot n_{2}\right)^{2} - k^{2} \left( 2\left(n_{1} \cdot n_{2}\right)^{2} + (n_{1} \cdot n_{1})(n_{2} \cdot n_{2}) \right) \right)$$

$$- \frac{3}{k^{2}} \left( -3\log\left(-k^{2}\right) + 6\log(m) + 5 \right) \left(k \cdot n_{1}\right)^{2} \left(k \cdot n_{2}\right)^{2} \right)$$
(9.18)

and

$$\mathcal{O}(m^{4}): -\frac{im^{4}}{16\pi^{2}k^{4}} \left(k^{2} \left(2\log\left(-\frac{k^{2}}{m^{2}}\right)-1\right)\right)$$

$$\cdot \left((n_{2} \cdot n_{2}) \left(k \cdot n_{1}\right)^{2}+4(n_{1} \cdot n_{2})(k \cdot n_{2})(k \cdot n_{1})+(n_{1} \cdot n_{1}) \left(k \cdot n_{2}\right)^{2}-3 \left(k \cdot n_{1}\right)^{2} \left(k \cdot n_{2}\right)^{2}\right)$$
(9.19)

$$+2k^{4}\left(2(n_{1}\cdot n_{2})^{2}+(n_{1}\cdot n_{1})(n_{2}\cdot n_{2})\right)\left(-2+\gamma-\log(4\pi)-\log\left(-k^{2}\right)+\frac{2}{\varepsilon}\right)\right)$$

$$\mathcal{O}(m^{6}):\frac{im^{6}}{36\pi^{2}k^{2}}\left(6\log\left(-\frac{k^{2}}{m^{2}}\right)+11\right)\Pi_{\frac{1}{2}}^{(2)}(k,n_{1},n_{2})$$

$$\dots:$$
(9.20)

 $\mathcal{O}(m^0)$  and all terms with even m power larger than 4 are conserved, while  $\mathcal{O}(m^2)$  and  $\mathcal{O}(m^4)$  are not. According to our prescription we have to subtract not only  $\mathcal{O}_{IR}(m^2)$  and  $\mathcal{O}_{IR}(m^4)$ , but also  $\mathcal{O}_{IR}(m^0)$  and  $\mathcal{O}_{IR}(\log(m))$ . We obtain

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = -\frac{im^4}{8\pi^2} \left( +2\log\left(-\frac{k^2}{m^2}\right) - 1 \right) \,\Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2) \tag{9.21}$$

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2}{36\pi^2} k^2 \left( 3\log\left(-\frac{k^2}{m^2}\right) - 5 \right) \,\Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2) \tag{9.22}$$

and

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) = \frac{i}{1800\pi^2} k^4 (-15\log\left(-\frac{k^2}{m^2}\right) + 46) \\ \cdot \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(9.23)

They are all conserved. (9.22) contains a nonlocal linearized version of the EH eom.

#### 9.1.3 4d msm: spin 3 current

The scheme is the same as above. In the IR the odd power of m vanish. The even powers  $m^{2n}$  with  $n \leq 0$  are conserved together with the term proportional to  $\log(m)$ . The terms  $\mathcal{O}_{\mathrm{IR}}(m^2), \mathcal{O}_{\mathrm{IR}}(m^6), \mathcal{O}_{\mathrm{IR}}(m^6)$  are not conserved. Of course  $\mathcal{O}(\log[m])$  diverges in the IR, while the term  $\mathcal{O}_{\mathrm{IR}}(m^0)$  diverges for  $\varepsilon \to 0$ . According to our prescription all these terms, which are local, have to be subtracted from the effective action. The result is as follows:

$$\mathcal{O}(m^{-2}): -\frac{ik^8}{5040m^2\pi^2} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(9.24)

$$\mathcal{O}(m^{-4}): -\frac{ik^{10}}{110880m^4\pi^2} \Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2)$$
(9.25)

In the UV the odd m power terms vanish. The even power of order 2, 4, 6 are not conserved, but

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m))$$

$$= -\frac{ik^6}{29400\pi^2} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left(-105 \log\left(-\frac{k^2}{m^2}\right) + 352\right)$$
(9.26)

and

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2k^4}{300\pi^2} \,\Pi^{(3)}_{\frac{3}{2}}(k, n_1, n_2) \left( (31 - 15\log\left(-\frac{k^2}{m^2}\right)\right) \tag{9.27}$$

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = -\frac{im^4k^2}{24\pi^2} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left( (7 - 6\log\left(\frac{k^2}{m^2}\right) \right)$$
(9.28)

$$\mathcal{O}_{\rm UV}(m^6) - \mathcal{O}_{\rm IR}(m^6) = \frac{im^6}{12\pi^2} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left( \left(1 - 6\log\left(-\frac{k^2}{m^2}\right)\right) \right)$$
(9.29)

$$\mathcal{O}_{\rm UV}(m^8) = \frac{im^8}{48k^2\pi^2} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left( (25 - 12\log\left(-\frac{k^2}{m^2}\right) \right)$$
(9.30)  
...

are all conserved. Eq. (9.28) is related to a nonlocal version of the spin 3 Fronsdal equation.

#### 9.2 4d fermion model

We consider now the same analysis for the fermion massive model. We start again from eqs. (2.7), (2.10), (2.27), (2.28) and the like for higher spins.

## 9.2.1 4d mfm: spin 1 current

The full formula for the 2pt correlator is similar to the scalar case and expressed in terms of parameter derivatives of hypergeometric functions, see section 11. A full expression in terms of simple functions can be found in appendix C. The m-power expansion in the IR is as follows

$$\mathcal{O}(m^2): \quad \frac{im^2}{4\pi^2\varepsilon} \left(n_1 \cdot n_2\right) \tag{9.31}$$

$$\mathcal{O}(\log(m)): \quad \frac{i\log(m)}{6\pi^2} k^2 (n_1 \cdot \pi^k \cdot n_2) \tag{9.32}$$

$$\mathcal{O}(m^{0}): -\frac{i}{24\pi^{2}} \left(k^{2} \left(-2\gamma + 1 + \log(16\pi^{2}) - \frac{4}{\varepsilon}\right) (n_{1} \cdot n_{2}) + 2\left(\gamma - \log(4\pi) + \frac{2}{\varepsilon}\right) (k, n_{1}) (k, n_{2})\right)$$
(9.33)

$$\mathcal{O}(m^{-2}): -\frac{ik^4}{60\pi^2 m^2} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(9.34)

All odd powers of m vanish. The above terms are all conserved except (9.31) and (9.33).  $\mathcal{O}(m^2)$  and  $\mathcal{O}(\log(m))$  are divergent in the IR and  $\mathcal{O}(m^0)$  is divergent in  $\varepsilon$ .

In the UV all odd powers of m vanish, while

$$\mathcal{O}(m^{0}): \frac{i}{24\pi^{2}} \left( \left( \frac{4}{\varepsilon} - \frac{13}{3} + 2\gamma + 2i\pi - \log(16\pi^{2}) \right) k^{2}(n_{1} \cdot n_{2}) - \frac{2}{3} \left( \frac{6}{\varepsilon} - 5 + 3\gamma + 3i\pi - \log(16\pi^{2}) \right) (n_{1} \cdot k) (k \cdot n_{2}) \right)$$
(9.35)

$$\mathcal{O}(m^2): -\frac{im^2}{4\pi^2 k^2} \left( -2(k \cdot n_1)(k \cdot n_2) + k^2(n_1 \cdot n_2) \right)$$
(9.36)

$$\mathcal{O}(m^4): -i\frac{m^4}{4\pi^2 k^2} \left( 2\log\left(-\frac{k^2}{m^2}\right) + 1 \right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
... (9.37)

All the terms are conserved except the first two. But, subtracting from them the corresponding local terms in the IR we get

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) = \frac{i}{36\pi^2} \left( 3\log\left(-\frac{k^2}{m^2}\right) - 5 \right) k^2 \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)$$
(9.38)  
$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2}{m^2} \left(n_1 \cdot \tau^{(k)} \cdot n_2\right)$$
(9.39)

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = -\frac{m}{2\pi^2} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(9.39)

Clearly (9.38) reproduces the Maxwell action.

# 9.2.2 4d mfm: e.m. tensor

A full expression in terms of simple functions can be found in appendix C. In this subsection every result must be multiplied by a factor of  $\frac{1}{16}$ . In the IR the odd powers of m vanish. The nonvanishing even powers are

$$\mathcal{O}(m^{4}): \frac{im^{4}}{8\pi^{2}} \left( (n_{1} \cdot n_{2})^{2} \left( 2\gamma - 1 - 2\log(4\pi) + 4\log(m) + \frac{4}{\varepsilon} \right) \right)$$

$$+ (n_{1} \cdot n_{1})(n_{2} \cdot n_{2}) \left( 2\gamma - 3 - 2\log(4\pi) + 4\log(m) + \frac{4}{\varepsilon} \right) \right)$$

$$\mathcal{O}(m^{2}): \frac{im^{2}}{12\pi^{2}} \left( \left( (n_{2} \cdot n_{2}) (k \cdot n_{1})^{2} + (n_{2} \cdot n_{2})(k \cdot n_{1})^{2} - k^{2}(n_{1} \cdot n_{1})(n_{2} \cdot n_{2}) \right)$$

$$\cdot \left( \gamma - 1 - \log(4\pi) + 2\log(m) + \frac{2}{\varepsilon} \right)$$

$$+ (n_{1} \cdot n_{2}) (k \cdot n_{1}) (k \cdot n_{2}) \left( 3 - 2\gamma + \log(4\pi) - 4\log(m) - \frac{4}{\varepsilon} \right) \right)$$
(9.40)
(9.41)

$$\mathcal{O}(\log(m)): -\frac{i\log(m)}{20\pi^2}k^4\Pi^{(2)}_{-\frac{1}{3}}(k,n_1,n_2)$$
(9.42)

and

$$\mathcal{O}(m^{0}): -\frac{i}{120\pi^{2}} \left( k^{4} \left( \frac{6}{\varepsilon} + 3\gamma - 1 - 3\log(4\pi) \right) (n_{1} \cdot n_{2})^{2} \right)$$

$$+k^{2} \left( -\frac{12}{\varepsilon} - 6\gamma + 1 + 6\log(4\pi) \right) (n_{1} \cdot n_{2}) (k \cdot n_{1}) (k \cdot n_{2}) + \left( \frac{2}{\varepsilon} + \gamma - \log(4\pi) \right)$$

$$\cdot \left( -k^{4} (n_{1} \cdot n_{1}) (n_{2} \cdot n_{2}) + k^{2} (n_{1} \cdot n_{1}) (k \cdot n_{2})^{2} + (k \cdot n_{1})^{2} \left( 2 (k \cdot n_{2})^{2} + k^{2} (n_{2} \cdot n_{2}) \right) \right)$$

$$\mathcal{O}(m^{-2}): \frac{i}{336m^{2}\pi^{2}} k^{6} \Pi^{(2)}_{-\frac{1}{5}} (k, n_{1}, n_{2})$$

$$(9.44)$$

The first two terms are not conserved, the logarithmic term is conserved but divergent in the IR, the  $m^0$  term is not conserved and divergent in the limit  $\varepsilon \to 0$ . They all must be subtracted. The remaining terms are conserved. In the UV all the odd powers of m vanish. The nonvanishing even powers are

$$\mathcal{O}(m^{0}): \frac{i}{1800\pi^{2}} \left( \left( \frac{30}{\varepsilon} + 15\log\left(-k^{2}\right) - 46 + 15\gamma - 15\log(4\pi) \right) \right)$$

$$\left( 9.45 \right)$$

$$\left( k^{4} \left( (n_{1} \cdot n_{1})(n_{2} \cdot n_{2}) - 3\left(n_{1} \cdot n_{2}\right)^{2} \right) - k^{2}(n_{1} \cdot n_{1})\left(k \cdot n_{2}\right)^{2} - k^{2}(n_{2} \cdot n_{2})\left(k \cdot n_{1}\right)^{2} \right)$$

$$+ 3k^{2}(n_{1} \cdot n_{2})(k \cdot n_{1})(k \cdot n_{2}) \left( \frac{60}{\varepsilon} + 30\log\left(-k^{2}\right) - 77 + 30\gamma - 30\log(4\pi) \right)$$

$$- 2\left(k \cdot n_{1}\right)^{2}\left(k \cdot n_{2}\right)^{2} \left( \frac{30}{\varepsilon} + 15\log\left(-k^{2}\right) - 31 + 15\gamma - 15\log(4\pi) \right) \right)$$

$$\mathcal{O}(m^{2}): - \frac{m^{2}}{36\pi^{2}k^{2}}i \left( \left(-6\left(k \cdot n_{1}\right)^{2}\left(k \cdot n_{2}\right)^{2} \right)$$

$$(9.46)$$

$$- k^{2}(n_{1} \cdot n_{2})(k \cdot n_{1})(k \cdot n_{2}) \left( -\frac{12}{\varepsilon} - 6\log\left(-k^{2}\right) + 7 - 6\gamma + 6\log(4\pi) \right)$$

$$-k^{2}(n_{1} \cdot n_{2})(k \cdot n_{1})(k \cdot n_{2})\left(-\frac{12}{\varepsilon} - 6\log(-k^{2}) + 7 - 6\gamma + 6\log(4\pi)\right) \\ + \left(k^{4}\left(-(n_{1} \cdot n_{1})\right)(n_{2} \cdot n_{2}) + k^{2}(n_{1} \cdot n_{1})(k \cdot n_{2})^{2} + k^{2}(n_{2} \cdot n_{2})(k \cdot n_{1})^{2}\right) \\ \cdot \left(\frac{6}{\varepsilon} + 3\log(-k^{2}) - 8 + 3\gamma - 3\log(4\pi)\right) \\ - k^{4}(n_{1} \cdot n_{2})^{2}\left(\frac{6}{\varepsilon} + 3\log(-k^{2}) - 3\log(4\pi) - 5 + 3\gamma\right)\right)$$

and

$$\mathcal{O}(m^{4}): \frac{im^{4}}{16\pi^{2}k^{4}} \left(k^{2} \left(2 \log \left(-\frac{k^{2}}{m^{2}}\right) - 1\right)$$

$$(9.47)$$

$$\cdot \left((n_{2} \cdot n_{2}) (k \cdot n_{1})^{2} + 4(n_{1} \cdot n_{2})(k \cdot n_{2})(k \cdot n_{1}) + (n_{1} \cdot n_{1}) (k \cdot n_{2})^{2} - 3 (k \cdot n_{1})^{2} (k \cdot n_{2})^{2}\right)$$

$$-2k^{4} \left(2 (n_{1} \cdot n_{2})^{2} + (n_{1} \cdot n_{1})(n_{2} \cdot n_{2})\right) \left(-2 + \gamma - \log(4\pi) + \log\left(-\frac{k^{2}}{m^{2}}\right) + \frac{2}{\varepsilon}\right)\right)$$

$$\mathcal{O}(m^{6}): -\frac{im^{6}}{12\pi^{2}k^{2}} \left(6 \log\left(-\frac{k^{2}}{m^{2}}\right) + 7\right) \left(n_{1} \cdot \pi^{(k)} \cdot n_{2}\right)^{2}$$

$$+\frac{im^{6}}{36\pi^{2}k^{2}} \left(6 \log\left(-\frac{k^{2}}{m^{2}}\right) + 11\pi\right) (n_{1} \cdot \pi^{(k)} \cdot n_{1})(n_{2} \cdot \pi^{(k)} \cdot n_{2})$$

$$\dots : \dots$$

All terms with even m power larger than 4, as well as  $\mathcal{O}(log(m))$ , are conserved, while  $\mathcal{O}(m^0), \mathcal{O}(m^2)$  and  $\mathcal{O}(m^4)$  are not. According to our prescription we have to subtract not only  $\mathcal{O}_{\mathrm{IR}}(m^0), \mathcal{O}_{\mathrm{IR}}(m^2)$  and  $\mathcal{O}_{\mathrm{IR}}(m^4)$ , but also  $\mathcal{O}_{\mathrm{IR}}(\log(m))$ . We obtain

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = \frac{im^4}{8\pi^2} \left( +2\log\left(-\frac{k^2}{m^2}\right) - 5\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^2$$

$$+ \frac{im^4}{8\pi^2} \left(2\log\left(-\frac{k^2}{m^2}\right) - 1\right) \left((n_1 \cdot \pi^{(k)} \cdot n_1)(n_2 \cdot \pi^{(k)} \cdot n_2)\right)$$
(9.49)

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2k^2}{36\pi^2} \left( 3\log\left(-\frac{k^2}{m^2}\right) + 1 \right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^2$$

$$-\frac{im^2k^2}{36\pi^2} \left( 3\log\left(-\frac{k^2}{m^2}\right) - 5 \right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2)$$
(9.50)

and

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) = \frac{i}{1800\pi^2} k^4$$

$$\cdot \left(9 \left(-5 \log\left(-\frac{k^2}{m^2}\right) + 12\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^2 - \left(-15 \log\left(-\frac{k^2}{m^2}\right) + 46\right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2)\right)$$
(9.51)

They are all conserved. (9.50) contains a nonlocal linearized version of the EH eom.

#### 9.2.3 4d mfm: spin 3 current

A full expression of the correlator in terms of simple functions can be found in appendix C. The scheme is the same as above. In the IR the odd power of m vanish. The even powers  $m^{2n}$  with n < 0 are conserved together with the term proportional to  $\log(m)$ . The terms  $\mathcal{O}_{IR}(m^0), \mathcal{O}_{IR}(m^2), \mathcal{O}_{IR}(m^6)$  and  $\mathcal{O}_{IR}(m^6)$  are not conserved. Of course  $\mathcal{O}(\log[m])$  diverges in the IR, while the term  $\mathcal{O}_{IR}(m^0)$  diverges for  $\varepsilon \to 0$ . According to our prescription all these terms, which are local, have to be subtracted from the effective action. The result is as follows.

$$\mathcal{O}(m^{-2}): -\frac{ik^8}{2835m^2\pi^2} \Pi^{(3)}_{-\frac{3}{16}}(k, n_1, n_2)$$
(9.52)

$$\mathcal{O}(m^{-4}): -\frac{2ik^{10}}{93555m^4\pi^2} \Pi^{(3)}_{\frac{49}{128}}(k, n_1, n_2)$$
(9.53)

In the UV the odd m power terms vanish. The even power of order 2, 4, 6 are not conserved, but again

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) \tag{9.54}$$

$$= -\frac{2ik^6}{99225\pi^2} \left(-210\log\left(-\frac{k^2}{m^2}\right) + 599\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^3$$

$$+ \frac{ik^6}{1587600\pi^2} \left(-3885\log\left(-\frac{k^2}{m^2}\right) + 13339\right) (n_1 \cdot \pi^{(k)} \cdot n_2)(n_1 \cdot \pi^{(k)} \cdot n_1)(n_2 \cdot \pi^{(k)} \cdot n_2)$$

and

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = -\frac{4im^2k^4}{2025\pi^2} \left(15\log\left(-\frac{k^2}{m^2}\right) - 16\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^3$$
(9.55)  
+
$$\frac{im^2k^4}{16200\pi^2} \left(480\log\left(-\frac{k^2}{m^2}\right) - 857\right) (n_1 \cdot \pi^{(k)} \cdot n_2)(n_1 \cdot \pi^{(k)} \cdot n_1)(n_2 \cdot \pi^{(k)} \cdot n_2)$$

$$\begin{aligned} \mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) &= \frac{4im^4k^2}{27\pi^2} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^3 \\ &\quad -\frac{im^4k^2}{144\pi^2} \left( 18\log\left( -\frac{k^2}{m^2} \right) - 23 \right) (n_1 \cdot \pi^{(k)} \cdot n_2) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \\ \mathcal{O}_{\rm UV}(m^6) - \mathcal{O}_{\rm IR}(m^6) &= \frac{4im^6}{81\pi^2} \left( 6\log\left( -\frac{k^2}{m^2} \right) - 7 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^3 \\ &\quad +\frac{im^6}{162\pi^2} \left( 69\log\left( -\frac{k^2}{m^2} \right) - 70 \right) (n_1 \cdot \pi^{(k)} \cdot n_2) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \\ \mathcal{O}_{\rm UV}(m^8) &= -\frac{2im^8}{81\pi^2k^2} \left( 12\log\left( -\frac{k^2}{m^2} \right) + 19 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^3 \\ &\quad -\frac{im^2k^4}{2592\pi^2k^2} \left( 1356\log\left( -\frac{k^2}{m^2} \right) + 1637 \right) (n_1 \cdot \pi^{(k)} \cdot n_2) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \end{aligned}$$

are all conserved. Eq. 
$$(9.56)$$
 is related to a nonlocal version of the spin 3 Fronsdal equation.

## 10 Tomography in 6d

### 10.1 6d scalar model

. . .

. . .

The basic formulas are again (2.1), (2.2), (2.4), (2.6) and (5.4) together with the analogous ones for higher spins, with  $d = 6 + \varepsilon$ . For the full two-point correlator formulas see next section. Here we limit ourselves to IR and UV expansions.

#### 10.1.1 6d msm: spin 1 current

Like in 4d, we have to consider also  $\log(m)$  and  $\frac{1}{\varepsilon}$  factors. In the IR the nonvanishing terms are

$$\mathcal{O}(m^4): \quad \frac{im^4}{128\pi^3} \left( 2\gamma - 3 - 2\log(4\pi) + 4\log(m) + \frac{4}{\varepsilon} \right) (n_1 \cdot n_2) \tag{10.1}$$

$$\mathcal{O}(m^2): \quad \frac{im^2k^2}{192\pi^3} \left(\gamma - 1 - \log(4\pi) + 2\log(m) + \frac{2}{\varepsilon}\right) (n_1 \cdot \pi^k \cdot n_2) \tag{10.2}$$

$$\mathcal{O}(\log(m)): \frac{i\log(m)}{960\pi^3} k^4(n_1 \cdot \pi^k \cdot n_2)$$
 (10.3)

$$\mathcal{O}(m^0): \ \frac{ik^4}{1920\pi^3} \left(\gamma - \log(4\pi) + \frac{2}{\varepsilon}\right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(10.4)

$$\mathcal{O}(m^{-2}): -\frac{ik^6}{26880\pi^3 m^2} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(10.5)

These coefficients are conserved except  $\mathcal{O}(m^4)$ . All the odd powers of m vanish.

In the UV we find:

$$\mathcal{O}(m^0): \frac{ik^4}{28800\pi^3} \left( -46 + 15\gamma - 15\log(4\pi) + 15\log(-k^2) + \frac{30}{\varepsilon} \right) (n_1 \cdot \pi^{(k)} \cdot n_2) \quad (10.6)$$

$$\mathcal{O}(m^2): \frac{im^2k^2}{576\pi^3} \left( \left( -3\log(-k^2) - 3\gamma + 8 + 3\log(4\pi) - \frac{6}{\varepsilon} \right) \right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(10.7)

$$\mathcal{O}(m^{4}): \frac{im^{4}}{128\pi^{3}k^{2}} \left( \left( -2\log\left(-\frac{k^{2}}{m^{2}}\right) + 1 \right) (k \cdot n_{1})(k \cdot n_{2}) + 2k^{2}(n_{1} \cdot n_{2}) \left( -2 + \gamma - \log(4\pi) + \log\left(-k^{2}\right) + \frac{2}{\varepsilon} \right) \right)$$
(10.8)

All odd powers of m vanish. The even powers are conserved except (9.7). Subtracting from the latter the analogous (local) non-conserved term in the IR we find a conserved term

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = -\frac{im^4}{128\pi^3} \left(-2\log\left(-\frac{k^2}{m^2}\right) + 1\right) (n_1 \cdot \pi^{(k)} \cdot n_2) \tag{10.9}$$

The  $\mathcal{O}(m^2), \mathcal{O}(\log(m))$  terms are divergent in the IR, and the  $\mathcal{O}(m^0)$  is divergent in the  $\varepsilon \to 0$  limit, but they are local and can be subtracted with the following result:

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2k^2}{576\pi^3} \left( \left( -3\log\left(-\frac{k^2}{m^2}\right) + 5\right) \right) (n_1 \cdot \pi^{(k)} \cdot n_2) (10.10)$$
  
$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) = -\frac{ik^4}{28800\pi^3} \left( -15\log\left(-\frac{k^2}{m^2}\right) + 46 \right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(10.11)

(10.10) corresponds to the linearized Maxwell action with an energy dependent coupling.

#### 10.1.2 6d msm: spin 2 current

In the IR the odd powers of m vanish. The nonvanishing even powers are

$$\mathcal{O}(m^6): \frac{im^6}{576\pi^3} \left( 2 \left( n_1 \cdot n_2 \right)^2 + (n_1 \cdot n_1)(n_2 \cdot n_2) \right) \left( 6\gamma - 11 - 6\log(4\pi) + 12\log(m) + \frac{12}{\varepsilon} \right)$$
(10.12)

$$\mathcal{O}(m^4): \frac{im^4}{384\pi^3} \left( (k \cdot n_2)^2 (n_1 \cdot n_1) + 4(n_1 \cdot n_2)(k \cdot n_2)(k \cdot n_1) + (n_2, n_2)(k \cdot n_1)^2 \right)$$
(10.13)

$$-k^{2} \left( 2 \left( n_{1} \cdot n_{2} \right)^{2} + \left( n_{1} \cdot n_{1} \right) \left( n_{2} \cdot n_{2} \right) \right) \left( 2\gamma - 3 - 2 \log(4\pi) + 4 \log(m) + \frac{4}{\varepsilon} \right)$$

$$\mathcal{O}(m^2): \quad \frac{im^2k^4}{480\pi^3} \left(\gamma - 1 - \log(4\pi) + 2\varepsilon \log(m) + \frac{2}{\varepsilon}\right) \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2) \tag{10.14}$$

$$\mathcal{O}(\log(m)): -\frac{i\log(m)}{3360\pi^3}k^6\Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(10.15)

$$\mathcal{O}(m^{0}): -\frac{i}{6720\pi^{3}}k^{6}\left(\gamma - \log(4\pi) + \frac{2}{\varepsilon}\right)\Pi_{\frac{1}{2}}^{(2)}(k, n_{1}, n_{2})$$
(10.16)

The first two terms are not conserved, the logarithmic term is conserved but divergent in the IR, the  $m^0$  term is divergent in the limit  $\varepsilon \to 0$ . They all must be subtracted (including the finite  $\mathcal{O}(m^0)$  part). The remaining terms are conserved.

In the UV all the odd powers of m vanish. The nonvanishing even powers are

$$\mathcal{O}(m^{0}): \frac{ik^{6}}{705600\pi^{3}} \left(352 - 105\gamma + 105\log(4\pi) - 105\log(-k^{2}) - \frac{210}{\varepsilon}\right) \quad (10.17)$$

$$\cdot \Pi_{\frac{1}{2}}^{(2)}(k, n_{1}, n_{2})$$

$$\mathcal{O}(m^{2}): -\frac{im^{2}k^{4}}{7200\pi^{3}} \left(-\frac{30}{\varepsilon} + 46 - 15\gamma + 15\log(4\pi) - 15\log(-k^{2})\right) \quad (10.18)$$

$$\cdot \Pi_{\frac{1}{2}}^{(2)}(k, n_{1}, n_{2})$$

and

$$\mathcal{O}(m^4): \quad \frac{im^4}{1152\pi^3 k^2} \left( 2k^2 \left( 3\log\left(-k^2\right) - 8 + 3\gamma - 3\log(4\pi) + \frac{6}{\varepsilon} \right) \left( (n_2 \cdot n_2) \left(k \cdot n_1\right)^2 + 4(n_1 \cdot n_2)(k \cdot n_2)(k \cdot n_1) + (n_1 \cdot n_1) \left(k \cdot n_2\right)^2 - k^2 \left( 2\left(n_1 \cdot n_2\right)^2 + (n_1 \cdot n_1) \left(n_2 \cdot n_2\right) \right) \right)$$
(10.19)

$$+3\left(-6\log\left(-\frac{k^{2}}{m^{2}}\right)+7\right)\left(k\cdot n_{1}\right)^{2}\left(k\cdot n_{2}\right)^{2}\right)$$

$$\mathcal{O}(m^{6}):-\frac{im^{6}}{576\pi^{3}k^{4}}\left(6\log\left(-\frac{k^{2}}{m^{2}}\right)-1\right)$$
(10.20)

$$\times \left( \left( (n_1 \cdot n_1) (k \cdot n_2)^2 + 4(n_1 \cdot n_2)(k \cdot n_2)(k \cdot n_1) + (k \cdot n_1)^2 (n_2 \cdot n_2) \right) k^2 - 3 (k \cdot n_1)^2 (k \cdot n_2)^2 \right) -6k^4 \left( -2 + \gamma - \log(4\pi) + \log(-k^2) + \frac{2}{\varepsilon} \right) \left( 2 (n_1 \cdot n_2)^2 + (n_1 \cdot n_1) (n_2 \cdot n_2) \right) \\ \mathcal{O}(m^8) : \frac{im^8}{1152\pi^3 k^6} \left( -25 - 12 \log\left(-\frac{k^2}{m^2}\right) \right) \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(10.21)

 $\mathcal{O}(m^0)$  and  $\mathcal{O}(m^2)$  as well as all terms with even m power larger than 6 are conserved, while  $\mathcal{O}(m^4)$  and  $\mathcal{O}(m^6)$  are not. According to our prescription we have to subtract not only  $\mathcal{O}_{\mathrm{IR}}(m^4)$  and  $\mathcal{O}_{\mathrm{IR}}(m^6)$ , but also  $\mathcal{O}_{\mathrm{IR}}(m^0)$ ,  $\mathcal{O}_{\mathrm{IR}}(\log(m))$  and  $\mathcal{O}_{\mathrm{IR}}(m^2)$ . We obtain

$$\mathcal{O}_{\rm UV}(m^6) - \mathcal{O}_{\rm IR}(m^6) = -\frac{im^6}{288\pi^3} \left( -6\log\left(-\frac{k^2}{m^2}\right) + 1 \right) \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2)$$
(10.22)

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = \frac{im^4}{576\pi^3} k^2 \left(-6\log\left(-\frac{k^2}{m^2}\right) + 7\right) \Pi^{(2)}_{\frac{1}{2}}(k, n_1, n_2) \qquad (10.23)$$

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2}{7200\pi^3} k^4 \left(-15\log\left(-\frac{k^2}{m^2}\right) + 31\right) \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2) \quad (10.24)$$

and

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) = \frac{i}{705600\pi^3} k^6 \left(-105 \log\left(-\frac{k^2}{m^2}\right) + 352\right) \\ \cdot \Pi_{\frac{1}{2}}^{(2)}(k, n_1, n_2) \tag{10.25}$$

They are all conserved. (10.23) contains a nonlocal linearized version of the EH eom.

#### 10.1.3 6d msm: spin 3 current

The scheme is the same as above. In the IR the odd power of m vanish. The even powers  $m^{2n}$  with  $n \leq 2$  are conserved together with the term proportional to  $\log(m)$ . The terms  $\mathcal{O}_{\mathrm{IR}}(m^4), \mathcal{O}_{\mathrm{IR}}(m^6), \mathcal{O}_{\mathrm{IR}}(m^8)$  are not conserved. Of course  $\mathcal{O}(\log(m)), \mathcal{O}_{\mathrm{IR}}(m^2)$  diverge in the IR, while the term  $\mathcal{O}_{\mathrm{IR}}(m^0)$  diverges for  $\varepsilon \to 0$ . According to our prescription all these terms, which are local, have to be subtracted from the effective action. Using again  $\Pi^{(3)}(k, n_1, n_2)$ , see (5.10), the result is as follows.

$$\mathcal{O}(m^{-2}): -\frac{ik^{10}}{443520m^2\pi^3} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(10.26)

$$\mathcal{O}(m^{-4}): -\frac{ik^{12}}{11531520m^4\pi^3} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2)$$
(10.27)

In the UV the odd m power terms vanish. The even power of order 4,6,8 are not conserved, but

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m))$$

$$= \frac{ik^8}{6350400\pi^3} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left(-315\log\left(-\frac{k^2}{m^2}\right) + 1126\right)$$
(10.28)

and

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2k^6}{117600\pi^3} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left( (247 - 105\log\left(-\frac{k^2}{m^2}\right) \right) (10.29)$$

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = -\frac{im^2\kappa^2}{4800\pi^3} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left( (47 - 30\log\left(-\frac{\kappa^2}{m^2}\right) \right)$$
(10.30)

$$\mathcal{O}_{\rm UV}(m^6) - \mathcal{O}_{\rm IR}(m^6) = -\frac{im^6k^2}{288\pi^3} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left( (-5 + 6\log\left(-\frac{k^2}{m^2}\right) \right)$$
(10.31)

$$\mathcal{O}_{\rm UV}(m^8) = -\frac{im^8}{384\pi^3} \Pi_{\frac{3}{2}}^{(3)}(k, n_1, n_2) \left( (-1 - 12\log\left(-\frac{k^2}{m^2}\right) \right)$$
(10.32)

are all conserved. Eq. (9.29) is related to a nonlocal version of the spin 3 Fronsdal equation.

## 10.2 6d fermion model

We consider now the same analysis for the fermion massive model. We start again from eqs. (2.7), (2.10), (2.27), (2.28) and the like for higher spins and set  $d = 6 + \varepsilon$ .

#### 10.2.1 6d mfm: spin 1 current

We will limit ourselves to the power of m expansions in the IR. The terms proportional to  $m^4, m^2, m^0$  and  $\log(m)$  are local, nonconserved and/or divergent. Thus they must be subtracted. Therefore the first nonvanishing term in the IR is:

$$\mathcal{O}(m^{-2}): -\frac{ik^6}{1120\pi^3 m^2} (n_1 \cdot \pi^{(k)} \cdot n_2)$$
(10.33)

In the UV, after subtracting the local terms we find:

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) \tag{10.34}$$
$$= -\frac{ik^4}{3600\pi^3} \left(-30\log\left(-\frac{k^2}{m^2}\right) + 77\right) (n_1 \cdot \pi^{(k)} \cdot n_2)$$
$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2k^2}{72\pi^3} \left(\left(-3\log\left(-\frac{k^2}{m^2}\right) + 2\right)\right) (n_1 \cdot \pi^{(k)} \cdot n_2) \tag{10.35}$$

and

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = \frac{im^4}{8\pi^3} (n_1 \cdot \pi^{(k)} \cdot n_2) \tag{10.36}$$

$$\mathcal{O}_{\rm UV}(m^6) = -\frac{im^6}{72\pi^3 k^2} \left( \left( -6\log\left(-\frac{k^2}{m^2}\right) - 5 \right) \right) (n_1 \cdot \pi^{(k)} \cdot n_2) \quad (10.37)$$

(10.35) corresponds to the linearized Maxwell action with an energy dependent coupling.

#### 10.2.2 6d mfm: spin 2 current

In this subsection all results must be multiplied by a factor of  $\frac{1}{16}$  In the IR the odd powers of m vanish. The terms proportional to  $m^6, m^4, m^2, m^0$  and  $\log(m)$  are local, nonconserved and/or divergent. Thus they must be subtracted.

$$\mathcal{O}(m^{-2}): \frac{ik^8}{8640\pi^3 m^2} \left( \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^2 - \frac{1}{7} (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$
(10.38)  
...

In the UV all the odd powers of m vanish. After subtracting the above local terms we have

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) \tag{10.39}$$

$$= \frac{i}{352800\pi^3} k^6 \left( 25 \left( -21 \log \left( -\frac{k^2}{m^2} \right) + 62 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^2 - \left( -105 \log \left( -\frac{k^2}{m^2} \right) + 352 \right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = -\frac{im^2}{3600\pi^3} k^4 \left( 9 \left( -5 \log \left( -\frac{k^2}{m^2} \right) + 7 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^2 \right) \tag{10.40}$$

$$+ \left( 15 \log \left( -\frac{k^2}{m^2} \right) - 31 \right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$

and

$$\mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = -\frac{im^4}{288\pi^3} k^2 \left( \left( 6\log\left(-\frac{k^2}{m^2}\right) + 5\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^2 + \left(-6\log\left(-\frac{k^2}{m^2}\right) + 7\right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$
(10.41)

$$\mathcal{O}_{\rm UV}(m^6) - \mathcal{O}_{\rm IR}(m^6) = -\frac{im^6}{144\pi^3} \left( \left( 6\log\left(-\frac{k^2}{m^2}\right) - 13\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^2 \quad (10.42) + \left( 6\log\left(-\frac{k^2}{m^2}\right) - 1\right) \left(n_1 \cdot \pi^{(k)} \cdot n_1\right) \left(n_2 \cdot \pi^{(k)} \cdot n_2\right) \right) \\ \mathcal{O}_{\rm UV}(m^8) = \frac{im^8}{576\pi^3 k^2} \left( 3\left(12\log\left(-\frac{k^2}{m^2}\right) + 17\right) \left(n_1 \cdot \pi^{(k)} \cdot n_2\right)^2 \quad (10.43) + \left(12\log\left(-\frac{k^2}{m^2}\right) + 25\right) \left(n_1 \cdot \pi^{(k)} \cdot n_1\right) \left(n_2 \cdot \pi^{(k)} \cdot n_2\right) \right)$$

They are all conserved. (10.41) contains a nonlocal linearized version of the EH eom.

### 10.2.3 6d mfm: spin 3 current

The scheme is the same as above. In the IR the odd power of m vanish. The even powers  $m^{2n}$  with  $n \leq 2$  are conserved together with the term proportional to  $\log(m)$ . The terms  $\mathcal{O}_{\mathrm{IR}}(m^4), \mathcal{O}_{\mathrm{IR}}(m^6), \mathcal{O}_{\mathrm{IR}}(m^8)$  are not conserved. Of course  $\mathcal{O}(\log[m]), \mathcal{O}_{\mathrm{IR}}(m^2)$  diverge in the IR, while the term  $\mathcal{O}_{\mathrm{IR}}(m^0)$  diverges for  $\varepsilon \to 0$ . According to our prescription all these terms, which are local, have to be subtracted from the effective action.

Using again  $\Pi^{(3)}(k, n_1, n_2)$ , see (5.10), the result is

$$\mathcal{O}(m^{-2}): -\frac{ik^{10}}{93555m^2\pi^3} \prod_{\frac{49}{128}}^{(3)}(k, n_1, n_2)$$
(10.44)

In the UV the odd m powers vanish. The even powers of order 4, 6, 8 are not conserved. After subtracting the local terms in the IR one gets

$$\mathcal{O}_{\rm UV}(m^0) - \mathcal{O}_{\rm IR}(m^0) - \mathcal{O}_{\rm IR}(\log(m)) \tag{10.45}$$

$$= -\frac{ik^8}{57153600\pi^3} \left( 32 \left( -315 \log\left(-\frac{k^2}{m^2}\right) + 1021 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^3 -3 \left( -630 \log\left(-\frac{k^2}{m^2}\right) + 3617 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$

and

$$\mathcal{O}_{\rm UV}(m^2) - \mathcal{O}_{\rm IR}(m^2) = \frac{im^2k^6}{3175200\pi^3} \bigg( 32 \bigg( -210 \log\bigg( -\frac{k^2}{m^2} \bigg) + 389 \bigg) \bigg( n_1 \cdot \pi^{(k)} \cdot n_2 \bigg)^3 \quad (10.46) \\ + \bigg( 3885 \log\bigg( -\frac{k^2}{m^2} \bigg) - 9454 \bigg) \bigg( n_1 \cdot \pi^{(k)} \cdot n_2 \bigg) \big( n_1 \cdot \pi^{(k)} \cdot n_1 \big) (n_2 \cdot \pi^{(k)} \cdot n_2 \big) \bigg) \\ \mathcal{O}_{\rm UV}(m^4) - \mathcal{O}_{\rm IR}(m^4) = \frac{im^4k^4}{64800\pi^3} \bigg( 16 \bigg( -17 + 30 \log\bigg( -\frac{k^2}{m^2} \bigg) \bigg) \bigg( n_1 \cdot \pi^{(k)} \cdot n_2 \bigg)^3 \quad (10.47) \\ + \bigg( -480 \log\bigg( -\frac{k^2}{m^2} \bigg) + 617 \bigg) \bigg( n_1 \cdot \pi^{(k)} \cdot n_2 \bigg) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \bigg) \bigg)$$

$$\mathcal{O}_{\rm UV}(m^6) - \mathcal{O}_{\rm IR}(m^6) = -\frac{im^6k^2}{864\pi^3} \left( \frac{64}{3} \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right)^3 \right)$$
(10.48)  
$$- \left( 18 \log \left( -\frac{k^2}{m^2} \right) - 17 \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$
(10.49)  
$$\mathcal{O}_{\rm UV}(m^8) - \mathcal{O}_{\rm IR}(m^8) = \frac{im^8}{5184\pi^3} \left( 16 \left( (11 - 12 \log \left( -\frac{k^2}{m^2} \right) \right) \right)$$
(10.49)  
$$+ \left( 211 - 276 \log \left( -\frac{k^2}{m^2} \right) \right) \left( n_1 \cdot \pi^{(k)} \cdot n_2 \right) (n_1 \cdot \pi^{(k)} \cdot n_1) (n_2 \cdot \pi^{(k)} \cdot n_2) \right)$$
(10.49)  
$$\cdots \cdots$$

are all conserved. Eq. (10.48) is related to a nonlocal version of the spin 3 Fronsdal equation.

#### 11 Spin s current two-point correlators in any dimension

In this section we derive general formulas for the two-point correlators for spin s = 1, 2 and 3 in any dimension. The procedure is slightly different from the one used so far. In the previous sections we fixed the dimension of space-time, that is we set  $d = 3, 4 + \varepsilon, 5, 6 + \varepsilon$  in the scalar integrals (see section 5). In this section we leave the parameter d free and evaluate specific cases at the end. The two procedures often lead to different intermediate results. Of course the results of physical interest must coincide.

In the following we focus on the massive case, while the massless computations are deferred to appendix B.

#### 11.1 Fermion model

In this subsection we compute the even part of two point correlator for a fermion in d dimensions for spin s = 1, 2, 3

$$\tilde{J}_{\mu_1\dots\mu_s\nu_1\dots\nu_s}(k) = -\int \frac{d^3p}{(2\pi)^3} \operatorname{Tr}\left(\frac{i}{\not\!\!p-m}\gamma_\sigma \frac{i}{\not\!\!p-\not\!\!k-m}\gamma_\tau\right) V^{\sigma}_{\mu_1\dots\mu_s} V^{\tau}_{\nu_1\dots\nu_s}$$
(11.1)

where the Feynman vertices are

$$V_{\mu_1\dots\mu_s}^{\sigma} = i \quad \delta_{\mu}^{\sigma} \sum_{j=0}^{\lfloor \frac{s-1}{2} \rfloor} \frac{i^{2s-2j-2}(k_{\mu}-2p_{\mu})^{s-2j-1}(2(p\cdot(p-k)-m^2)\eta_{\mu\mu}-4(p-k)_{\mu}p_{\mu})^j}{(2j+1)!(n-2j-1)!}$$
(11.2)

#### 11.1.1 Fermion model — massive case

Let us compute the even part of two point correlator for a massive fermion in d dimensions for spin 1. The one-loop contribution is

$$\tilde{J}_{\mu\nu}(k) = -\int \frac{d^d p}{(2\pi)^d} \operatorname{Tr}\left(\gamma_{\nu} \frac{1}{\not{p} - m} \gamma_{\mu} \frac{1}{\not{p} - \not{k} - m}\right)$$
$$= -\operatorname{Tr}(1) \int \frac{d^d p}{(2\pi)^d} \frac{p_{\nu}(p-k)_{\mu} + p_{\mu}(p-k)_{\nu} - p \cdot (p-k)\eta_{\mu\nu} + m^2 \eta_{\mu\nu}}{(p^2 - m^2)((p-k)^2 - m^2)} \quad (11.3)$$

Tr(1) is the trace of the identity operator on the vector space on which the Dirac matrices act. Since we are working with the lowest dimensional complex spinors in each dimension we have  $\text{Tr}(1) = 2^{\lfloor \frac{d}{2} \rfloor}$ . In the odd dimensional cases d can be simply replaced by the values  $3, 5, \ldots$ , so this factor is an overall numerical factor. In the even dimensional cases we have to replace d by  $2 + \varepsilon, 4 + \varepsilon, \ldots$ , so the same factor contains an  $\varepsilon$  dependence in addition to the overall numerical factor. The  $\varepsilon$  dependence cannot change the divergent pole part in dimensional regularization but will only affect the finite local part. However when subtracting the infinite and finite IR terms from the effective action this dependence disappears. So we will ignore it.

**Warning.** In evaluating the scalar integral  $\tilde{\mathcal{I}}^{(1)}(k)$ , which is our basic quantity, the results below have been obtained by choosing a reference value 4 for Tr(1). This value is appropriate only for d = 4, 5, but must be corrected for the other dimensions: for d = 2, 3 the results must be divided by 2, for d = 6, 7 they must be multiplied by 2, and so on.

By Davydychev tensor reduction procedure it is possible to rewrite such an amplitude in terms of scalar integrals as

$$\tilde{J}_{\mu\nu}(k) = \eta_{\mu\nu}\tilde{\mathcal{I}}^{(1)}(k) + k_{\mu}k_{\nu}\tilde{\mathcal{I}}^{(2)}(k) = -4\left(\eta_{\mu\nu}\left(m^{2}\tilde{I}^{(2)}(d, 1, 1; k, m)\right)\right)$$
(11.4)

$$-2\pi \left( 2k \left( I^{(2)}(d+2,2,1;k,m) + 8\pi I^{(2)}(d+4,3,1;k,m) \right) - (d-2)\tilde{I}^{(2)}(d+2,1,1;k,m) \right) \right)$$
  
+8\pi k\_\mu k\_\nu \left( \tilde{I}^{(2)}(d+2,2,1;k,m) + 8\pi I^{(2)}(d+4,3,1;k,m) \right) \right)

Because of dimensional reasons the superficial degree of divergence of  $\tilde{\mathcal{I}}^{(1)}$  and  $\tilde{\mathcal{I}}^{(2)}$  are always such that  $\deg(\tilde{\mathcal{I}}^{(1)}) = \deg(\tilde{\mathcal{I}}^{(2)}) + 2 = d - 2$ . One can check the Ward identity

$$k^{2}\tilde{\mathcal{I}}^{(2)}(k) + \tilde{\mathcal{I}}^{(1)}(k) = 0, \qquad (11.5)$$

which implies we can rewrite the amplitude (11.3) as

$$\tilde{J}_{\mu\nu}(k) = \left(\eta_{\mu\nu} - \frac{k_{\mu}k_{\nu}}{k^2}\right)\tilde{\mathcal{I}}^{(1)}(k),$$
(11.6)

This is in fact the main advantage of the regularization we are using in this section: the two point function is conserved without any subtraction. This will not be true for higher spin currents (see below).

The explicit form of  $\tilde{\mathcal{I}}^{(1)}(k)$  is

$$\tilde{\mathcal{I}}^{(1)}(k) = -\frac{2^{1-d} i e^{-\frac{1}{2}i\pi d} \pi^{-\frac{d}{2}} (-m^2)^{d/2-1} \Gamma\left(2-\frac{d}{2}\right)}{(d-2)}$$

$$\cdot \left(8\frac{k^2}{4m^2} {}_2F_1\left(2,1-\frac{d}{2};\frac{3}{2};\frac{k^2}{4m^2}\right) - 4\left(d\frac{k^2}{4m^2}+1\right) {}_2F_1\left(1,1-\frac{d}{2};\frac{3}{2};\frac{k^2}{4m^2}\right) + 4\right) \\
= \frac{2^{3-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{\frac{d-2}{2}} \Gamma\left(2-\frac{d}{2}\right) \left(\left((d-2)z+1\right) {}_2F_1\left(1,-\frac{d}{2};\frac{1}{2};z\right) + 2z-1\right)}{(d-2)d(z-1)z}$$
(11.7)

where  $z = \frac{k^2}{4m^2}$ . For d = 2 there is no pole. For even d > 2 the relevant  $\varepsilon$ -expansion is given by

$$\tilde{\mathcal{I}}^{(1)}(k) = \frac{1}{\varepsilon} \frac{2^{4-d} i^{d+1} \pi^{-d/2} m^{d-2} \left( \left( (d-2)z+1 \right) {}_2F_1 \left( 1, -\frac{d}{2}; \frac{1}{2}; z \right) + 2z - 1 \right)}{(d-2)d(z-1)z\Gamma\left(\frac{d}{2} - 1\right)}$$
(11.8)  
$$2^{3-d} i^{d+1} e^{-d/2} m^{d-2}$$

$$-\frac{2^{o} - a_{1}a_{1} + \pi - a_{2} - 2ma^{a-2}}{(d-2)^{2}d^{2}(z-1)z\Gamma\left(\frac{d}{2}-1\right)}$$

$$\cdot \left(-2(d-2)d\left(z_{2}F_{1}\left(1,-\frac{d}{2};\frac{1}{2};z\right)-\frac{1}{2}((d-2)z+1)_{2}F_{1}^{(0,1,0,0)}\left(1,-\frac{d}{2},\frac{1}{2},z\right)\right)\right)$$

$$+ \left(((d-2)z+1)_{2}F_{1}\left(1,-\frac{d}{2};\frac{1}{2};z\right)+2z-1\right)$$

$$\cdot \left(d^{2}\log(4\pi)-2(d-2)d\log(m)-2d(\log(4\pi)-2)-4\right)$$

$$+ (d-2)d\left(H_{\frac{d}{2}-2}-\gamma\right)\left(((d-2)z+1)_{2}F_{1}\left(1,-\frac{d}{2};\frac{1}{2};z\right)+2z-1\right)\right)$$
(11.9)

where  $_2F_1^{(0,1,0,0)}\left(1,-\frac{d}{2},\frac{1}{2},z\right)$  means the derivative of  $_2F_1^{(0,1,0,0)}\left(1,\beta,\frac{1}{2},z\right)$  with respect to  $\beta$  at  $\beta = -\frac{d}{2}$ , and  $H_{\frac{d}{2}-2}$  are the harmonic numbers.

## 11.1.2 d even

Let us examine first the even d case. For even d the hypergeometric functions boil down to finite order polynomials according to the formulas

$${}_{2}F_{1}\left(2,1-\frac{d}{2};\frac{3}{2};z\right) = \sum_{n=0}^{\frac{d}{2}-1} \frac{(2)_{n}\left(1-\frac{d}{2}\right)_{n}}{n!\left(\frac{3}{2}\right)_{n}} z^{n}$$
$${}_{2}F_{1}\left(1,1-\frac{d}{2};\frac{3}{2};z\right) = \sum_{n=0}^{\frac{d}{2}-1} \frac{(1)_{n}\left(1-\frac{d}{2}\right)_{n}}{n!\left(\frac{3}{2}\right)_{n}} z^{n}$$

so that one can easily check the  $1/\varepsilon$  part is just

$$d = 4 \qquad \frac{2im^2 z}{3\pi^2}, \tag{11.10}$$

$$d = 6 \qquad \frac{im^4 z (4z - 5)}{30\pi^3}, \tag{11.10}$$

$$d = 8 \qquad \frac{im^6 z (24z^2 - 56z + 35)}{1680\pi^4}, \qquad$$

$$d = 10 \qquad \frac{im^8 z (64z^3 - 216z^2 + 252z - 105)}{60480\pi^5}, \qquad$$

$$d = 12 \qquad \frac{im^{10} z (640z^4 - 2816z^3 + 4752z^2 - 3696z + 1155)}{10644480\pi^6}, \qquad$$
...

So in general the divergent part (for  $\varepsilon \to 0$ ) is a polynomial in  $z = \frac{k^2}{4m^2}$  of degree d/2 - 1, where the constant term is missing because the front factor of highest dimension is  $m^{d-2}$ .

According to Weinberg's theorem this corresponds to the degree of divergence d-4 of the two point functions, which is therefore lower than the expected one d-2 because of gauge invariance. The above divergent terms are local and appear both in the IR and the UV limit, as we have seen many times above. Some of them are divergent for  $m \to \infty$  and must be subtracted to guarantee finiteness of the IR limit (or decoupling of infinite mass modes). Others are of order  $m^0$ . The reason why we subtract them from the effective action is, according to our attitude, because the physical information is contained in the difference between the IR and UV limits (rather than in their absolute value).

As for the finite part we cannot give a closed formula for generic d, but thanks to the formulae

$${}_{2}F_{1}\left(1, -\frac{d}{2}; \frac{1}{2}; z\right) = \sum_{n=0}^{d/2} \frac{(1)_{n} \left(-\frac{d}{2}\right)_{n} z^{n}}{\left(\frac{1}{2}\right)_{n}} \frac{z^{n}}{n!}$$

$$= \sum_{n=0}^{d/2} \frac{\Gamma(-\frac{d}{2}+n)}{\Gamma(-\frac{d}{2})} \frac{\Gamma(\frac{1}{2})}{\Gamma(\frac{1}{2}+n)} z^{n}$$

$${}_{2}F_{1}^{(0,1,0,0)}\left(1, -\frac{d}{2}; \frac{1}{2}; z\right) = -2\sum_{n=0}^{\infty} \frac{\partial}{\partial d} \frac{\Gamma(-\frac{d}{2}+n)}{\Gamma(-\frac{d}{2})} \frac{\Gamma(\frac{1}{2})}{\Gamma(\frac{1}{2}+n)} z^{n}$$

$$= -\sum_{n=0}^{\infty} \frac{\sqrt{\pi}\Gamma\left(n-\frac{d}{2}\right)\left(\psi^{(0)}\left(-\frac{d}{2}\right)-\psi^{(0)}\left(n-\frac{d}{2}\right)\right)}{\Gamma\left(-\frac{d}{2}\right)\Gamma\left(n+\frac{1}{2}\right)} z^{n}$$

$$(11.11)$$

we can recognize the IR behavior is analytic in z, which is to be expected as m acts as an IR regulator. More explicitly, we get the following behaviors

$$\begin{split} d &= 4: \frac{im^2 z (\log(\frac{m^2}{4\pi}) + \gamma)}{3\pi^2} - \frac{4im^2 z^2}{15\pi^2} - \frac{4im^2 z^3}{35\pi^2} - \frac{64im^2 z^4}{945\pi^2} + \mathcal{O}(z^5), \quad (11.12) \\ d &= 6: \frac{im^4 z (-\log(\frac{m^2}{4\pi}) - \gamma + 1)}{12\pi^3} + \frac{im^4 z^2 (\log(\frac{m^2}{4\pi}) + \gamma)}{15\pi^3} \\ &- \frac{im^4 z^3}{35\pi^3} - \frac{8im^4 z^4}{945\pi^3} + \mathcal{O}(z^5), \\ d &= 8: \frac{im^6 z (2\log(\frac{m^2}{4\pi}) + 2\gamma - 3)}{192\pi^4} + \frac{im^6 z^2 (-\log(\frac{m^2}{4\pi}) - \gamma + 1)}{60\pi^4} \\ &+ \frac{im^6 z^3 (\log(\frac{m^2}{4\pi}) + \gamma)}{140\pi^4} - \frac{2im^6 z^4}{945\pi^4} + \mathcal{O}(z^5), \\ d &= 10: \frac{im^8 z (-6\log(\frac{m^2}{4\pi}) - 6\gamma + 11)}{6912\pi^5} + \frac{im^8 z^2 (2\log(\frac{m^2}{4\pi}) + 2\gamma - 3)}{960\pi^5} \\ &+ \frac{im^8 z^3 (-\log(\frac{m^2}{4\pi}) - \gamma + 1)}{560\pi^5} + \frac{im^8 z^4 (\log(\frac{m^2}{4\pi}) + \gamma)}{1890\pi^5} + \mathcal{O}(z^5), \end{split}$$

So the dominating term is  $\sim m^{d-2}(A^{(1)}\log m + B^{(1)})z \equiv m^{d-4}(A^{(1)}\log m + B^{(1)})k^2$ , whereas the term with highest power of momentum and dimensionless constant is  $\sim m^{d-2}(A^{(d/2-1)}\log m + B^{(d/2-1)})z^{d/2-1} \equiv (A^{(d/2-1)}\log m + B^{(d/2-1)})k^{d-2}$ . So, in coordinate space the following terms are dominating for large m

$$\sim (A^{(d/2-1)}\log m + B^{(d/2-1)})F^{\mu\nu} \Box^{d/2-1}F_{\mu\nu}, \dots, m^{d-4}(A^{(1)}\log m + B^{(1)})F^{\mu\nu}F_{\mu\nu},$$

whereas all the others are suppressed by negative powers of m. For d > 4 those terms would give a non-decoupling of IR dynamics from high-energy physics, but we can notice they are the same as the local counterterms appearing in the divergent part.<sup>7</sup> So they have to be subtracted, as we have done many times before. In (11.12) there are also terms of order  $m^0$ . They have to be subtracted from the effective action for the same reason explained above: the physical meaning is contained in the difference between the UV and the IR.

In d = 2 no pole shows up. In fact we have

$$\frac{i\left({}_{2}F_{1}^{(0,1,0,0)}\left(1,-1,\frac{1}{2},z\right)+2z(2z-1)\right)}{2\pi(z-1)z},$$

whose IR expansion is

$$-\frac{4iz}{3\pi} - \frac{16iz^2}{15\pi} - \frac{32iz^3}{35\pi} - \frac{256iz^4}{315\pi} - \frac{512iz^5}{693\pi} - \frac{2048iz^6}{3003\pi} + O\left(z^7\right), \qquad (11.13)$$

meaning all the local terms are suppressed by negative powers of the mass. There is no need to remove them by finite subtraction in order to have decoupling.

The asymptotic behavior of the finite part in the UV  $(z \to \infty)$  is determined by the formulae

$${}_{2}F_{1}\left(1,-\frac{d}{2};\frac{1}{2};z\right) = \frac{(1)_{\frac{d}{2}}(-z)^{\frac{d}{2}}}{\left(\frac{1}{2}\right)_{\frac{d}{2}}} {}_{2}F_{1}\left(-\frac{d}{2},\frac{1-d}{2};-\frac{d}{2};\frac{1}{z}\right)$$

$$= \Gamma\left(\frac{1}{2}\right) \frac{\Gamma\left(\frac{d+2}{2}\right)}{\Gamma\left(\frac{d+1}{2}\right)} (-z)^{\frac{d}{2}} {}_{2}F_{1}\left(-\frac{d}{2},\frac{1-d}{2};-\frac{d}{2};\frac{1}{z}\right)$$

$$= (-z)^{d/2} \sum_{n=0}^{\frac{d}{2}} \frac{\left(\cos\left(\frac{\pi d}{2}\right)\Gamma\left(\frac{d}{2}+1\right)\Gamma\left(-\frac{d}{2}+n+\frac{1}{2}\right)\right)}{\sqrt{\pi}\Gamma(n+1)} \left(\frac{1}{z}\right)^{n}$$

$$(11.14)$$

and

$${}_{2}F_{1}^{(0,1,0,0)}\left(1,-\frac{d}{2};\frac{1}{2};z\right) = -2\frac{\partial}{\partial d} {}_{2}F_{1}\left(1,-\frac{d}{2};\frac{1}{2};z\right)$$
(11.15)  
$$= -z^{d/2}\log(-z)\sum_{n=0}^{\frac{d}{2}}\frac{\Gamma\left(\frac{d}{2}+1\right)\Gamma\left(-\frac{d}{2}+n+\frac{1}{2}\right)}{\sqrt{\pi}\Gamma(n+1)}\left(\frac{1}{z}\right)^{n}$$
$$+z^{d/2}\sum_{n=0}^{\infty}\frac{\Gamma\left(\frac{d}{2}+1\right)\Gamma\left(-\frac{d}{2}+n+\frac{1}{2}\right)\left(H_{-\frac{d}{2}+n-\frac{1}{2}}-H_{\frac{d}{2}}\right)}{\sqrt{\pi}\Gamma(n+1)}\left(\frac{1}{z}\right)^{n}$$

<sup>&</sup>lt;sup>7</sup>This case corresponds, in ordinary interacting gauge theories, to the fact that the vertex for the spin 1 current is a non-renormalizable interaction for d > 4 and in fact higher derivatives operators are generated. d = 4 is the special case when just the two-derivative operator with dimensionless front constant is generated and it corresponds to the fact that spin 1 vertex in d = 4 is a renormalizable interaction. For d < 4 the spin 1 vertex is super-renormalizable.

We notice, as before, that the asymptotic expansion of  ${}_{2}F_{1}^{(0,1,0,0)}\left(1,-\frac{d}{2};\frac{1}{2};z\right)$  contains  $\log(-z)$ , which can have an imaginary part even for real -z. We should keep in mind that the expansion is valid under the assumption  $|\arg(-z)| < \pi$ , which means -z, if real, must be assumed to be positive. This is for example the case for Euclidean momenta which are such that  $\frac{k_{E}^{2}}{4m^{2}} = -\frac{k_{M}^{2}}{4m^{2}} = -z$ . When the metric is Lorentzian a cut appears at z = 1 and we have to rely on the analytic continuation by choosing a definite Riemann sheet: it is not surprising that in this case an imaginary part of the correlator appear. Dimension by dimension the UV expansions for the finite part are:

$$\begin{split} d &= 4: \quad \frac{im^2 z(3\log(\frac{m^2}{\pi}) + 3\log(-z) + 3\gamma - 5)}{9\pi^2} - \frac{im^2(2\log(-4z) + 1)}{16\pi^2 z} \quad (11.16) \\ &- \frac{im^2}{2\pi^2} + O\left(\left(\frac{1}{z}\right)^2\right), \\ d &= 6: \quad \frac{im^4 z^2(30\log(\frac{m^2}{\pi}) + 30\log(-z) + 30\gamma - 77)}{450\pi^3} + \frac{im^4(6\log(-4z) + 5)}{576\pi^3 z} \\ &- \frac{im^4 z(3\log(\frac{m^2}{\pi})) + 3\log(-z) + 3\gamma - 5)}{36\pi^3} + \frac{im^4}{16\pi^3} + O\left(\left(\frac{1}{z}\right)^2\right), \\ d &= 8: \quad \frac{im^6 z^3(105\log(\frac{m^2}{\pi}) + 105\log(-z) + 105\gamma - 317)}{14700\pi^4} \\ &- \frac{im^6 z^2(30\log(\frac{m^2}{\pi}) + 3\log(-z) + 3\gamma - 5)}{288\pi^4} - \frac{im^6(12\log(-4z) + 13)}{18432\pi^4 z} \\ &- \frac{im^6}{192\pi^4} + O\left(\left(\frac{1}{z}\right)^2\right), \\ d &= 10: \quad \frac{im^8 z^4(1260\log(\frac{m^2}{\pi}) + 1260\log(-z) + 1260\gamma - 4189)}{2381400\pi^5} \\ &- \frac{im^8 z^3(105\log(\frac{m^2}{\pi}) + 1260\log(-z) + 105\gamma - 317)}{5880\pi^5} \\ &+ \frac{im^8 z^2(30\log(\frac{m^2}{\pi}) + 3\log(-z) + 30\gamma - 77)}{14400\pi^5} \\ &- \frac{im^8 z(3\log(\frac{m^2}{\pi}) + 3\log(-z) + 30\gamma - 77)}{3456\pi^5} + \frac{im^8(60\log(-4z) + 77)}{1843200\pi^5 z} + O\left(\left(\frac{1}{z}\right)^2\right), \\ \end{array}$$

As it is to be expected the leading behavior is  $\sim m^{d-2}z^{d/2-1}\log(-z)$  corresponding to the UV behavior of the divergent part. The presence of logarithms is to be interpreted as the consequence of running of parameters. In ordinary (interacting) gauge theories, once these logarithms are reabsorbed in the running parameters, the remaining polynomial behavior can itself be subtracted by proper counterterms leading to a well-behaved amplitude in the UV.

## 11.1.3 d odd

Let us discuss now the odd dimensional case. In odd dimensions there is no divergent part in the  $\epsilon$ -expansion. For the IR we get the expansion

$$d = 3 \quad \sqrt{-m^2} \left( -\frac{2z}{3\pi} - \frac{4z^2}{15\pi} - \frac{6z^3}{35\pi} - \frac{8z^4}{63\pi} - \frac{10z^5}{99\pi} + O(z^6) \right),$$

$$d = 5 \quad (-m^2)^{3/2} \left( -\frac{z}{3\pi^2} + \frac{2z^2}{15\pi^2} + \frac{z^3}{35\pi^2} + \frac{4z^4}{315\pi^2} + \frac{5z^5}{693\pi^2} + O(z^6) \right),$$

$$d = 7 \quad (-m^2)^{5/2} \left( -\frac{z}{18\pi^3} + \frac{z^2}{15\pi^3} - \frac{z^3}{70\pi^3} - \frac{2z^4}{945\pi^3} - \frac{z^5}{1386\pi^3} + O(z^6) \right),$$

$$d = 9 \quad (-m^2)^{7/2} \left( -\frac{z}{180\pi^4} + \frac{z^2}{90\pi^4} - \frac{z^3}{140\pi^4} + \frac{z^4}{945\pi^4} + \frac{z^5}{8316\pi^4} + O(z^6) \right),$$

$$d = 11 \quad (-m^2)^{9/2} \left( -\frac{z}{2520\pi^5} + \frac{z^2}{900\pi^5} - \frac{z^3}{840\pi^5} + \frac{z^4}{1890\pi^5} - \frac{z^5}{16632\pi^5} + O(z^6) \right),$$
(11.17)

Again the dominating term is  $\sim m^{d-2}z^2$ . The case d = 3 is the one in which no term with positive power of m shows up.

For the UV behavior we get

$$\begin{aligned} d &= 3 \quad \sqrt{-m^2} \left( \frac{i\sqrt{z}}{4} + \frac{1}{4}i\sqrt{\frac{1}{z}} + O\left(\left(\frac{1}{z}\right)^{3/2}\right) \right), \\ d &= 5 \quad \left(-m^2\right)^{3/2} \left( -\frac{3iz^{3/2}}{32\pi} + \frac{i\sqrt{z}}{16\pi} + \frac{i\sqrt{\frac{1}{z}}}{32\pi} - \frac{1}{15\pi^2 z} + O\left(\left(\frac{1}{z}\right)^{3/2}\right) \right), \\ d &= 7 \quad \left(-m^2\right)^{5/2} \left( \frac{5iz^{5/2}}{384\pi^2} - \frac{3iz^{3/2}}{128\pi^2} + \frac{i\sqrt{z}}{128\pi^2} + \frac{i\sqrt{\frac{1}{z}}}{384\pi^2} - \frac{1}{210\pi^3 z} + O\left(\left(\frac{1}{z}\right)^{13/2}\right) \right), \\ d &= 9 \quad \left(-m^2\right)^{7/2} \left( -\frac{7iz^{7/2}}{6144\pi^3} + \frac{5iz^{5/2}}{1536\pi^3} - \frac{3iz^{3/2}}{1024\pi^3} + \frac{i\sqrt{z}}{1536\pi^3} + \frac{i\sqrt{\frac{1}{z}}}{6144\pi^3} \right) \\ &- \frac{1}{3780\pi^4 z} + O\left(\left(\frac{1}{z}\right)^{13/2}\right) \right), \\ d &= 11 \quad \left(-m^2\right)^{9/2} \left( \frac{3iz^{9/2}}{40960\pi^4} - \frac{7iz^{7/2}}{24576\pi^4} + \frac{5iz^{5/2}}{12288\pi^4} - \frac{iz^{3/2}}{4096\pi^4} + \frac{i\sqrt{z}}{24576\pi^4} \right) \\ &+ \frac{i\sqrt{\frac{1}{z}}}{122880\pi^4} - \frac{1}{83160\pi^5 z} + O\left(\left(\frac{1}{z}\right)^{13/2}\right) \right), \end{aligned}$$

$$(11.18)$$

# 11.1.4 Spin 2

For spin s = 2 the two-point correlator in any dimension d is

$$\begin{split} \tilde{J}_{\mu\mu\nu\nu}(k) &= \frac{2^{2-d+\left\lfloor \frac{d}{2} \right\rfloor} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} \Gamma\left(1-\frac{d}{2}\right)}{d(d+1)k^2} \\ &\times \left[ \left( -\left((d-1)k^2 + 8m^2\right) \,_2F_1\left[1, -\frac{d}{2}; \frac{1}{2}; \frac{k^2}{4m^2}\right] - (d+1)k^2 + 8m^2 \right) \pi^2_{\mu\nu} \right. \\ &\left. - \left( \left(4m^2 - k^2\right) \,_2F_1\left[1, -\frac{d}{2}; \frac{1}{2}; \frac{k^2}{4m^2}\right] + (d+1)k^2 - 4m^2 \right) \pi_{\mu\mu}\pi_{\nu\nu} \right] \\ &\left. + \frac{2^{2-d+\left\lfloor \frac{d}{2} \right\rfloor} i \,\pi^{-d/2}m^d\Gamma\left(1-\frac{d}{2}\right) \left(\eta^2_{\mu\nu} + \eta_{\mu\mu}\eta_{\nu\nu}\right)}{d} \right] \end{split}$$

We can check the Ward identity

$$k^{\mu}\tilde{J}_{\mu\mu\nu\nu}(k) = -2^{2-d+\left\lfloor\frac{d}{2}\right\rfloor} i \pi^{-d/2} m^{d} \Gamma\left(-\frac{d}{2}\right) \left(k_{\nu}\eta_{\mu\nu} + k_{\mu}\eta_{\nu\nu}\right)$$
(11.19)

For spin s = 2 the two-point correlator is transverse up to local counterterms.

# 11.1.5 Spin 3

For spin s = 3 the two-point correlator in any dimension d can be written as

$$\begin{split} \tilde{J}_{\mu\mu\mu\nu\nu\nu\nu} &= -\frac{2^{1-d+\left\lfloor \frac{d}{2} \right\rfloor} i \, e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} (d+2) \Gamma \left(2-\frac{d}{2}\right)}{9d(d+3) \left(d^4-5d^2+4\right) k^4} \pi_{\mu\nu}} \\ &\times \left[ 32 \left( -\left(k^2-4m^2\right)^2 \left(dk^2+12m^2\right) \, _2F_1 \left[1,-\frac{d}{2};-\frac{1}{2};\frac{k^2}{4m^2}\right] \right. \\ &\left. +4m^2 \left(48m^4+8(2d-3)k^2m^2+\left(3-2d(d+1)\right)k^4\right)\right) \pi_{\mu\nu}^2 \\ &- \left(4k^4m^2 \left( \left(d(9d-44\right)+219\right) \, _2F_1 \left[1,-\frac{d}{2};-\frac{1}{2};\frac{k^2}{4m^2}\right] + d(95d+68) - 219\right) \right. \\ &\left. +k^6 \left( \left(d((d-6)d+3)-54\right) \, _2F_1 \left[1,-\frac{d}{2};-\frac{1}{2};\frac{k^2}{4m^2}\right] - 18(d-1)(d+1)(d+3)\right) \right. \\ &\left. +9216m^6 \left( \, _2F_1 \left[1,-\frac{d}{2};-\frac{1}{2};\frac{k^2}{4m^2}\right] - 1\right) \right. \\ &\left. +128k^2m^4 \left( \left(5d-39\right) \, _2F_1 \left[1,-\frac{d}{2};-\frac{1}{2};\frac{k^2}{4m^2}\right] - 23d+39\right) \right) \pi_{\mu\mu}\pi_{\nu\nu} \right] \\ &\left. -\frac{2^{1-d+\left\lfloor \frac{d}{2} \right]} i \, e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} \Gamma \left(2-\frac{d}{2}\right)}{9d(d^2-4)} \right. \\ &\times \left[ \eta_{\mu\nu} \left[ \eta_{\nu\nu} \left(2\eta_{\mu\mu} \left(208m^2-9(d+2)k^2\right) + 21(d+2)k_{\mu}k_{\mu}\right) + 256m^2\eta_{\mu\nu}\eta_{\mu\nu}} \right] \\ &\left. +21(d+2)\eta_{\mu\mu}\eta_{\mu\nu}k_{\nu}k_{\nu} + 18(d+2)\eta_{\mu\mu}\eta_{\nu\nu}k_{\mu}k_{\nu} \right] \end{split}$$

For spin s = 3 the two-point correlator is conserved up to local counterterms

$$k^{\mu}\tilde{J}_{\mu\mu\mu\nu\nu\nu\nu} = -\frac{2^{1-d+\left\lfloor\frac{d}{2}\right\rfloor}ie^{-\frac{1}{2}i\pi d}\pi^{-d/2}\left(-m^{2}\right)^{d/2}\Gamma\left(2-\frac{d}{2}\right)}{9d\left(d^{2}-4\right)} \times \left[k_{\nu}\left(\eta_{\nu\nu}\left(57(d+2)k_{\mu}k_{\mu}+416m^{2}\eta_{\mu\mu}\right)+768m^{2}\eta_{\mu\nu}\eta_{\mu\nu}\right)\right. \\ \left.+2\eta_{\mu\nu}\eta_{\nu\nu}k_{\mu}\left(3(d+2)k^{2}+416m^{2}\right)+21(d+2)\eta_{\mu\mu}k_{\nu}k_{\nu}k_{\nu}\right. \\ \left.+42(d+2)\eta_{\mu\nu}k_{\nu}k_{\nu}\right]$$
(11.21)

### 11.2 Scalar model

Let us compute the even part of two point correlator for a scalar in d dimensions for spin s

$$\tilde{J}_{\mu_1\dots\mu_s\nu_1\dots\nu_s}(k) = \int \frac{d^d p}{(2\pi)^d} \frac{i}{p^2 - m^2} V_{\mu_1\dots\mu_s} \frac{i}{(p-k)^2 - m^2} V_{\nu_1\dots\nu_s}$$
(11.22)

where the vertex for an incoming scalar with momentum p and outgoing scalar with momentum p' and an outgoing spin-s field with momentum k is

$$V_{\mu_1\dots\mu_s} = i(p+p')_{\mu_1}\dots(p+p')_{\mu_s}\delta^{(d)}(p-p'-k)$$
(11.23)

## 11.2.1 Scalar model — massive case

Let us compute the two point correlator for the massive scalar in any dimension d for spin s=1

$$\tilde{J}_{\mu\nu}(k) = \frac{2^{2-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} \Gamma\left(2-\frac{d}{2}\right)}{(d-2)k^2 m^2} \times \left(\left(k^2 \eta_{\mu\nu} - k_{\mu} k_{\nu}\right) \ _2F_1\left[1, 1-\frac{d}{2}; \frac{3}{2}; \frac{k^2}{4m^2}\right] + k_{\mu} k_{\nu}\right)$$
(11.24)

 $\tilde{J}_{\mu\nu}$  is not conserved

$$k^{\mu}\tilde{J}_{\mu\nu}(k) = \frac{2^{-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} d\Gamma\left(-\frac{d}{2}\right)}{m^2} k_{\nu}$$
(11.25)

To make  $\tilde{J}_{\mu\nu}(k)$  conserved let us add a local counterterm with an arbitrary constant a

$$\tilde{J}_{\mu\nu}(k) = \frac{2^{2-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} \Gamma\left(2-\frac{d}{2}\right)}{(d-2)k^2 m^2} \times \left(ak^2 \eta_{\mu\nu} + k_{\mu}k_{\nu} + \left(k^2 \eta_{\mu\nu} - k_{\mu}k_{\nu}\right) \,_2F_1\left[1, 1-\frac{d}{2}; \frac{3}{2}; \frac{k^2}{4m^2}\right]\right) \quad (11.26)$$

We get conservation for a = -1. The conserved 2pt is

$$\tilde{J}_{\mu\nu}(k) = \frac{2^{2-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} \Gamma\left(2-\frac{d}{2}\right)}{(d-2)m^2} \pi_{\mu\nu} \\ \times \left({}_2F_1\left[1, 1-\frac{d}{2}; \frac{3}{2}; \frac{k^2}{4m^2}\right] - 1\right)$$
(11.27)

For massive scalar for spin s = 2 we get

$$\tilde{J}_{\mu\mu\nu\nu}(k) = 2^{2-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} \Gamma\left(-\frac{d}{2}\right) \left(2\pi_{\mu\nu}^2 + \pi_{\mu\mu}\pi_{\nu\nu}\right) \\
\times \left({}_2F_1\left[1, -\frac{d}{2}; \frac{3}{2}; \frac{k^2}{4m^2}\right] - 1\right) \\
+ \frac{2^{-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} (-m^2)^{d/2} \Gamma\left(-\frac{d}{2}\right)}{k^2 m^2} \\
\times \left(d k_{\mu} k_{\mu} k_{\nu} k_{\nu} + 4k^2 m^2 \left(2\eta_{\mu\nu}\eta_{\mu\nu} + \eta_{\mu\mu}\eta_{\nu\nu}\right)\right) \tag{11.28}$$

 $\tilde{J}_{\mu\mu\nu\nu}(k)$  is not conserved

$$k^{\mu}\tilde{J}_{\mu\mu\nu\nu}(k) = \frac{2^{1-d} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} (-m^2)^{d/2} \Gamma\left(-\frac{d}{2}\right)}{m^2} \times \left(k_{\mu} \left(d k_{\nu} k_{\nu} + 4m^2 \eta_{\nu\nu}\right) + 8m^2 \eta_{\mu\nu} k_{\nu}\right)$$
(11.29)

For spin s = 3 we have

$$\begin{split} \tilde{J}_{\mu\mu\mu\nu\nu\nu\nu}(k) &= -3 \cdot 2^{1-d} \, i \, e^{-\frac{1}{2}i\pi d} \pi^{-d/2} \left(-m^2\right)^{d/2} \Gamma\left(-\frac{d}{2}-1\right) \pi_{\mu\nu} \left(2\pi_{\mu\nu}^2 + 3\pi_{\mu\mu}\pi_{\nu\nu}\right) \\ & \times \left(8m^2 \left(_2F_1\left[2, -\frac{d}{2}-1; \frac{3}{2}; \frac{k^2}{4m^2}\right] - 1\right) + (d+2)k^2 {}_2F_1\left[1, -\frac{d}{2}; \frac{3}{2}; \frac{k^2}{4m^2}\right]\right) \\ & -\frac{i2^{-d-1}e^{-\frac{1}{2}i\pi d} \pi^{-d/2} (-m^2)^{\frac{d}{2}} \Gamma\left(-\frac{d}{2}-1\right)}{k^4 m^2} \\ & \times \left[(d+2)k^2 \, k_\mu k_\nu (k_\mu k_\mu (d \, k_\nu k_\nu + 12m^2 \eta_{\nu\nu}) + 12m^2 \eta_{\mu\mu} k_\nu k_\nu + 36m^2 \eta_{\mu\nu} k_\mu k_\nu) \right. \\ & \left. -20(d+2)m^2 k_\mu k_\mu k_\mu k_\nu k_\nu k_\nu + 48k^4 m^4 \eta_{\mu\nu} (2\eta_{\mu\nu} \eta_{\mu\nu} + 3\eta_{\mu\mu} \eta_{\nu\nu}))\right] \tag{11.30}$$

 $\tilde{J}_{\mu\mu\mu\nu\nu\nu}(k)$  is not conserved

$$k^{\mu}\tilde{J}_{\mu\mu\nu\nu}(k) = \frac{3 \cdot 2^{-d-1} i e^{-\frac{1}{2}i\pi d} \pi^{-d/2} (-m^2)^{\frac{d}{2}} \Gamma\left(-\frac{d}{2}-1\right)}{m^2} \\ \times \left[24m^2 \eta_{\mu\nu} k_{\mu} ((d+2)k_{\nu}k_{\nu}+4m^2 \eta_{\nu\nu}) + 4m^2 k_{\nu} (\eta_{\mu\mu} ((d+2)k_{\nu}k_{\nu}+12m^2 \eta_{\nu\nu}) + 24m^2 \eta_{\mu\nu} \eta_{\mu\nu}) + (d+2)k_{\mu}k_{\mu}k_{\nu} (d k_{\nu}k_{\nu}+12m^2 \eta_{\nu\nu})\right]$$
(11.31)

#### 11.2.2 Concluding remark

In this section we have produced two-point correlator formulas for spin 1,2 and 3 in any dimension. In all the cases where it is possible to make a comparison between the results obtained in this section and the previous ones (spin 1 in d=3,4,5,6) the results coincide.<sup>8</sup> However they do *only* if we subtract the infinite and finite IR terms from the effective action. In other words this confirms that only the difference between the UV and the IR can have a physical meaning.

<sup>&</sup>lt;sup>8</sup>In making the comparison one should not forget to correct the results of section 11.1.2 and 11.1.3 for the Tr(1) factor as explained before eq. (11.4), i.e. by dividing the d = 2, 3 results by 2, multiplying the d = 6, 7 ones by 2, etc.

#### 12 Conclusion

We have seen a large number of examples that the one-loop effective action of a free massive model coupled to external sources contains complete information about the (classical) equations of motion of the sources. In this paper we have considered only the two-point functions and so the relevant information involve the linearized equations of motion. Moreover we have considered only completely symmetric bosonic external sources. Within these limitations we have produced overwhelming evidence that our previous statement is correct. We have considered both a free scalar model and a free fermion model in different dimensions, and shown that in all cases the two-point functions of conserved currents are built out of the differential operators which define the linearized (Fronsdal) equations of motion of the fields that couple to the currents. There is no doubt that such free field theories know about the dynamics of the fields that can couple to them (via a conserved current).<sup>9</sup>

At this stage a specification is in order. Our intent in this paper was to show the universal appearance of non-local Fronsdal (as well as Maxwell and EH) linearized eom in the one-loop effective actions of a free scalar and boson field coupled to external currents, while postponing other subtler questions to future research. In particular we did not tackle the problem raised by [47, 48], concerning the form of the Fronsdal equation that guarantees the right propagator for the relevant higher spin field. In order to do that one must first of all specify to what equations one refers to, for we have seen that in the IR and UV limits of the OLEA's the conserved structures very often are different, and different from the various tomographic sections, although for spin s they are all characterized by the presence of the leading (4.1) term (the scalar model is in this sense a particular, though less interesting, case, because the conserved structures are always the same for given spin). This part of our research is work in progress, see [49].

The results of this paper opens a new research territory. Beside the just mentioned problem, we would like to know whether the above results extend to other external sources, fermionic fields as well as not completely symmetric fields. The next question is interaction, which requires analyzing three-point functions. In this context interactions have been considered for the simplest cases (spin 1 and 2) in 3d in [1]. From three-point functions one expects to find information about the consistency of the (field or fields) interaction with the source field symmetry. For instance, for spin 1 with gauge symmetry, for spin 2 with diffeomorphisms. For higher spins we do not know, in general, neither the interaction nor the full form of the symmetry transformations. But knowing the three-point functions may be the key to constructing both. There are anyhow some exceptions to our ignorance in this field (higher spin theories in 3d, or Vasiliev's higher spin theory in  $AdS_4$ , or string field theory). One can hopefully use this knowledge to test the approach suggested here.

If our conjecture is correct, that is if the analysis of three-point correlators in theories coupled to external sources confirms their consistency with the dynamics of the latter, as we believe, an obvious question comes next: what does this mean? The correspondence between one free field theory and higher (or low) spin theories is not a type of duality we

<sup>&</sup>lt;sup>9</sup>The limitation to free field theories does not seem to be essential provided the currents are conserved, but of course explicit calculations are far more complicated in the case of interacting theories.

are familiar with, like AdS/CFT. First of all it concerns models in the same dimension. Secondly, from one free theory we retrieve knowledge about (infinite) many theories. So the correspondence would be one to (infinite) many. And this is clearly not satisfactory. The results of this paper points rather toward the possibility of a correspondence between theories with infinite many fields. If, say, a starting free (or interacting) theory knows about the dynamics of other fields, why shouldn't the latter be included with the initial one in a unique theory? Arguing this way one is led to a (for the time being, generic) concept of *involutive theory*: a theory is *involutive* if it includes all the fields it is able to couple with (in the OLEA) while preserving a fundamental symmetry.

A good playground to test this concept could be string field theory (SFT). Such a theory is formulated in terms of a basic string field  $\Phi$ . The latter, in the field theory regime, is a superposition of Fock space states, each with a coefficient given by a suitable ordinary spacetime field. Restricting ourselves to bosonic SFT, the action formulated by Witten is well known, and is given by the formula below with  $\Psi$  replaced by  $\Phi$ . Analyzing it in the spirit of this paper amounts to studying the theory

$$S = \frac{1}{2g_o} \int \left( \Phi * Q\Phi + \frac{2}{3} \Phi * \Phi * \Psi \right)$$
(12.1)

where the first piece is the free SFT and the second is the simplest interaction with the source term ( $\Psi$  is the source string field). The first piece is invariant under the BRST transformation  $\delta \Phi = Q\Lambda$ . The second term carries the invariance under  $\delta \Psi = Q\Lambda$  provided that  $\Phi$  is on shell, i.e.  $Q\Phi = 0$ . This mimics what we have done previously for simple field theory models.

#### Acknowledgments

We would like to thank Ivica Smolić for his collaboration in the early stage of this research. One of us (L.B.) would like to thank Dario Francia and Mikhail Vasiliev for useful discussions. The research has been supported by the Croatian Science Foundation under the project No. 8946 and by the University of Rijeka under the research support No. 13.12.1.4.05.

# A Proof that a conserved structure can be written in terms of products of $\pi$ alone

By induction in steps of 2. In the lowest case (spin 1), the most general Lorentz covariant (dimensionless) conserved even structure can be written in terms of  $\eta_{\mu\nu}$  and  $\frac{k_{\mu}k_{\nu}}{k^2}$ . Imposing conservation the result is  $\sim \eta_{\mu\nu} - \frac{k_{\mu}k_{\nu}}{k^2} = \pi_{\mu\nu}$ . In the same way one can prove the property for the case s = 2. Now we suppose that the proposition is true for s. So it is true for the combination  $T^{(s)}(k \cdot n_1^s \cdot n_2^s) = \tilde{E}^{(s)}(k) = \sum_{l=0}^{[s/2]} a_l \tilde{A}_l^{(s)}$  (see above), meaning that  $k^{\mu} \frac{\partial}{\partial n_1^{\mu}} T^{(s)} = 0$ . In order to construct  $T^{(s+1)}$  we can multiply  $T^{(s)}$  by  $(n_1n_2)$  or  $\frac{(n_1k)(n_2k)}{k^2}$  or multiplying  $T^{(s-1)}$  by  $(n_1n_1)(n_2n_2), \frac{(n_1k)(n_1k)}{k^2}(n_2n_2), (n_1n_1)\frac{(n_2k)(n_2k)}{k^2}$  or by

 $\frac{(n_1k)(n_1k)}{k^2}\frac{(n_2k)(n_2k)}{k^2}$ , because the construction is in steps or 2. So we can have only

$$T^{(s+1)} = a_1(n_1n_2)T^{(s)} + a_2(n_1n_1)(n_2n_2)T^{(s-1)} + b_1\frac{(n_1k)(n_2k)}{k^2}T^{(s)} + b_2\frac{(n_1k)(n_1k)}{k^2}(n_2n_2)T^{(s-1)} + b_3(n_1n_1)\frac{(n_2k)(n_2k)}{k^2}T^{(s-1)} + b_4\frac{(n_1k)(n_1k)}{k^2}\frac{(n_2k)(n_2k)}{k^2}T^{(s-1)}$$
(A.1)

Now applying  $k^{\mu} \frac{\partial}{\partial n_1^{\mu}}$  to this expression we find that conservation requires  $a_1 = -b_1, a_2 = -b_2 = -b_3 = b_4$ . So that (A.1) becomes

$$T^{(s+1)} = a(n_1, \pi^{(k)} \cdot n_2) T^{(s)} + b(n_1, \pi^{(k)} \cdot n_1)(n_2, \pi^{(k)} \cdot n_2) T^{(s-1)}$$
(A.2)

with arbitrary a and b.

### **B** Massless models

In this appendix we consider the massless case both for the scalar and the fermion models, i.e. we set m = 0 in their action, and derive the relevant two-point functions for several tensorial currents in any dimension. These results are based on the scalar integral (5.8).

The results we report below have to be compared with the results obtained in the section 6-9, precisely with the  $\mathcal{O}_{\text{UV}}(m^0) - \mathcal{O}_{\text{IR}}(m^0) - \mathcal{O}_{\text{IR}}(\log(m))$  terms therein. It can be easily checked that for odd d the results coincide exactly, as far as the even parity part of the correlators is concerned, while the massless model approach is unable to reproduce the odd parity part (at least perturbatively). In the even dimensional case the results of the two approaches do not, in general, coincide. Only the terms proportional to  $\log(-k^2)$  are the same in the two approaches. These differences are due to the lack of IR regularization in the massless model approach.

For conciseness in this appendix we use a simplified notation, taken from the literature on higher spin fields: the same repeated subscript, say  $\mu \dots \mu$  repeated s times, stand for s completely symmetrized labels.

In this appendix we also construct traceless two-point correlators. As a matter of fact, in this paper we are only marginally interested in zero trace currents. But the tracelessness condition may be relevant for further developments.

#### B.1 Massless fermion model

Let us start with spin 1 case for massless fermions. The two point correlator is

$$\tilde{J}_{\mu\nu}(k) = -\int \frac{d^d p}{(2\pi)^d} \operatorname{Tr}\left(\gamma_{\nu} \frac{1}{\not{p}} \gamma_{\mu} \frac{1}{\not{p} - \not{k}}\right) = -\operatorname{Tr}(1) \int \frac{d^d p}{(2\pi)^d} \frac{p_{\nu}(p-k)_{\mu} + p_{\mu}(p-k)_{\nu} - p \cdot (p-k)\eta_{\mu\nu}}{p^2(p-k)^2}$$
(B.1)

where  $\text{Tr}(1) = 2^{\lfloor \frac{d}{2} \rfloor}$ . Using Davydychev methods, the two point correlator for spin s=1 and any dimension *d* reads

$$\tilde{J}_{\mu\nu} = -\frac{2^{2-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}-1} \left(d-2\right)}{\left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+1}{2}\right)} \pi_{\mu\nu}$$
(B.2)

where  $\pi_{\mu\nu}$  is the usual projector.

In a similar way for spin 2 in the massless case we get

$$\tilde{J}_{\mu\mu\nu\nu} = \frac{2^{1-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{a}{2}} \left(d-1\right)}{\left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+3}{2}\right)} \left(\pi_{\mu\nu}^2 - \frac{1}{d-1}\pi_{\mu\mu}\pi_{\nu\nu}\right)$$
(B.3)

The two-point correlator in s = 2 case is traceless for any d.

For spin 3 we obtain

$$\tilde{J}_{\mu\mu\mu\nu\nu\nu} = -\frac{2^{-2-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2} - \frac{d}{2}} \left(k^2\right)^{\frac{d}{2} + 1}}{9\left(-1 + e^{i\pi d}\right) \Gamma\left(\frac{d+5}{2}\right)} \pi_{\mu\nu}}{\times \left(32 \, d \, \pi_{\mu\nu}^2 + \left(d((d-6)d+3) - 54\right) \pi_{\mu\mu} \pi_{\nu\nu}\right)}$$
(B.4)

We can check if this expression is traceless

$$\eta^{\mu\mu}\tilde{J}_{\mu\mu\mu\nu\nu\nu} = -\frac{2^{-1-2d+\left\lfloor\frac{d}{2}\right\rfloor}(d-3)^2 \left(d^2+d-6\right) \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}+1}}{9 \left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+5}{2}\right)} \pi_{\mu\nu}\pi_{\nu\nu} \tag{B.5}$$

For dimension d = 3, the two-point correlator  $J_{\mu\mu\mu\nu\nu\nu}$  is traceless. To obtain tracelessness for other dimensions we can add local counterterms to the two point function

$$\tilde{J}_{\mu\mu\mu\nu\nu\nu} = -\frac{2^{-2-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}+1}}{9\left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+5}{2}\right)} \pi_{\mu\nu}} \times \left( (32\,d+a_1)\,\pi^2_{\mu\nu} + \left(d((d-6)d+3) - 54 + a_2)\pi_{\mu\mu}\pi_{\nu\nu}\right) \right)$$
(B.6)

The counterterms are local only for even d. In this case it is easy to verify that for the full expression to be traceless we must have

$$a_1 = \frac{1}{3} \left( -a_2(d+1) - \left( d^2 + d - 6 \right) (d-3)^2 \right)$$

The traceless 2pt function is therefore

$$\tilde{J}_{\mu\mu\mu\nu\nu\nu\nu} = -\frac{2^{-2-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}+1} (1+d)(-54+3d-6d^2+d^3+a_2)}{9 \left(-1+e^{i\pi d}\right) \Gamma \left(\frac{d+5}{2}\right)} \times \pi_{\mu\nu} \left(\pi_{\mu\nu}^2 - \frac{3}{(d+1)} \pi_{\mu\mu} \pi_{\nu\nu}\right)$$
(B.7)

For spin s = 4 the two-point correlator is

$$\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu\nu} = \frac{2^{-3-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}+2}}{9\left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+7}{2}\right)} \left((24(d+1))\pi^4_{\mu\nu} + \left((d-7)(d(d+2)+9)\right)\pi^2_{\mu\nu}\pi_{\mu\mu}\pi_{\nu\nu} + \left((9-(d-4)d)\right)\pi^2_{\mu\mu}\pi^2_{\nu\nu}\right)$$
(B.8)

Again, unless d = 3, the expression for  $\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu\nu}$  is not traceless For even d > 3 we can add local counterterms

$$\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu\nu} = \frac{2^{-3-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}+2}}{9\left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+7}{2}\right)} \left((24(d+1)+a_1)\pi^4_{\mu\nu} + ((d-7)(d(d+2)+9)+a_2)\pi^2_{\mu\nu}\pi_{\mu\mu}\pi_{\nu\nu} + ((9-(d-4)d)+a_3)\pi^2_{\mu\mu}\pi^2_{\nu\nu}\right)$$
(B.9)

and we obtain a traceless 2pt correlator for

$$a_1 = -(1/3)(1+d)((-3+d)^2(5+d) - a_3(3+d)), \qquad a_2 = (d-3)^2(d+5) - 2a_3(d+1)$$

The traceless two point correlator is now

$$\tilde{J}_{\mu\mu\mu\mu\mu\nu\nu\nu\nu} = -\frac{2^{-3-2d+\left\lfloor \frac{d}{2} \right\rfloor} \pi^{\frac{3}{2}-\frac{d}{2}} \left(-a_3 + d^2 - 4d - 9\right) \left(k^2\right)^{\frac{d}{2}+2} (d+1)(d+3)}{27 \left(-1 + e^{i\pi d}\right) \Gamma \left(\frac{d+7}{2}\right)} \times \left(\pi_{\mu\nu}^4 - \frac{6}{(d+3)} \pi_{\mu\mu} \pi_{\mu\nu}^2 \pi_{\nu\nu} + \frac{3}{(d+1)(d+3)} \pi_{\mu\mu}^2 \pi_{\nu\nu}^2\right)$$
(B.10)

In general, conserved and traceless 2pt function for spin s in any (even) dimension is proportional to

$$\tilde{J}_{\mu_1\dots\mu_s\nu_1\dots\nu_s} \sim \sum_{l=0}^{\left\lfloor \frac{s}{2} \right\rfloor} \frac{(-1)^l}{2^{2l}l!} \frac{s!}{(s-2l)!} \frac{\Gamma\left(s + \frac{d-3}{2} - l\right)}{\Gamma\left(s + \frac{d-3}{2}\right)} \pi_{\mu\nu}^{s-2l} (\pi_{\mu\mu}\pi_{\nu\nu})^l \tag{B.11}$$

We can write the sum as

$$\tilde{J}_{\mu_1\dots\mu_s\nu_1\dots\nu_s} \sim {}_2F_1\left(\frac{1-s}{2}, -\frac{s}{2}; \frac{5-d-2s}{2}; \frac{\pi_{\mu\mu}\pi_{\nu\nu}}{\pi_{\mu\nu}^2}\right)\pi_{\mu\nu}^s \tag{B.12}$$

## B.2 Massless scalar model

Let us compute the two point correlator for a scalar in d dimensions for spin s = 1. Using Davydychev methods we get:

$$\tilde{J}_{\mu\nu} = -\frac{2^{3-2d} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}-1}}{\left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+1}{2}\right)} \pi_{\mu\nu}$$
(B.13)

The expression for  $\tilde{J}_{\mu\nu}$  is traceless and conserved.

For spin 2 we get

$$\tilde{J}_{\mu\mu\nu\nu} = \frac{2^{2-2d} \pi^{\frac{3}{2} - \frac{d}{2}} (k^2)^{\frac{d}{2}}}{(-1 + e^{i\pi d}) \Gamma(\frac{d+3}{2})} \left(2\pi^2_{\mu\nu} + \pi_{\mu\mu}\pi_{\nu\nu}\right)$$
(B.14)

which is conserved, but not traceless:

$$\eta^{\mu\mu}\tilde{J}_{\mu\mu\nu\nu} = \frac{2^{3-2d} \pi^{\frac{3}{2}-\frac{d}{2}} (k^2)^{\frac{d}{2}}}{(-1+e^{i\pi d}) \Gamma(\frac{d+3}{2})} (d+1) \pi_{\nu\nu}$$
(B.15)

Let us consider the counterterms  $\tilde{J}_{\mu\mu\nu\nu}$ 

$$\tilde{J}_{\mu\mu\nu\nu} = \frac{2^{2-2d} \pi^{\frac{3}{2} - \frac{d}{2}} (k^2)^{\frac{d}{2}}}{(-1 + e^{i\pi d}) \Gamma(\frac{d+3}{2})} \left( (2 + a_1) \pi^2_{\mu\nu} + (1 + a_2) \pi_{\mu\mu} \pi_{\nu\nu} \right)$$
(B.16)

which is local for even d. Adding it to  $\tilde{J}_{\mu\mu\nu\nu}$ , the trace becomes

$$\eta^{\mu\mu}\tilde{J}_{\mu\mu\nu\nu} = \frac{2^{3-2d} \pi^{\frac{3}{2}-\frac{d}{2}} (k^2)^{\frac{d}{2}}}{(-1+e^{i\pi d}) \Gamma(\frac{d+3}{2})} (d+1+a_1+a_2(d-1)) \pi^2_{\nu\nu}$$
(B.17)

If we choose

$$a_1 = -(1+d) - a_2(d-1)$$

we get a traceless  $\tilde{J}_{\mu\mu\nu\nu}$ 

$$\tilde{J}_{\mu\mu\nu\nu} = -\frac{2^{2-2d} \pi^{\frac{3}{2}-\frac{d}{2}} (k^2)^{\frac{d}{2}}}{(-1+e^{i\pi d}) \Gamma(\frac{d+3}{2})} (1+a_2)(d-1) \left(\pi_{\mu\nu}^2 - \frac{1}{d-1}\pi_{\mu\mu}\pi_{\nu\nu}\right)$$
(B.18)

This is possible only for even d.

For spin 3 we have

$$\tilde{J}_{\mu\mu\mu\nu\nu\nu} = -\frac{3 \cdot 2^{1-2d} \pi^{\frac{3}{2} - \frac{d}{2}} (k^2)^{\frac{d}{2} + 1}}{(-1 + e^{i\pi d}) \Gamma(\frac{d+5}{2})} \pi_{\mu\nu} \left(2\pi_{\mu\nu}^2 + 3\pi_{\mu\mu}\pi_{\nu\nu}\right)$$
(B.19)

This expression is transverse but not traceless

$$\eta^{\mu\mu}\tilde{J}_{\mu\mu\mu\nu\nu\nu} = -\frac{9 \cdot 2^{2-2d} \pi^{\frac{3}{2} - \frac{d}{2}} (k^2)^{\frac{d}{2} + 1}}{(-1 + e^{i\pi d}) \Gamma(\frac{d+5}{2})} (3+d) \pi_{\mu\nu}\pi_{\nu\nu}$$
(B.20)

In even d we can add local counterterms and obtain

$$\tilde{J}_{\mu\mu\mu\nu\nu\nu} = -\frac{3 \cdot 2^{1-2d} \pi^{\frac{3}{2} - \frac{d}{2}} \left(k^2\right)^{\frac{d}{2} + 1}}{\left(-1 + e^{i\pi d}\right) \Gamma\left(\frac{d+5}{2}\right)} \pi_{\mu\nu} \left((2 + a_1)\pi_{\mu\nu}^2 + (3 + a_2)\pi_{\mu\mu}\pi_{\nu\nu}\right)$$
(B.21)

To make  $\tilde{J}_{\mu\mu\mu\nu\nu\nu}$  traceless we must have

$$a_1 = -\frac{1}{3}(3(3+d) + a_2(1+d))$$

A traceless  $\tilde{J}_{\mu\mu\mu\nu\nu\nu}$  can now be written as

$$\tilde{J}_{\mu\mu\mu\nu\nu\nu\nu} = \frac{2^{1-2d} \pi^{\frac{3}{2}-\frac{d}{2}} \left(k^2\right)^{\frac{d}{2}+1}}{\left(-1+e^{i\pi d}\right) \Gamma\left(\frac{d+5}{2}\right)} (3+a_2)(d+1)\pi_{\mu\nu} \left(\pi_{\mu\nu}^2 - \frac{3}{d+1}\pi_{\mu\mu}\pi_{\nu\nu}\right) \tag{B.22}$$

For spin 4 the two-point correlator is

$$\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu} = -\frac{3 \cdot 2^{-2d} \pi^{\frac{3}{2} - \frac{d}{2}} (k^2)^{\frac{d}{2} + 2}}{(-1 + e^{i\pi d}) \Gamma(\frac{d+7}{2})} \left(8\pi_{\mu\nu}^4 + 24\pi_{\mu\nu}^2 \pi_{\mu\mu}\pi_{\nu\nu} + 3\pi_{\mu\mu}^2 \pi_{\nu\nu}^2\right)$$
(B.23)

The expression for  $\tilde{J}_{\mu_1...\mu_4\nu_1...\nu_4}$  is not traceless

$$\eta^{\mu\mu}\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu\nu} = -\frac{9\cdot 2^{2-2d}\pi^{\frac{3}{2}-\frac{d}{2}}(k^2)^{\frac{d}{2}+2}}{(-1+e^{i\pi d})\Gamma(\frac{d+7}{2})}(5+d)\pi_{\nu\nu}\left(4\pi_{\mu\nu}^2+\pi_{\mu\mu}\pi_{\nu\nu}\right)$$
(B.24)

Again in even d we can add local counterterms

$$\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu} = -\frac{3 \cdot 2^{-2d} \pi^{\frac{3}{2} - \frac{d}{2}} (k^2)^{\frac{a}{2} + 2}}{(-1 + e^{i\pi d}) \Gamma(\frac{d+7}{2})} \\ \cdot \left( (8 + a_1) \pi^4_{\mu\nu} + (24 + a_2) \pi^2_{\mu\nu} \pi_{\mu\mu} \pi_{\nu\nu} + (3 + a_3) \pi^2_{\mu\mu} \pi^2_{\nu\nu} \right)$$
(B.25)

The tracelessness condition is now

$$a_1 = -5 + a_3 + \frac{4}{3}d(3 + a_3) + \frac{1}{3}d^2(3 + a_3), \qquad a_2 = -2(15 + 3d + (1 + d)a_3)$$

so that a traceless  $\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu}$  can be written

$$\tilde{J}_{\mu\mu\mu\mu\nu\nu\nu\nu} = -\frac{2^{-2d} \pi^{\frac{3}{2} - \frac{d}{2}} (k^2)^{\frac{d}{2} + 2}}{(-1 + e^{i\pi d}) \Gamma(\frac{d+7}{2})} (3 + a_3)(d+1)(d+3) \cdot \left(\pi^4_{\mu\nu} - \frac{6}{(d+3)} \pi^2_{\mu\nu} \pi_{\mu\mu} \pi_{\nu\nu} + \frac{3}{(d+1)(d+3)} \pi^2_{\mu\mu} \pi^2_{\nu\nu}\right)$$
(B.26)

As in the case of fermions, the transverse and traceless 2pt function for spin s in even dimensions is proportional to

$$\tilde{J}_{\mu_1\dots\mu_s\nu_1\dots\nu_s} \sim \sum_{l=0}^{\lfloor \frac{s}{2} \rfloor} \frac{(-1)^l}{2^{2l}l!} \frac{s!}{(s-2l)!} \frac{\Gamma\left(s + \frac{d-3}{2} - l\right)}{\Gamma\left(s + \frac{d-3}{2}\right)} \pi_{\mu\nu}^{s-2l} (\pi_{\mu\mu}\pi_{\nu\nu})^l = {}_2F_1\left(\frac{1-s}{2}, -\frac{s}{2}; \frac{5-d-2s}{2}; \frac{\pi_{\mu\mu}\pi_{\nu\nu}}{\pi_{\mu\nu}^2}\right) \pi_{\mu\nu}^s$$
(B.27)

# C 4d full amplitudes

In view of the importance of the 4d case, we give in the following complete explicit formulas in terms of elementary functions of the two point correlator for spin 1,2,3 in the fermionic model.

# C.1 Spin 1

$$\frac{ik^2}{\pi^2}(n_1 \cdot \pi \cdot n_2) \left( -\frac{m^2}{3k^2} + \sqrt{4m^2 - k^2} \left(\frac{m^2}{3k^3} + \frac{1}{6k}\right) \csc^{-1}\left(\frac{2m}{k}\right) + \frac{1}{12}L_1 - \frac{1}{18} \right)$$

normalized with the understanding that in UV:

$$\sqrt{4m^2 - k^2} \to -ik + i\frac{2m}{k^2} + \dots \tag{C.1}$$

$$\csc^{-1}\left(\frac{2m}{k}\right) \to \frac{1}{2}i\log\left(-\frac{k^2}{m^2}\right) - \frac{im^2}{k^2} + \dots$$
(C.2)

Here,

$$L_n = \frac{2}{\varepsilon} + \log\left(\frac{m^2}{4\pi}\right) + \gamma - \sum_{k=1}^n \frac{1}{k}$$
(C.3)

C.2 Spin 2

$$\begin{aligned} \frac{i}{\pi^2} \frac{m^4}{4k^4} \left(k^2 \left(k \cdot n_2\right)^2 (n_1 \cdot \pi \cdot n_1) + 2k^2 (k \cdot n_1) (k \cdot n_2) (n_1 \cdot \pi \cdot n_2) + k^2 \left(k \cdot n_1\right)^2 (n_2 \cdot \pi \cdot n_2) \right. \\ \left. + 2 \left(k \cdot n_1\right)^2 \left(k \cdot n_2\right)^2\right) L_2 \\ \left. + \frac{i}{\pi^2} k^4 \left(n_1 \cdot \pi \cdot n_2\right)^2 \left(-\frac{8m^4}{15k^4} + \frac{7m^2}{360k^2} + \left(\frac{m^4}{4k^4} + \frac{m^2}{12k^2} - \frac{1}{40}\right) L_2 \right. \\ \left. + \sqrt{4m^2 - k^2} \left(\frac{8m^4}{15k^5} + \frac{m^2}{15k^3} - \frac{1}{20k}\right) \csc^{-1} \left(\frac{2m}{k}\right) + \frac{9}{400}\right) \\ \left. + \frac{i}{\pi^2} k^4 (n_1 \cdot \pi \cdot n_1) (n_2 \cdot \pi \cdot n_2) \left(-\frac{4m^4}{15k^4} + \frac{41m^2}{360k^2} + \left(\frac{m^4}{4k^4} - \frac{m^2}{12k^2} + \frac{1}{120}\right) L_2 \right. \\ \left. + \sqrt{4m^2 - k^2} \left(\frac{4m^4}{15k^5} - \frac{2m^2}{15k^3} + \frac{1}{60k}\right) \csc^{-1} \left(\frac{2m}{k}\right) - \frac{47}{3600}\right) \end{aligned}$$

# C.3 Spin 3

$$\begin{split} i\left(n_{1}\cdot\pi\cdot n_{2}\right)^{3}k^{6}\frac{1}{\pi^{2}}\left(-\frac{128m^{6}}{315k^{6}}+\frac{32m^{4}}{315k^{4}}+\frac{218m^{2}}{14175k^{2}}+\left(\frac{8m^{6}}{27k^{6}}-\frac{4m^{2}}{135k^{2}}+\frac{4}{945}\right)L_{3}\right.\\ &\left.+\sqrt{4m^{2}-k^{2}}\left(\frac{128m^{6}}{315k^{7}}-\frac{64m^{4}}{945k^{5}}-\frac{8m^{2}}{189k^{3}}+\frac{8}{945k}\right)\csc^{-1}\left(\frac{2m}{k}\right)-\frac{428}{99225}\right)\\ &\left.+i\left(k\cdot n_{1}\right)^{2}\left(k\cdot n_{2}\right)^{2}(n_{1}\cdot\pi\cdot n_{2})k^{2}\frac{1}{\pi^{2}}\left(\frac{m^{4}}{18k^{4}}+\left(\frac{37m^{6}}{27k^{6}}+\frac{m^{4}}{6k^{4}}\right)L_{3}\right)\right.\\ &\left.+i\left(k\cdot n_{1}\right)^{3}\left(k\cdot n_{2}\right)^{3}\frac{1}{\pi^{2}}\left(\frac{7m^{4}}{72k^{4}}+\left(\frac{7m^{6}}{9k^{6}}+\frac{7m^{4}}{24k^{4}}\right)L_{3}\right)\right.\\ &\left.+i\left(k\cdot n_{1}\right)(k\cdot n_{2})\left(n_{1}\cdot\pi\cdot n_{2}\right)^{2}\frac{8}{9\pi^{2}k^{2}}L_{3}\right.\\ &\left.+i\left(n_{1}\cdot\pi\cdot n_{1}\right)\left(n_{1}\cdot\pi\cdot n_{2}\right)\left(n_{2}\cdot\pi\cdot n_{2}\right)\frac{k^{6}}{\pi^{2}}\left(-\frac{64m^{6}}{105k^{6}}+\frac{257m^{4}}{1512k^{4}}-\frac{1877m^{2}}{56700k^{2}}\right)\right.\\ &\left.+\left(\frac{13m^{6}}{27k^{6}}-\frac{m^{4}}{8k^{4}}+\frac{4m^{2}}{135k^{2}}-\frac{37}{15120}\right)L_{3}\right.\\ &\left.+\sqrt{4m^{2}-k^{2}}\left(\frac{64m^{6}}{105k^{7}}-\frac{152m^{4}}{945k^{5}}+\frac{187m^{2}}{3780k^{3}}-\frac{37}{7560k}\right)\csc^{-1}\left(\frac{2m}{k}\right)+\frac{12433}{3175200}\right)\right.\\ &\left.+i\left(k^{2}(k\cdot n_{2})(n_{2}\cdot\pi\cdot n_{2})\left(k\cdot n_{1}\right)^{3}+k^{2}\left(k\cdot n_{2}\right)^{3}(n_{1}\cdot\pi\cdot n_{1})(k\cdot n_{1})\right)\right.\\ &\left.\cdot\frac{1}{\pi^{2}}\left(\frac{7m^{4}}{144k^{4}}+\left(\frac{13m^{6}}{27k^{6}}+\frac{7m^{4}}{48k^{4}}\right)L_{3}\right)\right.\\ &\left.+13im^{6}(k\cdot n_{1})(k\cdot n_{2})(n_{1}\cdot\pi\cdot n_{2})k^{4}+(k\cdot n_{1})^{2}(n_{1}\cdot\pi\cdot n_{2})(n_{2}\cdot\pi\cdot n_{2})k^{4}\right)\right.\\ &\left.\cdot\frac{1}{\pi^{2}}\left(\frac{m^{4}}{144k^{4}}+\left(\frac{13m^{6}}{27k^{6}}+\frac{m^{4}}{48k^{4}}\right)L_{3}\right)\right.\\ \end{array}$$

**Open Access.** This article is distributed under the terms of the Creative Commons Attribution License (CC-BY 4.0), which permits any use, distribution and reproduction in any medium, provided the original author(s) and source are credited.

#### References

- L. Bonora, M. Cvitan, P. Dominis Prester, B. Lima de Souza and I. Smolić, Massive fermion model in 3d and higher spin currents, JHEP 05 (2016) 072 [arXiv:1602.07178] [INSPIRE].
- [2] E.E. Boos and A.I. Davydychev, A method of evaluating massive Feynman integrals, Theor. Math. Phys. 89 (1991) 1052 [Teor. Mat. Fiz. 89 (1991) 56] [INSPIRE].
- [3] A.I. Davydychev, A simple formula for reducing Feynman diagrams to scalar integrals, Phys. Lett. B 263 (1991) 107 [INSPIRE].
- [4] A.I. Davydychev, Recursive algorithm of evaluating vertex type Feynman integrals, J. Phys. A 25 (1992) 5587 [INSPIRE].
- [5] C. Fronsdal, Massless fields with integer spin, Phys. Rev. D 18 (1978) 3624 [INSPIRE].
- [6] J. Fang and C. Fronsdal, Massless fields with half integral spin, Phys. Rev. D 18 (1978) 3630 [INSPIRE].
- [7] T. Curtright, Massless field supermultiplets with arbitrary spin, Phys. Lett. B 85 (1979) 219
   [INSPIRE].
- [8] L.P.S. Singh and C.R. Hagen, Lagrangian formulation for arbitrary spin 1. The boson case, Phys. Rev. D 9 (1974) 898 [INSPIRE].
- [9] L.P.S. Singh and C.R. Hagen, Lagrangian formulation for arbitrary spin 2. The fermion case, Phys. Rev. D 9 (1974) 910 [INSPIRE].
- [10] D. Francia and A. Sagnotti, On the geometry of higher spin gauge fields, Class. Quant. Grav. 20 (2003) S473 [Comment. Phys. Math. Soc. Sci. Fenn. 166 (2004) 165] [PoS(JHW2003)005] [hep-th/0212185] [INSPIRE].
- [11] D. Francia and A. Sagnotti, Free geometric equations for higher spins, Phys. Lett. B 543 (2002) 303 [hep-th/0207002] [INSPIRE].
- [12] J.M. Maldacena and G.L. Pimentel, On graviton non-Gaussianities during inflation, JHEP 09 (2011) 045 [arXiv:1104.2846] [INSPIRE].
- [13] X.O. Camanho, J.D. Edelstein, J. Maldacena and A. Zhiboedov, *Causality constraints on corrections to the graviton three-point coupling*, *JHEP* 02 (2016) 020 [arXiv:1407.5597]
   [INSPIRE].
- [14] M.A. Vasiliev, Consistent equation for interacting gauge fields of all spins in (3+1)-dimensions, Phys. Lett. B 243 (1990) 378 [INSPIRE].
- [15] M.A. Vasiliev, Properties of equations of motion of interacting gauge fields of all spins in (3+1)-dimensions, Class. Quant. Grav. 8 (1991) 1387 [INSPIRE].
- [16] M.A. Vasiliev, Algebraic aspects of the higher spin problem, Phys. Lett. B 257 (1991) 111 [INSPIRE].
- [17] M.A. Vasiliev, More on equations of motion for interacting massless fields of all spins in (3+1)-dimensions, Phys. Lett. B 285 (1992) 225 [INSPIRE].
- [18] Higher-spin gauge theories, in Proceedings of the First Solvay Workshop, held in Brussels on May 12–14 2004, R. Argurio, G. Barnich, G. Bonelli and M. Grigoriev eds., Int. Solvay Institutes, Belgium (2006).
- [19] D. Sorokin, Introduction to the classical theory of higher spins, AIP Conf. Proc. 767 (2005) 172 [hep-th/0405069] [INSPIRE].

- [20] D. Francia and A. Sagnotti, Higher-spin geometry and string theory, J. Phys. Conf. Ser. 33 (2006) 57 [hep-th/0601199] [INSPIRE].
- [21] A. Fotopoulos and M. Tsulaia, Gauge invariant lagrangians for free and interacting higher spin fields. A review of the BRST formulation, Int. J. Mod. Phys. A 24 (2009) 1
   [arXiv:0805.1346] [INSPIRE].
- [22] C. Iazeolla, On the algebraic structure of higher-spin field equations and new exact solutions, arXiv:0807.0406 [INSPIRE].
- [23] A. Campoleoni, Metric-like lagrangian formulations for higher-spin fields of mixed symmetry, Riv. Nuovo Cim. 33 (2010) 123 [arXiv:0910.3155] [INSPIRE].
- [24] A. Sagnotti, *Higher spins and current exchanges*, PoS(CORFU2011)106 [arXiv:1002.3388] [INSPIRE].
- [25] D. Francia, Low-spin models for higher-spin lagrangians, Prog. Theor. Phys. Suppl. 188 (2011) 94 [arXiv:1103.0683] [INSPIRE].
- [26] A.D. Sakharov, Vacuum quantum fluctuations in curved space and the theory of gravitation, Sov. Phys. Dokl. 12 (1968) 1040 [Dokl. Akad. Nauk Ser. Fiz. 177 (1967) 70] [Sov. Phys. Usp. 34 (1991) 394] [Gen. Rel. Grav. 32 (2000) 365] [INSPIRE].
- [27] J. Maldacena and A. Zhiboedov, Constraining conformal field theories with a higher spin symmetry, J. Phys. A 46 (2013) 214011 [arXiv:1112.1016] [INSPIRE].
- [28] J. Maldacena and A. Zhiboedov, Constraining conformal field theories with a slightly broken higher spin symmetry, Class. Quant. Grav. 30 (2013) 104003 [arXiv:1204.3882] [INSPIRE].
- [29] A. Campoleoni, Higher spins in D = 2 + 1, Subnucl. Ser. 49 (2013) 385 [arXiv:1110.5841] [INSPIRE].
- [30] C. Closset, T.T. Dumitrescu, G. Festuccia, Z. Komargodski and N. Seiberg, Comments on Chern-Simons contact terms in three dimensions, JHEP 09 (2012) 091 [arXiv:1206.5218] [INSPIRE].
- [31] S. Giombi, S. Prakash and X. Yin, A note on CFT correlators in three dimensions, JHEP 07 (2013) 105 [arXiv:1104.4317] [INSPIRE].
- [32] S. Giombi, S. Minwalla, S. Prakash, S.P. Trivedi, S.R. Wadia and X. Yin, Chern-Simons theory with vector fermion matter, Eur. Phys. J. C 72 (2012) 2112 [arXiv:1110.4386] [INSPIRE].
- [33] C.N. Pope and P.K. Townsend, Conformal higher spin in (2+1)-dimensions, Phys. Lett. B 225 (1989) 245 [INSPIRE].
- [34] M.P. Blencowe, A consistent interacting massless higher spin field theory in D = (2 + 1), Class. Quant. Grav. 6 (1989) 443 [INSPIRE].
- [35] B. de Wit and D.Z. Freedman, Systematics of higher spin gauge fields, Phys. Rev. D 21 (1980) 358 [INSPIRE].
- [36] T. Damour and S. Deser, 'Geometry' of spin 3 gauge theories, Ann. Inst. H. Poincaré Phys. Theor. 47 (1987) 277 [INSPIRE].
- [37] R. Jackiw and S. Templeton, How superrenormalizable interactions cure their infrared divergences, Phys. Rev. D 23 (1981) 2291 [INSPIRE].
- [38] I. Vuorio, Parity violation and the effective gravitational action in three-dimensions, Phys. Lett. B 175 (1986) 176 [INSPIRE].

- [39] T. Appelquist, M.J. Bowick, E. Cohler and L.C.R. Wijewardhana, *Chiral symmetry breaking in* (2+1)-dimensions, *Phys. Rev. Lett.* 55 (1985) 1715 [INSPIRE].
- [40] T.W. Appelquist, M.J. Bowick, D. Karabali and L.C.R. Wijewardhana, Spontaneous chiral symmetry breaking in three-dimensional QED, Phys. Rev. D 33 (1986) 3704 [INSPIRE].
- [41] T. Appelquist, M.J. Bowick, D. Karabali and L.C.R. Wijewardhana, Spontaneous breaking of parity in (2+1)-dimensional QED, Phys. Rev. D 33 (1986) 3774 [INSPIRE].
- [42] K.S. Babu, A.K. Das and P. Panigrahi, Derivative expansion and the induced Chern-Simons term at finite temperature in (2+1)-dimensions, Phys. Rev. D 36 (1987) 3725 [INSPIRE].
- [43] R. Delbourgo and A.B. Waites, Induced parity violation in odd dimensions, Austral. J. Phys. 47 (1994) 465 [hep-th/9404164] [INSPIRE].
- [44] G.V. Dunne, Aspects of Chern-Simons theory, hep-th/9902115 [INSPIRE].
- [45] F.S. Gama, J.R. Nascimento and A. Yu. Petrov, Derivative expansion and the induced Chern-Simons term in N = 1, D = 3 superspace, Phys. Rev. D 93 (2016) 045015
   [arXiv:1511.05471] [INSPIRE].
- [46] X. Bekaert, E. Joung and J. Mourad, Effective action in a higher-spin background, JHEP 02 (2011) 048 [arXiv:1012.2103] [INSPIRE].
- [47] D. Francia, J. Mourad and A. Sagnotti, Current exchanges and unconstrained higher spins, Nucl. Phys. B 773 (2007) 203 [hep-th/0701163] [INSPIRE].
- [48] D. Francia, On the relation between local and geometric lagrangians for higher spins, J. Phys. Conf. Ser. 222 (2010) 012002 [arXiv:1001.3854] [INSPIRE].
- [49] L. Bonora, M. Cvitan, P. Dominis Prester, S. Giaccari, B. Lima de Souza and T. Štemberga, Massive Dirac field in 3D and induced equations for higher spin fields, in Group31 Proceedings, to appear, Rio de Janeiro Brazil (2016).
- [50] L. Bonora and B. Lima de Souza, *Pure contact term correlators in CFT*, arXiv:1511.06635 [INSPIRE].
- [51] L. Bonora, S. Giaccari and B. Lima de Souza, Trace anomalies in chiral theories revisited, JHEP 07 (2014) 117 [arXiv:1403.2606] [INSPIRE].
- [52] L. Bonora, S. Giaccari and B.L.D. Souza, *Revisiting trace anomalies in chiral theories*, Springer Proc. Math. Stat. 111 (2014) 3 [INSPIRE].
- [53] L. Bonora, A.D. Pereira and B. Lima de Souza, Regularization of energy-momentum tensor correlators and parity-odd terms, JHEP 06 (2015) 024 [arXiv:1503.03326] [INSPIRE].
- [54] B.L. de Souza, DavyFeyn: a mathematica package for automated analytical Feynman diagram computations webpage, https://github.com/blimasouza/DavyFeyn.